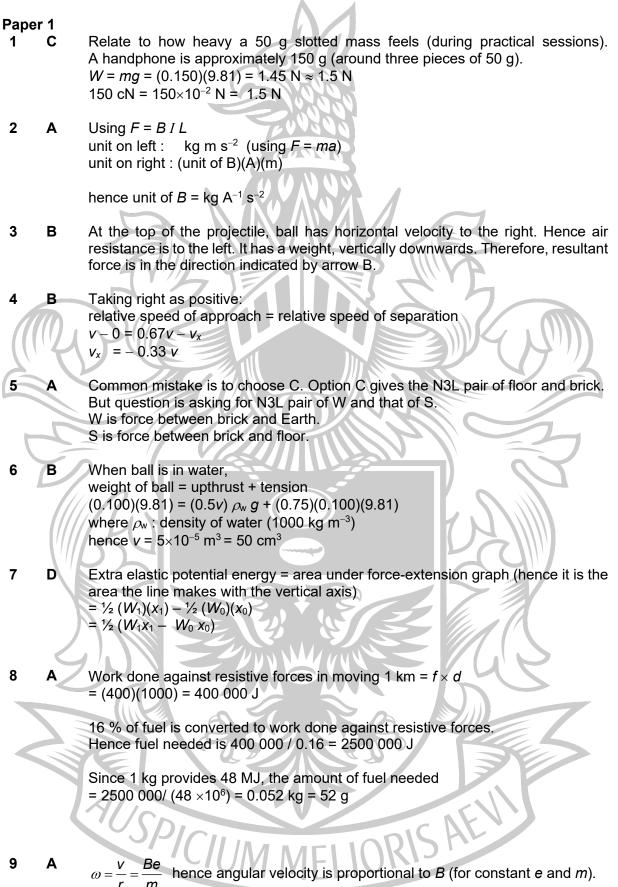
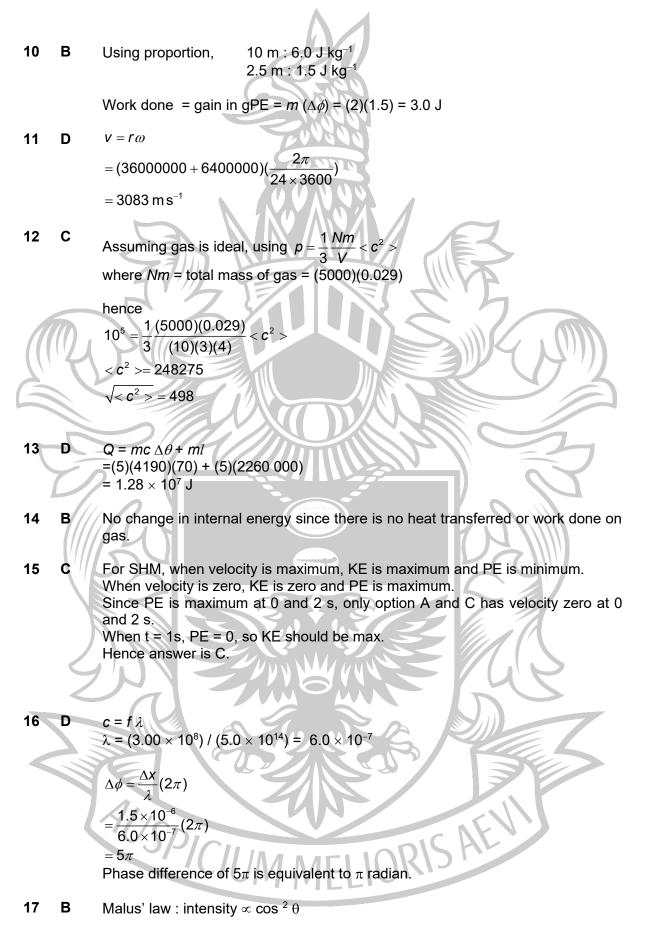
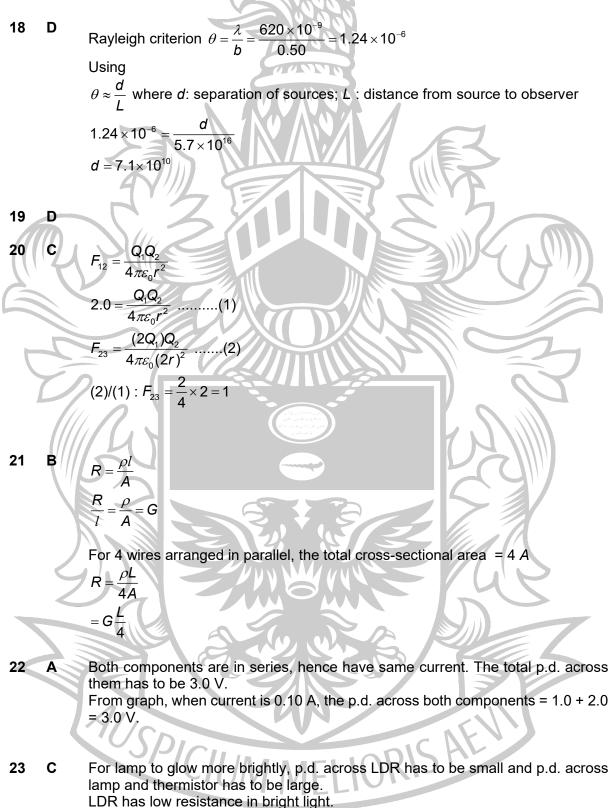
2021 A-Level H2 Physics Suggested Solutions





Initially, when middle filter is 45° to X and Y, the intensity of light emerging from Y is non-zero.

When middle filter is 90° to X and Y, the intensity of light emerging from Y is zero (since $\cos^2 90^\circ = 0$).



Thermistor has high resistance in low temperature.

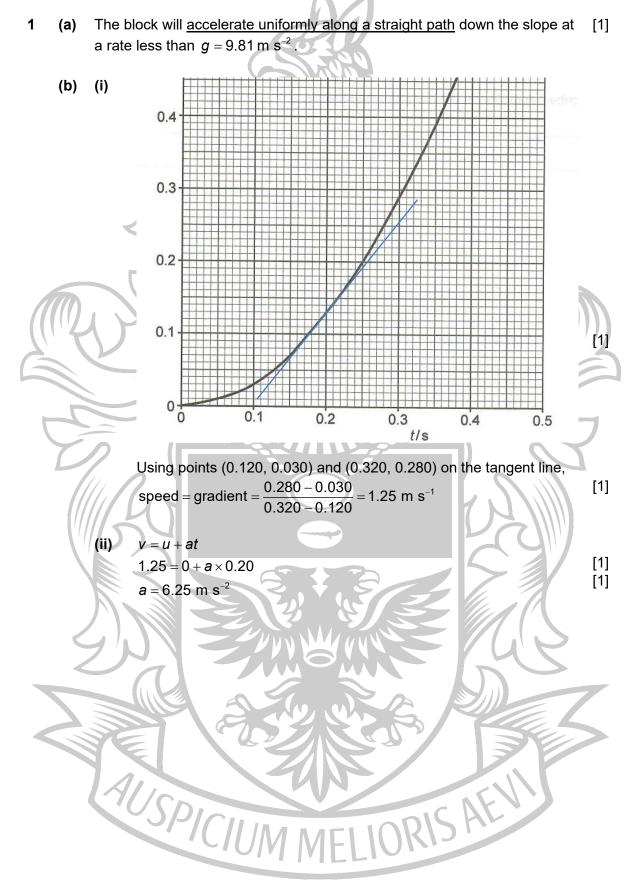
24 A Using right-hand grip rule for the circular coil, the B-field is pointing into paper. Using FLHR, the force on short wire is upwards.

25 D
26 A

$$\langle E \rangle = \frac{\Delta \Phi}{t} = \frac{\Delta (NBA)}{t}$$

 $= \frac{(3000)(1.8)(x0.01^2)}{0.06} = 0$
 $= 28 \vee$
27 B
 $\frac{V_s}{V_s} = \frac{16}{1.88} = 8.5$
 $\frac{V_s}{V_s} = \frac{1}{1.8}$
Hence $I_s = 0.32 \pm 8.5 = 0.038 \text{ A} = 38 \text{ mA}$
28 B
Longer wavelength corresponds to transition between smaller AE.
Hence 590 nm corresponds to transition from $n = 3$ to 2.
 $\Delta E_{3-1} = \Delta E_{3-2} + \Delta E_{2-1}$
 $\frac{hC}{h} = \frac{hC}{hC}}$
 $\lambda = 252$
29 D
intensity $= \frac{P}{A} = \frac{Nhf}{tA}$
30 B
Energy released = (measure - masspocatels) C^2
 $= (136.90709 - 136.90583 - 5.49 \times 10^{-4})(1.66 \times 10^{-27})(3.00 \times 10^{9})^2$
 $= 1.1 \times 10^{-33}$





2 (a) From graph, the momentum of the ball as it makes contact with the surface (at 0.53 s) is 3.20 N s. [1]

$$E_{k} = \frac{p^{2}}{2m} = \frac{3.20^{2}}{2 \times 0.62}$$
= 8.26 J
[1]

(b) From graph, the ball is in contact with the surface from t = 0.53 s to 0.68 s, during which its momentum changes from 3.20 N s to -1.80 N s.

The magnitude of the average force on the ball is equal to the magnitude of the total change in momentum over the total time interval. (This is basically from Newton's 2nd Law.)

$$\langle F \rangle = \left| \frac{\Delta p}{\Delta t} \right| = \left| \frac{-1.80 - 3.20}{0.68 - 0.53} \right| = 33.3 \text{ N}$$

But note that **this average force represents only the net force on the ball** – not the force the surface acts on the ball! That would be the normal contact force, which is what we are tasked to find.

When in contact with the surface, two forces act on the ball – the normal contact force *N* upwards, and the weight of the ball *mg* downwards. (Note that the magnitude of *N* can vary with time.) However, since $\langle F \rangle$ on the ball

is upwards, the average normal contact force $\langle N \rangle$ must be larger than the [1] weight of the ball. [1]

$$\langle F \rangle = \langle N \rangle - mg$$

 $\langle N \rangle = \langle F \rangle + mg = 33.3 + 0.62(9.81)$
= 39.4 N

Note: The final answer has to be positive, since it is the magnitude.

(c) KE after the first rebound =
$$\frac{1.80^2}{2 \times 0.62}$$
 = 2.61

fraction of KE retained after every rebound = $\frac{2.61}{8.26} = 0.316$

[1]

[1]

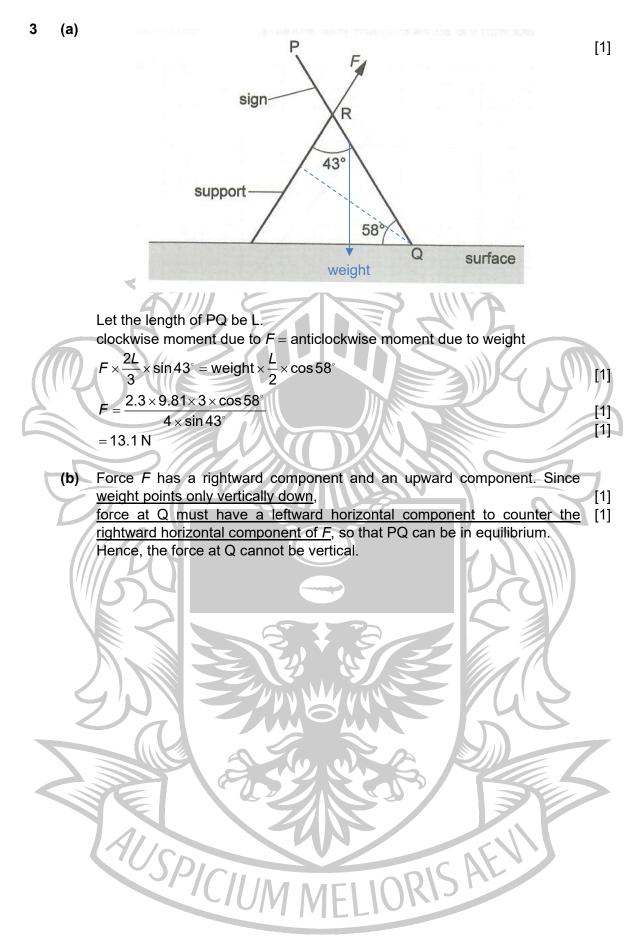
[1]

[1]

[1]

After *n*-th rebound, the fraction of the initial KE retained is 0.316^n $0.316^n = 5.0\% = 0.050$ (solved by taking logarithm on both sides) n = 2.60 = 3 (round up, integer number of times of rebounding)

Note: Since the fractional loss of KE is constant for every rebound, the fraction of KE retained is also constant, because (fractional loss) + (fraction retained) = 1.



4 (a) Consider a small test mass with mass *m*, located a distance *r* from *M*.

Newton's Law of gravitation: gravitational force $F = \frac{GMm}{r^2}$

Gravitational field strength g is defined as the gravitational force experienced per unit mass.

$$g = \frac{\text{gravitational force}}{m} = \frac{1}{m} \frac{GMm}{r^2} = \frac{GM}{r^2}$$
[1]

(b) Since the centripetal force is provided by the gravitational force, the centripetal acceleration a_c is equal to g.

$$a_{\rm c} = g, \ \omega^2 r = \frac{GM}{r^2} \Rightarrow r = \left(\frac{GM}{\omega^2}\right)^{\frac{1}{3}}$$
 [1]

angular speed of the satellite, $\omega = \frac{2\pi}{110 \times 60} = 9.52 \times 10^{-4} \text{ rad s}^{-1}$ [1]

Hence,

g

$$=\frac{GM}{r^{2}}=\frac{GM}{\left(\frac{GM}{\omega^{2}}\right)^{\frac{2}{3}}}=\left(GM\omega^{4}\right)^{\frac{1}{3}}=(6.67\times10^{-11}\times6.0\times10^{24}\times(9.52\times10^{-4})^{4})^{\frac{1}{3}}$$
[1]

= 6.90 m s

(c) Using the satellite from (b) as a comparison, <u>Earth's angular velocity ω is [1] $\frac{2\pi}{T} = \frac{2\pi}{24(60)(60)} = 7.27 \times 10^{-5} \text{ rad s}^{-1}, \text{ which is an order of magnitude}$ <u>smaller than the satellite's $\omega = 9.52 \times 10^{-4} \text{ rad s}^{-1}$. Since $a_c = r\omega^2$, the a_c on</u></u>

Earth's surface must be about two orders of magnitude smaller than the satellite's centripetal acceleration $a_c = g = 6.90 \text{ m s}^{-2}$, due to the ω^2 term and the smaller orbital radius *r* for an object on Earth.

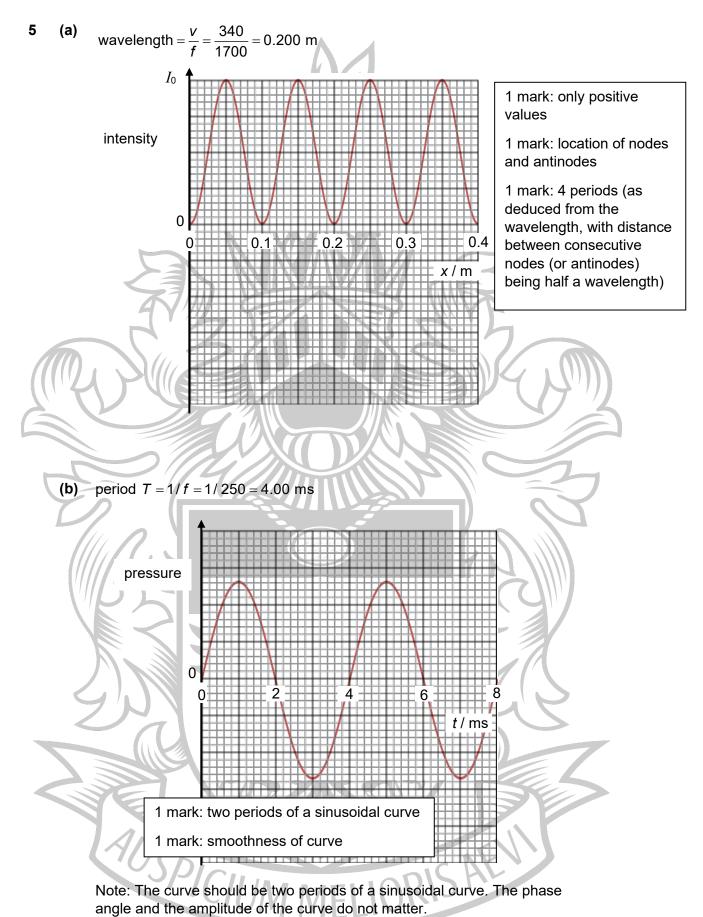
Therefore, the <u>centripetal acceleration a_c of an object moving with the</u> <u>surface of Earth must also be about two orders of magnitude smaller than</u> <u>the gravitational acceleration g</u>, which is approximately 9.81 m s⁻².

Since $g = a_c + a_{\text{freefall}}$, with a_c being negligible, g and a_{freefall} are approximately [1] equal. They are exactly equal at the South and North poles, where the radius of circular motion (hence a_c) is zero.

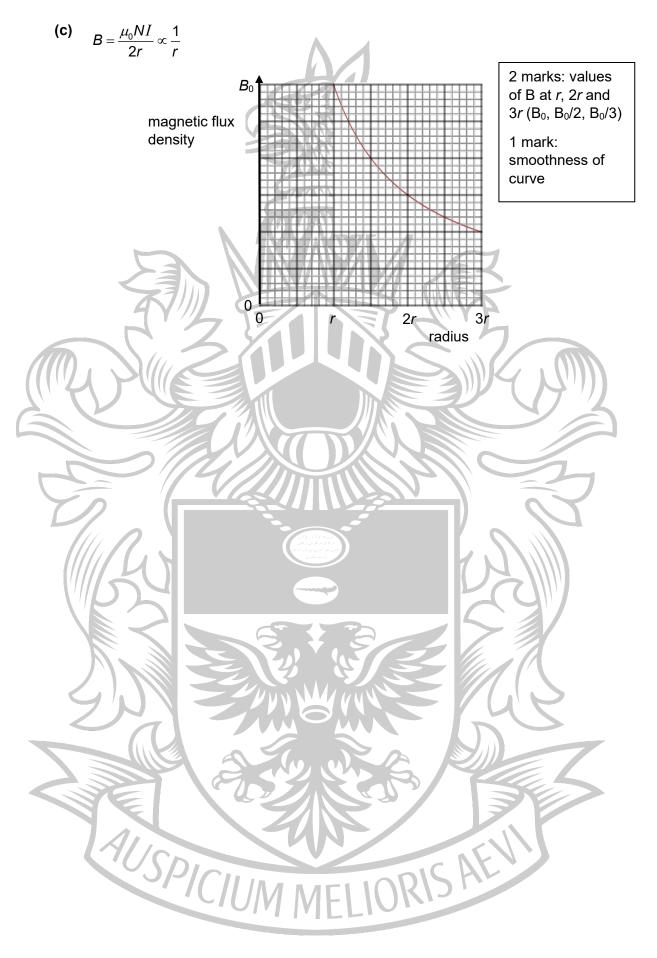
(Note: Earth's radius r is approximately 6370 km, and its angular velocity ω

$$a_c = \frac{2\pi}{T} = \frac{2\pi}{24(60)(60)} = 7.27 \times 10^{-5} \text{ rad s}^{-1}$$
. Hence, $a_c = r\omega^2 = 0.034 \text{ m s}^{-2}$

which is almost 300 times smaller than $g = 9.81 \text{ m s}^{-2}$. Hence, $a_c \ll g$.)



5



6 (a)

(i) Potential difference across R = potential difference across lamp

$$= 0.30 \times 5.0$$

= 1.5 V [1]

Potential difference across R = potential difference across lamp = 0.30×5.0 = 1.5 V

[1]

[1]

[1]

[1]

Potential difference across the 3.0 Ω resistor = 6.0 - 1.5 = 4.5 V

Current flowing through the 3.0 Ω resistor $=\frac{4.5}{3.0}=1.5$ A Hence, current flowing through *R* is 1.5 - 0.30 = 1.2 A. Resistance of $R = \frac{1.5}{1.2} = 1.25 \approx 1.3 \Omega$ [1]

Alternatively:

Potential difference across the 3.0 Ω resistor = 6.0 – 1.5 = 4.5 V 11

Let the effective resistance of R and the lamp be R_{eff} . Then

$$\frac{R_{eff}}{3} = \frac{1.5}{4.5} \implies R_{eff} = 1.0 \Omega$$
Hence,
$$\frac{5 \times R}{5 + R} = 1.0$$

$$R = 1.3 \Omega$$
(1]

(ii) The ratio of the power dissipated by the 3.0 Ω , *R* and the lamp is: $1.5^2 \times 3.0: 1.2^2 \times 1.25: 0.3^2 \times 5.0$ [1] = 6.75: 1.8: .45

= 15 : 4 : 1

(iii)

Hence, the energy transferred in the lamp is

$$\frac{1}{15+4+1}$$
 × 120 = 6.0 J

At the beginning, the <u>temperature of the filament is at its lowest (room</u> [1] <u>temperature), which results in a large current as the resistance is low</u> as well.

As the current flows through the lamp, the <u>temperature of the filament</u> <u>lamp will increase</u>, which will cause the <u>resistance of the filament to</u> <u>increase</u>, resulting in a decrease in current.

Hence, the current is the greatest when the current is first switched on as the temperature and resistance of the filament is the lowest at the start.

(b) Since X and Y are connected in series, the current flowing through is the [1] same.



7 Some of the alpha particles were observed to be scattered at large angles. (a) [1] This shows that they encountered a positive charge that is concentrated [1] into a very small volume known as the nucleus so that the coulomb repulsion is very large.

> Since the α particles are back scattered, the positive charge they collided [1] with must also have a much larger mass, otherwise it would be the positive charges that were pushed away rather than the alpha particles scattered backward.

> Hence, the mass of the atom is also largely concentrated within the nucleus. [1]

> > [1]

[1]

11

[1]

[1]

[1]

[1]

For minimum separation, all the kinetic energy of the α particle is converted (b) into electric potential energy. We are essentially finding the distance of closest approach.

Loss in kinetic energy = Gain in electric potential energy $5.59 \times 1.6 \times 10^{-13} - 0 = \frac{(2e)(79e)}{2}$ 0

 $4\pi\varepsilon_r$

 $2 \times 79 \times (1.6 \times 10^{-19})$ $4\pi (8.85 \times 10^{-12})r$

 $r = 4.07 \times 10^{-14} \text{ m}$

²²²₈₆Rn (C) $^{218}_{84}$ Po + $^{4}_{2}$ He

> Let the binding energy per nucleon of polonium be E. Polonium has 222 - 4 = 218 nucleons.

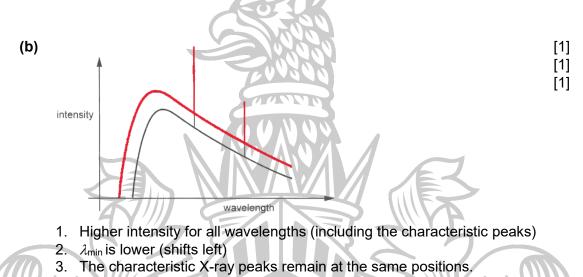
Total binding energy of products = (4)(7.08) + (218)ETotal binding energy of reactant (radon) = (222)(7.69)

 E_{κ} released = $BE_{products} - BE_{reactants}$ 6.62 = (4)(7.08) + (218)E - (222)(7.69)E = 7.73 MeV

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8 It is evacuated to prevent electrons from the electron beam colliding with (a) [1] the gas molecules in the tube, which would reduce the efficiency of X-ray production as scattered electrons may either lose too much energy during the collisions to produce X-rays, or they may not even hit the metal target.



Note: the peak of maximum intensity should shift left to a shorter wavelength.

(i) Number of electrons used in one image =

(C)

$$= \frac{0.12 \times 1.1}{1.6 \times 10^{-19}}$$

= 8.25 × 10¹⁷
Thermal energy produced = 0.99 × 8.25 × 10¹⁷ × (65 × 1.6 × 10⁻¹⁶)
= 8494.2 ≈ 8500 J

[1]

[1]

[1]

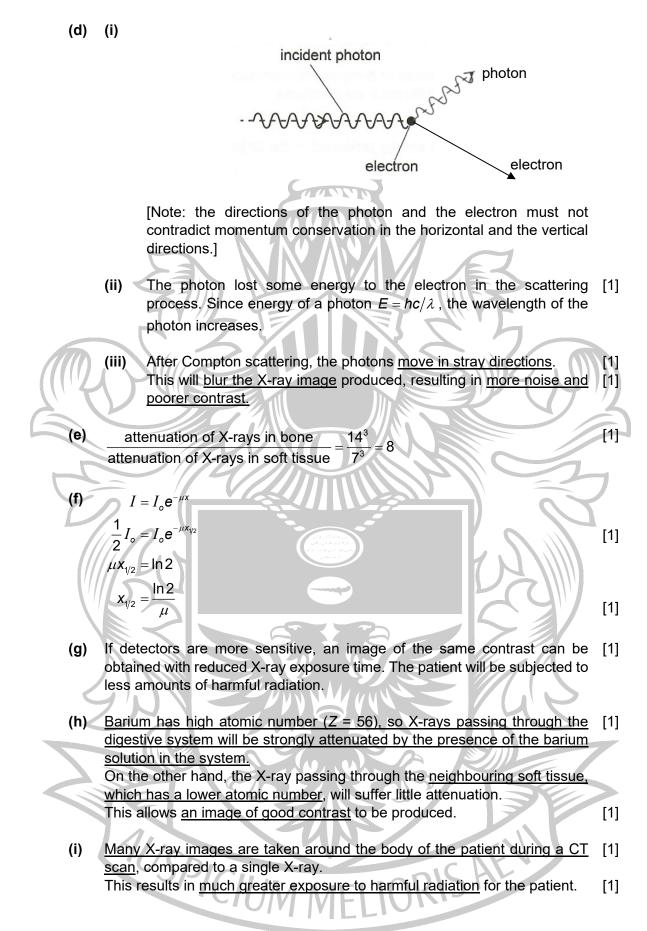
[1]

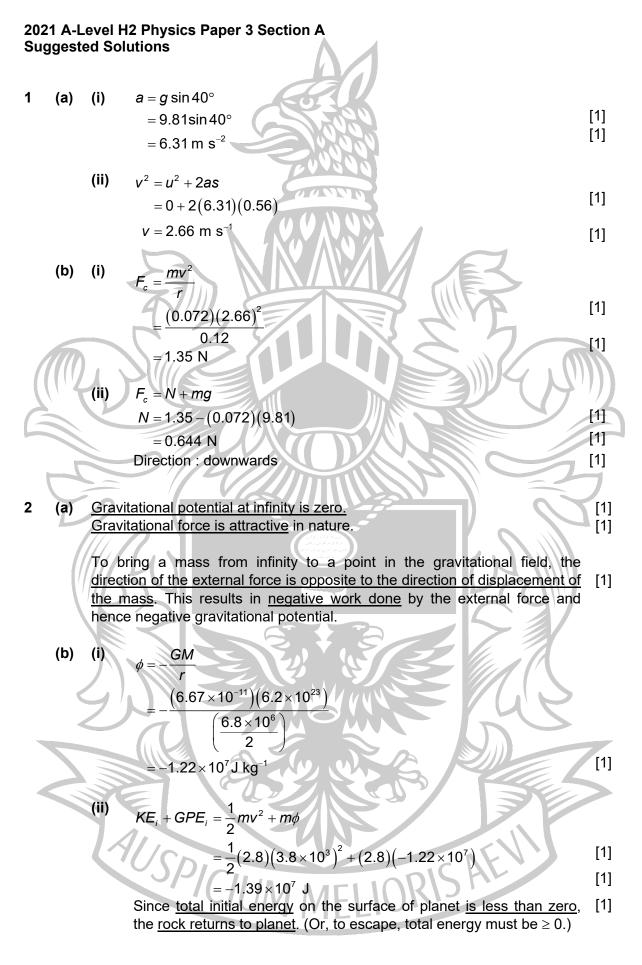
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(ii) Thermal energy = $mc\Delta\theta$ $8500 = 0.012 \times 130 \times \Delta\theta$ $\Delta \theta = 5445 \approx 5400$ °C

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(iii) This is to prevent any part of the target from being exposed to the [1] electron beam for a prolonged period (e.g. 1.1 s), which may cause the target to be heated beyond its melting point and start to melt.





3 The internal energy of an ideal gas is the <u>sum of the kinetic energies</u> of the [1] (a) molecules due to their random motion. [1]

> Note: An ideal gas has no potential energy as the molecules do not experience intermolecular forces except during collisions.

Since initial and final gas has the same number of moles of particles, (b) (i) $(\mathbf{n}\mathbf{V})$ $(\mathbf{n}\mathbf{V})$

$$\left(\frac{p_{V}}{RT}\right)_{1} = \left(\frac{p_{V}}{RT}\right)_{2}$$

$$\frac{(1.0 \times 10^{5})(3.2 \times 10^{-3})}{12 + 273.15} = \frac{(1.0 \times 10^{5})(3.6 \times 10^{-3})}{T_{2}}$$
[1]

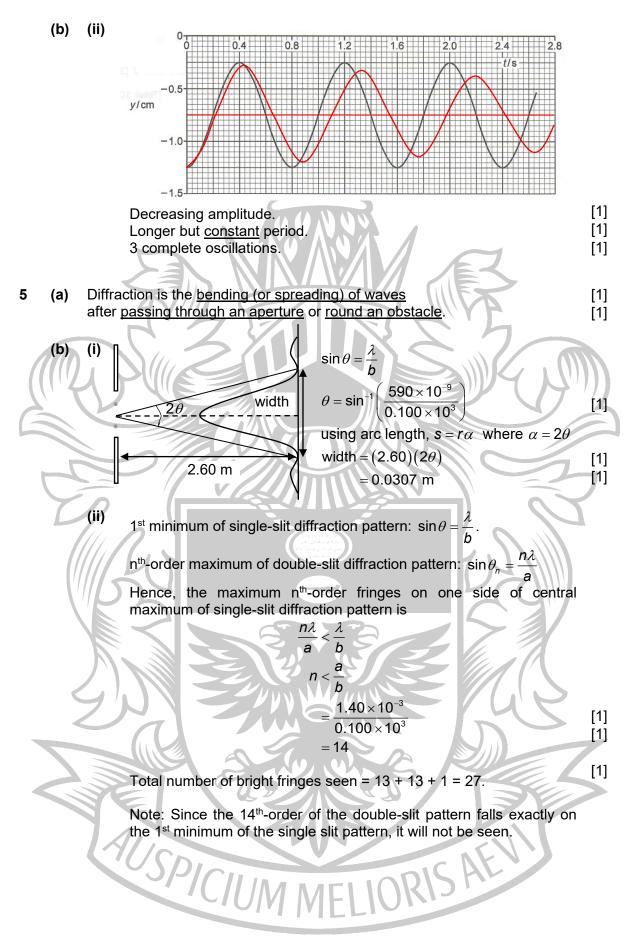
$$T_2 = 320.79 \text{ K}$$

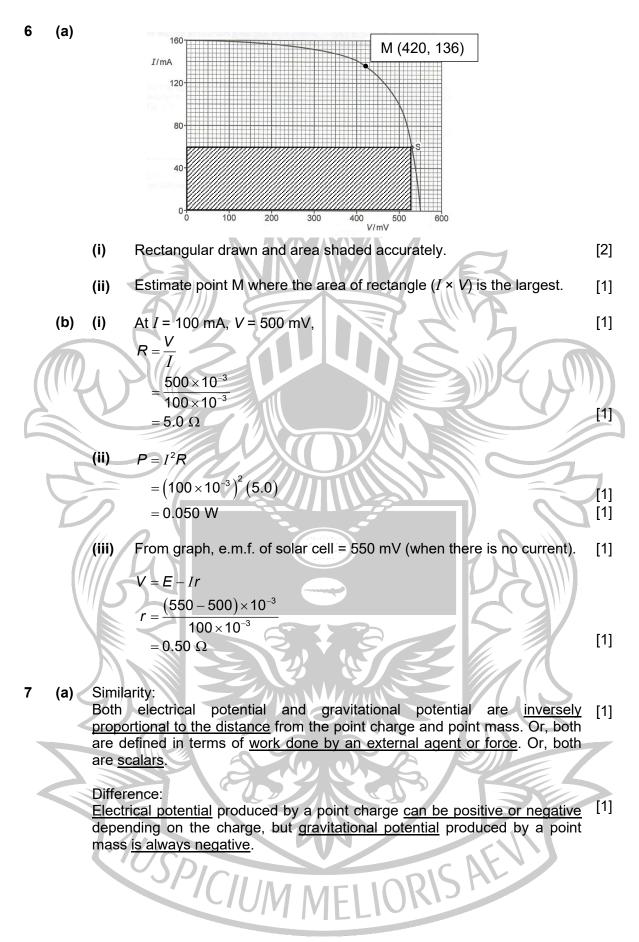
= 47.6 °C [1]

/::.\

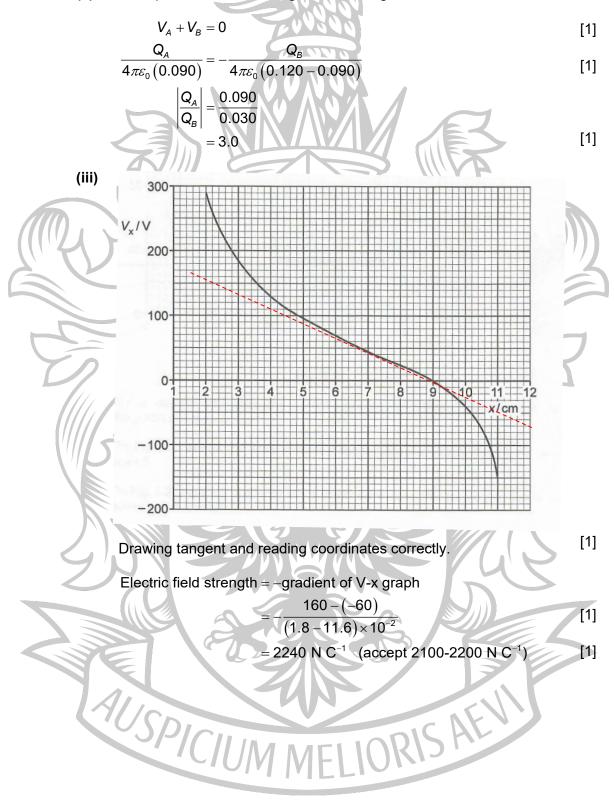
(ii) Work done by gas against atmosphere =
$$p_{AV}$$

= $(1.0 \times 10^5)(3.6 - 3.2) \times 10^3$ [1]
= 40 J
(i)
 $AU = W + Q = -40 + 101 = 61 J$
(i)
 $N = \frac{pV}{kT} = \frac{(1.0 \times 10^5)(3.2 \times 10^{-3})}{(1.38 \times 10^{-23})(12 + 273.15)} = 8.132 \times 10^{-22}$ [1]
 $\frac{AU}{N} = \frac{61}{8.132 \times 10^{-22}}$ [1]
 $\frac{AU}{N} = \frac{61}{8.132 \times 10^{-22}}$ [1]
 $\frac{AU}{N} = \frac{-0.250 - (-1.250)}{2} = 0.500 \text{ cm}$ [1]
(ii)
 $w = \frac{2\pi}{7}$ [1]
(iii)
 $w_0 = wx_0$
 $= (7.85)(0.50)$
 $= 3.93 \text{ cm s}^{-1}$ [1]
(1)
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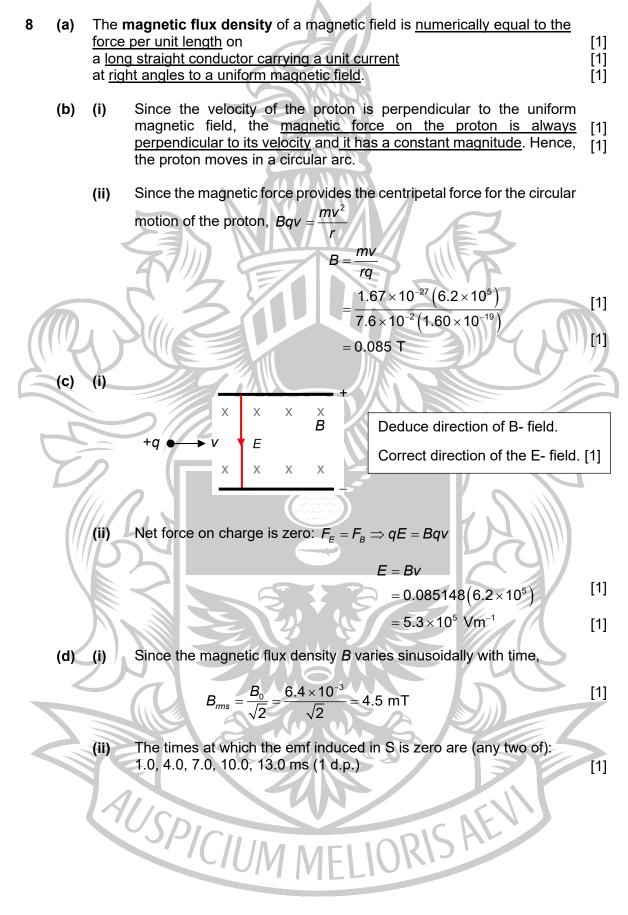




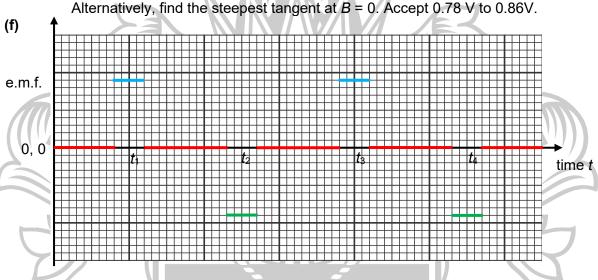
- (b) (i) The potential between the two charges is the addition of the individual potential produced by each charge. <u>Since potential is positive nearer to A and negative nearer to B</u>, it shows that charge A is positive and charge B is negative. Hence the charges have <u>opposite signs</u>.
 - (ii) Total potential due to charge A and charge B is zero at x = 9.0 cm.



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(e)
$$B_0 = 6.4 \times 10^{-3} \text{ T}$$
 [1]
 $T = 6.0 \times 10^{-3} \text{ s}$ [1]
 $\varepsilon = -\frac{d\Phi}{dt} = -NB_0 A \frac{d}{dt} \sin(\omega t + \theta) = -NB_0 A \omega \cos(\omega t + \theta),$
 $\varepsilon_{\text{max}} = NB_0 A \omega$
 $= (270) (6.4 \times 10^{-3}) \left(\pi \left(\frac{2.4 \times 10^{-2}}{2} \right)^2 \right) \left(\frac{2\pi}{6.0 \times 10^{-3}} \right)$ [1]
 $= 0.82 \text{V}$ [1]



Emf is zero during times when flux is constant – shown by the red lines [1]

Emf is **positive** and **constant** during times when flux is decreasing uniformly with time – shown by the **blue** lines [1]

Emf is **negative** and **constant** during te times when flux is increasing uniformly with time – shown by the **green** lines [1]

Magnitude and duration of blue and green lines are the same [1]

9 (a)

(i)

- Photoelectrons are emitted from the surface of a metal only if [1] the frequency of the light incident on the metal is higher than a certain minimum or threshold value.
- The maximum kinetic energy of the photoelectrons is dependent [1] only on the frequency, but not on the intensity of the incident light.

(ii) <u>Emission spectrum is produced when an atom de-excite by releasing</u> [1] <u>energy as a photon.</u>

<u>The energy of the emitted photon is given by difference in energies</u> [1] <u>between two energy-levels</u> in the atom $\left(\frac{hc}{\lambda} = E_{higher} - E_{lower}\right)$.

Since <u>only certain wavelengths are present</u> ($\lambda = \frac{hc}{E_{higher} - E_{lower}}$), it [1] means that discrete energy levels exist in the atom.

(b) (i)
$$E = \frac{hc}{\lambda}$$

 $= \frac{6.63 \times 10^{-34} \times 3.00 \times 10^8}{340 \times 10^{-9}} = 5.85 \times 10^{-19} \text{ J}$ [1]
 $= \frac{5.85 \times 10^{-19}}{1.60 \times 10^{-19}} \text{ eV}$ [1]
 $= 3.66 \text{ eV}$ [1]

<u>De-excitation between the two lowest levels gives rise to photons of</u> [1] <u>energy -3.4 - (-13.6) = 10.2 eV, which is higher than 3.66 eV of a photon of purple light.</u>

This means that the <u>wavelengths of photons due to transitions to the</u> [1] -13.6 eV level are shorter than 340 nm, which are <u>non-visible</u> since they correspond to uv light. Thus, a visible line spectrum cannot result from transitions to the -13.6 eV energy level.

[1]

[1]

[1] [1]

Decay constant λ = probability of decay per unit time. Hence, probability of decay in time Δt is given by $\lambda \Delta t$.

$$\Delta t = \frac{\ln 2}{53 \text{ days}} \times 1 \text{ day} = 0.0131$$

 $N_0 = \frac{\text{mass of sample}}{\text{molar mass}} \times \text{Avogadro number}$ $= \left(\frac{5.7 \times 10^{-12}}{7.0 \times 10^{-3}}\right) \times 6.02 \times 10^{23} = 4.902 \times 10^{14}$

$$= N_{a}e^{-\lambda t}$$

(ii)

(C)

(i)

(ii)

 $\cdot \lambda$

N

 $= 4.902 \times 10^{14} \times e^{-0.0131(120)}$

 $= 1.02 \times 10^{14}$

(iii) By emitting a γ-ray photon, a <u>radioactive beryllium nucleus only loses</u> [1] <u>energy</u> but <u>its composition remains unchanged</u>. Hence, the number [1] of beryllium nuclei is a constant.

