H2 Physics Syllabus 9749

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> Marcus Soh 2018-2019 Victoria Junior College

"The only limits you have are the limits you believe."

Wayne Dyer

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Certain images were made using the online tool https://www.draw.io.

Conventions

Every definition that should be memorised will be highlighted in green.

Every formula that accompanies a definition and / or should be memorised will be highlighted in yellow.

Where applicable, content not in syllabus is boxed up with the heading "EXTRA!".

Disclaimer

Although I have tried to ensure that everything in these notes are factually correct, some errors may arise. In which case, please send an email to:

marcussoh38@gmail.com

with the details of the error, with "H2PHYE" as the subject.

These notes are <u>free to distribute</u>, however please credit me if you do. I appreciate your honesty.

Chapter 1: Physical Quantities and Measurements

1 Quantities

Base quantities (units in brackets): mass (kg), length (m), time (s), current (A), temperature (K), amount of substance (mol) Common derived quantities: force (N or kgms⁻²), work (J or kgm²s⁻²), power (W or kgm²s⁻³)

2 Homogeneity

A physical equation is homogeneous if <u>every term in the equation has the same</u> <u>units</u>.

3 Measurements

Prefixes:

 10^{12} : tera- (T) 10^{9} : giga- (G) 10^{6} : mega- (M) 10^{3} : kilo- (k) 10^{-1} : deci- (d) 10^{-2} : centi- (c) 10^{-3} : milli- (m) 10^{-6} : micro- (µ) 10^{-9} : nano- (n) 10^{-12} : pico- (p)

Common measurements:

Radius of the earth: $6.4 \times 10^6 \text{ m}$

Thickness of paper: 1×10^{-4} m

Diameter of H atom: 1×10^{-10} m

Length of football field: $1 \times 10^2 \text{ m}$

Mass of basketball: 0.60 kg

4 Conversions

e.g. $10 \text{ cm}^3 \text{ to } \text{m}^3$: $10 \text{ cm}^3 = 10 \times (10^{-2})^3 \text{ m}^3 = 10 \times 10^{-6} \text{ m}^3$ e.g. 50 kmh^{-1} to ms^{-1} : $50 \text{ kmh}^{-1} = \frac{50 \text{ km}}{1 \text{ h}} = \frac{50 \times 10^3 \text{ m}}{60 \times 60 \text{ s}} = (50 \div 3.6) \text{ms}^{-1}$

5 Vectors and Scalars

Vectors: magnitude + direction

Note: area is a vector whose direction is perpendicular (normal) to its surface

6 Vector manipulation

To find the resultant vector,

For a right-angled triangle, use Pythagoras' Theorem $(A^2 + B^2 = R^2)$ For a non-right-angled triangle, use cosine rule $(R = \sqrt{A^2 + B^2 + 2AB \cos \theta})$ Vector subtraction: a - b = a + (-b)Vector resolution: resolve vector into 2 mutually perpendicular components $(|b_x| = |b| \cos \theta, |b_y| = |b| \sin \theta)$

b

Chapter 2: Errors and Uncertainties

1 Systematic errors

Systematic errors are errors of measurements that occur according to some <u>'fixed</u> rule or pattern' such that they yield a <u>consistent</u> overestimation or underestimation of the true value.

The error is systematic if (1) <u>repeated measurements taken under the same</u> <u>conditions yield the same error in magnitude and sign</u> or if (2) <u>its value changes in a</u> <u>predictable manner based on the surroundings</u>.

e.g. zero error, faulty instruments, wrong assumptions, wrong experimental technique

Systematic errors are <u>NOT</u> reducible by taking the average. Instead, the error can be removed by calibrating instruments or by changing the experimental technique.

2 Random errors

Random errors are errors with <u>different magnitudes and signs</u> in repeated measurements. They occur due to the inability to obtain the true value of the measured quantity.

The error may be random due to (1) variation in the condition of <u>measuring</u> <u>instruments</u>, (2) variation due to the inability of an <u>observer</u> to measure <u>small</u> <u>intervals</u>, (3) variation due to <u>fluctuating external conditions</u>, or (4) <u>irregularity</u> of a quantity to be measured.

e.g. human reaction time, parallax error

Random errors are reducible by <u>taking a number of readings and finding the</u> <u>average</u>.

3 Precision and Accuracy in data

Precision: the 'spread' of the data Accuracy: how close the average value of the data is to the 'true' value

4 Precision and Accuracy in instruments

Precision can also refer to the smallest value an instrument can measure. Precision is reflected by the actual uncertainty of an instrument (to 1 s.f.).

Accuracy refers to how close the measured value is from the true value. Accuracy is reflected by the fractional or percentage uncertainty (to 2 s.f.).

5 Actual, Fractional and Percentage uncertainties

Rules: (for $R \pm \Delta R$)

1. Actual uncertainty ΔR is always expressed to 1 s.f.

2. R is expressed to the same number of decimal places as in ΔR .

e.g. (6.6 ± 0.1) cm

6 Uncertainties of calculated values

The First Principles is a way to calculate the uncertainty of a function. It can be used in all scenarios.

For example,

$$R = (1.2 \pm 0.1) \,\mathrm{m}$$

Therefore,

$$\Delta R = \frac{R_{max} - R_{min}}{2} = \frac{1.3 - 1.1}{2} = 0.1$$

Case	Method to find uncertainty (actual / fractional)	
Addition or subtraction	Final <u>actual</u> uncertainty is the sum of all absolute uncertainties. OR First principles	
Multiplication or division	Final <u>fractional</u> uncertainty is the sum of the individual fractional uncertainties. OR First principles	
$R = y^n$ where $n \in \mathbb{R}$	$\frac{\Delta R}{R} = n \frac{\Delta y}{y}$ where $n \in \mathbb{R}$ OR First principles	
$R = a^n + b$ where $n \in \mathbb{R}$	Only First principles can be used.	

Chapter 3: Kinematics

1 Displacement (symbol *s*, base units m)

Shortest distance from the initial to the final position of a body along a specific direction.

The area under a velocity – time graph gives the change in displacement.

2 Velocity (symbol v, base units ms⁻¹)

Rate of change of displacement with respect to time.

The gradient of a displacement – time graph gives the instantaneous velocity. The area under an acceleration – time graph gives the change in velocity.

$$v = \frac{ds}{dt}$$

3 Acceleration (symbol *a*, base units ms^{-2})

Rate of change of velocity with respect to time.

The gradient of a velocity – time graph gives the instantaneous accleration.

$$a = \frac{dv}{dt}$$

4 Free Fall

No air resistance: Downwards acceleration of 9.81 ms⁻²

5 Terminal Velocity in a Viscous Liquid

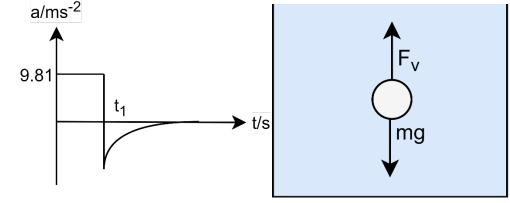
Terminal velocity is the maximum velocity obtained when an object is falling under the influence of gravity.

Assume that a small sphere is dropped from a height *h* above a viscous fluid. Since the sphere is small and density of air is low, we can assume that a = g when sphere is falling in air (see below for the case of air resistance).

When the sphere enters the fluid (e.g. water), the sphere will decelerate and its velocity decreases.

$$mg - F_v = ma$$

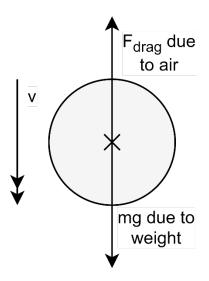
When $mg = F_v$, the sphere has reached terminal velocity.



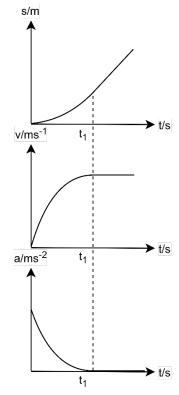
6 Terminal Velocity in air

Air resistance present: Still accelerate downwards, but as velocity increases, air resistance increases until acceleration reaches 0.

The maximum velocity attained is the terminal velocity.



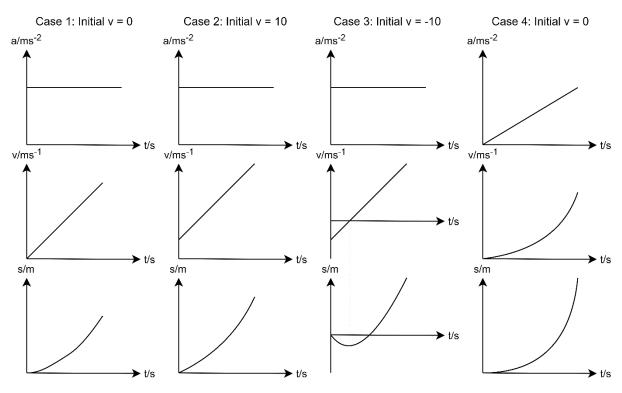
- 1. When object starts from rest, initial v = 0 and only experiences weight = mg a. F = ma becomes F = mg ($g = 9.81 \text{ ms}^{-2}$)
- 2. As object gains speed in the downwards direction, it experiences increasing air resistance against its motion (in the upward direction)
 - a. Net force experienced is the resultant of weight (mg) and air resistance (F_{air})
 - b. Net force = $mg F_{air}$ (which magnitude is smaller than mg)
 - c. $mg F_{air} = ma$ $\therefore a = g - \frac{F_{air}}{m}$
- 3. As speed of body increases, air resistance increases until it is equal to the body's weight
 - a. $mg F_{air} = ma$
 - $\Rightarrow 0 = ma$
 - $\Rightarrow a = 0$
 - \Rightarrow Body falls at constant velocity (terminal velocity at time t_1)



6 Graphs of Motion

Note that

- Gradient of the s-t graph is the velocity at that time
- Gradient of the v-t graph is the acceleration at that time
- Area under the a-t graph is the change of velocity
- Area under the v-t graph is the change of displacement.

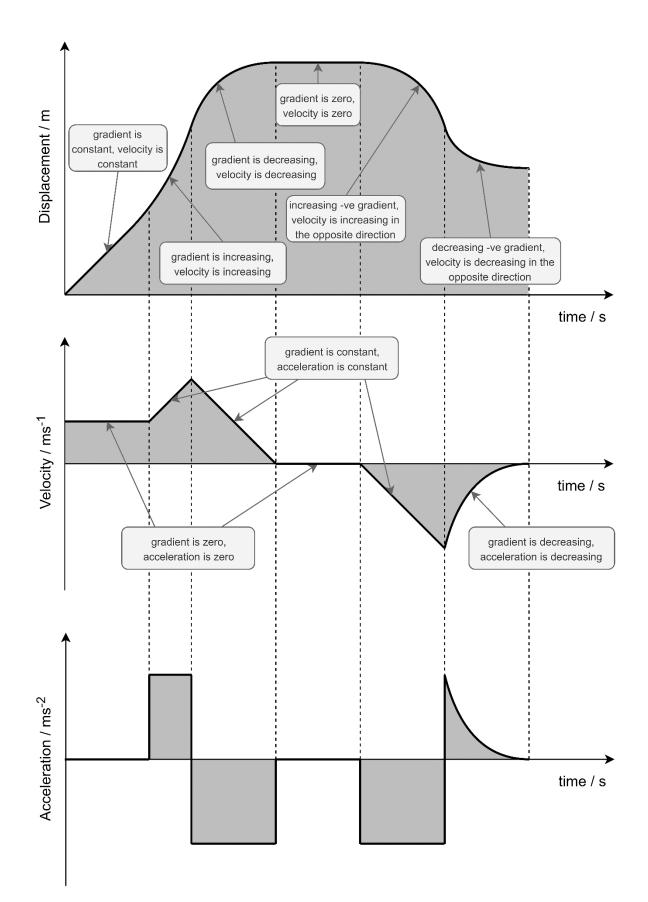


Relationship between displacement (s), velocity (v) and acceleration (a)

displacement *s*
differentiate
$$\left(v = \frac{ds}{dt}\right) \downarrow \uparrow$$
 integrate $\left(\Delta s = \int_{t_1}^{t_2} v \, dt\right)$
velocity *v*
differentiate $\left(a = \frac{ds}{dt}\right) \downarrow \uparrow$ integrate $\left(\Delta v = \int_{t_1}^{t_2} a \, dt\right)$
acceleration *a*

EXTRA!

Jerk comes after acceleration and is the rate of change of acceleration. Then comes snap, crackle, pop. These names are not standard, but they have been used to refer to the 4th, 5th and 6th derivatives respectively.



7 Equations of motion

1. Deriving v = u + at

 $a = \frac{dv}{dt} =>$ acceleration is equal to the gradient.

From the graph of velocity against time for an object undergoing constant acceleration, the gradient is equal to the acceleration. Hence,

$$a = \frac{v - u}{t}$$
$$v = u + at$$

2. Deriving $s = \frac{1}{2}(u+v)t$

Undergoing constant *a*,

total s travelled

Average $v = \frac{\text{total s travene}}{\text{total t taken}}$

- \Rightarrow Total s travelled = average v × total t taken
- \Rightarrow Velocity under constant *a* will increase uniformly with *t*, from initial value *u* to final value *v*, within *t* seconds

$$\Rightarrow$$
 Average v (mean v) = $\frac{v+u}{2}$

$$\Rightarrow \therefore s = \frac{v+u}{2} t = \frac{1}{2} (v+u)t$$

3. Deriving $s = ut + \frac{1}{2}at^2$

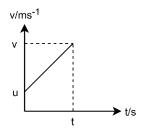
The area of a velocity-time graph is the displacement. Hence,

$$s = \frac{1}{2}(u+v)(t)$$

$$\frac{1}{2}(u+u+at)(t)$$

$$s = ut + \frac{1}{2}at^{2}$$
v/ms⁻¹
v
u

4. Deriving $v^2 = u^2 + 2as$ From (1), $t = \frac{v-u}{a}$ Substituting (1) into (3) $\Rightarrow s = u\left(\frac{v-u}{a}\right) + \frac{1}{2}a\left(\frac{v-u}{a}\right)^2$ $\Rightarrow as = uv - u^2 + \frac{1}{2}(v-u)^2$ $\Rightarrow as = uv - u^2 + \frac{1}{2}(v^2 + u^2 - 2vu)$ $\Rightarrow as = \frac{1}{2}v^2 - \frac{1}{2}u^2$ $\Rightarrow \therefore 2as = v^2 - u^2$



→ t/s

t

8 Projectile Motion

An object thrown at angle will have its weight as the only force acting vertically downwards.

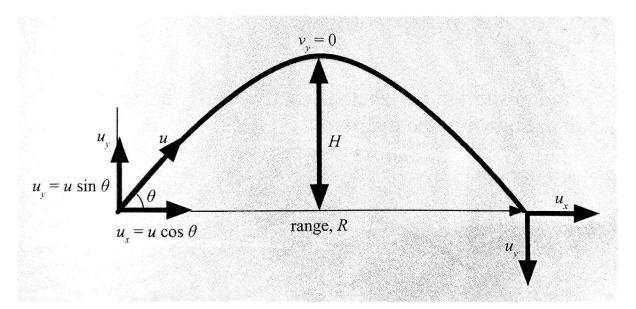
The only acceleration the object experiences is due to gravity, $g = 9.81 \text{ ms}^{-2}$. Air resistance is negligible.

There will be constant horizontal velocity.

Resolve the angle of motion to its horizontal and vertical components, then equations of kinematics to solve the problem.

	Horizontally (<i>x</i> -direction)	Vertically (y-direction)	
Displacement	$s_x = u_x t$	$s_y = u_y t - \frac{1}{2}gt^2$	
Velocity $u_x = u \cos \theta$		$v_y = u_y - gt$	
Acceleration $a_x = 0$ $a_y = -g$			
Note: Assuming upwards and to the right as positive for sign convention.			

- Path of projectile motion without air resistance is always symmetrical parabola (if starting and ending position are on the same level.)
- At the highest point of motion, vertical velocity = 0 (turning point)



Note: $u_y = -v_y$ and $u_x = v_x$

u and v can also be calculated using the formulae:

$$u = \sqrt{u_x^2 + u_y^2} \qquad \qquad v = \sqrt{v_x^2 + v_y^2}$$

Chapter 4: Forces

1 Non-contact forces

Gravitational, electric, magnetic

2 Contact forces

Normal contact, frictional, drag, elastic, upthrust

3 Hooke's Law

Hooke's Law: Within the proportional limit, the extension *e*, produced in a material is directly proportional to the force *F* applied.

F = ke

4 Spring constant (symbol *k*, units Nm^{-1} , base units kgs^{-2}) The spring constant is a measure of the <u>stiffness</u> of an object.

Springs connected in series:

$$\frac{1}{k_{eff}} = \frac{1}{k_1} + \frac{1}{k_2} + \dots + \frac{1}{k_N} \text{ or } k_{eff} = \frac{k}{N}$$

Springs connected in parallel:

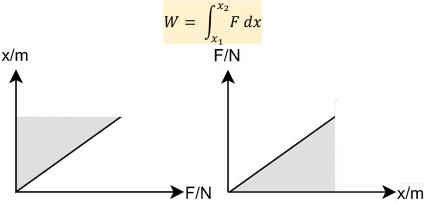
 $k_{eff} = k_1 + k_2 + \dots + k_N$ or $k_{eff} = Nk$

5 Elastic Potential Energy (EPE)

The energy stored in a <u>deformed</u> (extended or compressed) wire or spring which obeys Hooke's law is known as the EPE.

$$EPE = \frac{1}{2}Fx = \frac{F^2}{2k} = \frac{1}{2}kx^2$$

It can be found from the area under the force-extension graph.



6 Centre of Gravity

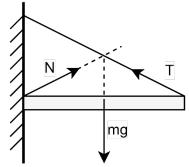
The entire mass of a body can be considered to <u>act at a single point</u> known as the centre of mass.

If a body is symmetrical and of uniform composition, then its centre of mass is at the <u>geometrical centre</u> of the body.

In uniform gravitational fields, the centre of mass coincides with the centre of gravity.

7 Free body diagrams

The purpose of free body diagrams is to analyse the forces acting <u>on</u> the object only. For forces in equilibrium, the 3 forces' line of action <u>must</u> coincide at the same point.



8 Moments (symbol τ , units Nm, base units kgm²s⁻²)

The turning effect of a force is called its moment or torque.

A moment is defined as the product of the force and the <u>perpendicular</u> distance from the line of action of the force to the pivot.

 $\tau = F \times d$

Resolve F if perpendicular distance not given.

Couple: 2 equal and opposite forces whose lines of action do <u>not</u> coincide. They produce rotation only as the net force is zero.

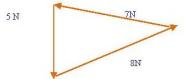
 $\tau = F \times perpendicular distance between the 2 forces$

Principle of moments: For an object in rotational equilibrium, the sum of the anticlockwise moments about the pivot is equal to the sum of clockwise moments about the same pivot.

 $\tau_{anticlockwise} = \tau_{clockwise}$

9 Equilibrium

A closed vector triangle can be formed from the force vectors.



Condition 1: The resultant force acting on the object must be zero.

$$\sum F = 0$$

Condition 2: The resultant moment acting on the object about any point must be zero.

$$\sum \tau = 0$$

10 Pressure (symbol p, units Pa, base units kgm⁻¹s⁻²) The pressure at a point X due to a column of liquid is given by

$$p = \rho g h$$

Derivation:

$$p = \frac{F}{A} = \frac{mg}{A} = \frac{\rho Vg}{A} = \frac{\rho Ahg}{A} = \rho gh$$

11 Upthrust

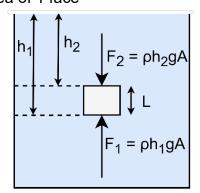
For an object in a liquid, the pressure difference between the top and bottom of the object provides an upward force, or upthrust.

Upthrust can be calculated as the weight of liquid displaced by the object.

Upthrust =
$$(\rho h_1 g)A - (\rho h_2 g)A = \rho LgA$$

where h_1 is the height of the bottom surface of the object under water h_2 is the height of the top surface of the object under water L is the height of the object

A is the surface area of 1 face



12 Principle of Flotation

For an object in equilibrium, the upthrust is equal to the weight of the object.

 $\rho LgA = mg$

Chapter 5: Dynamics

1 Inertia

Inertia is defined as the reluctance of a body to change its state of motion or rest. Inertia is affected by mass.

2 Mass and Weight

Mass is the amount of matter in a body. Weight is due to the effect of gravity on a mass.

3 Momentum (symbol p, base units kgms⁻¹) The product of mass and velocity of an object. p = mv

4 Newton's Laws

Newton's First Law: Every body continues in its <u>state of rest</u> or <u>uniform motion in a</u> <u>straight line</u> unless an unbalanced (resultant) force acts on it to change its state.

Newton's Second Law: The <u>rate of change of momentum of a body</u> is proportional to the unbalanced (resultant) force acting on it, and <u>occurs in the direction of the force</u>.

$$F \propto \frac{dp}{dt}$$

Derivation of F = ma

$$F = \frac{dp}{dt} = \frac{d(mv)}{dt} = m\frac{dv}{dt} = ma$$

Note: resultant force is always in the same direction as the acceleration.

Newton's Third Law: If body A exerts a force (action) on body B, then body B will exert an <u>equal and opposite force</u> (reaction) on body A.

5 Impulse (symbol *I*, units Ns, base units $kgms^{-1}$) Impulse is defined as the integral of the force over the time interval in which the force acts.

$$Impulse = \int_{t_1}^{t_2} F \, dt$$
$$I = F_{avg} \Delta t$$

Impulse can be thought of as the area under the force-time graph.

Impulse-momentum theorem states that impulse is equal to the change in momentum.

$$F_{avg}\Delta t = \Delta p \text{ or } F_{avg} = \frac{\Delta p}{\Delta t}$$

A small Δt leads to a large *F* and vice versa.

Hence, when falling, we usually bend our knees when we land after a jump to have a larger Δt and hence a smaller force on our knees. This reduces the risk of injury.

5 Collisions

Principle of Conservation of Momentum (p.o.c.o.m.): The total momentum of a system is conserved <u>provided no external unbalanced</u> <u>force acts on the system</u>.

$$m_1 u_1 + m_2 u_2 = m_1 v_1 + m_2 v_2$$

p.o.c.o.m. is always obeyed for every type of collision.

Elastic collisions: <u>Kinetic energy</u> of the system is conserved.

$$\frac{1}{2}m_1u_1^2 + \frac{1}{2}m_2u_2^2 = \frac{1}{2}m_1v_1^2 + \frac{1}{2}m_2v_2^2$$

Conservation of momentum and kinetic energy (p.o.c.o.m & c.o.e.) leads to the relative speed equation:

$$u_1 - u_2 = v_2 - v_1$$

or relative speed of approach = relative speed of separation. (r.s.a. = r.s.s.)

Inelastic collisions: Kinetic energy of the system is not conserved.

Totally inelastic collisions: Kinetic energy of the system is <u>not conserved</u> and both objects <u>stick to each other</u> after collision. (coalescence)

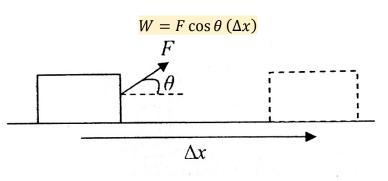
$$m_1 u_1 + m_2 u_2 = (m_1 + m_2) v$$

EXTRA!

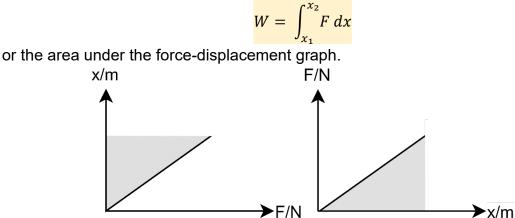
If an object collides into another object of equal mass that is originally stationary obliquely (not head on) and they travel off at an angle relative to one another, if the collision is elastic, the angle at which they separate will be 90°.

Chapter 6: Work, Energy, Power

1 Work done by a force (symbol W, units J, base units kgm^2s^{-2}) The work done by a constant force on an object is defined as the <u>product of the</u> <u>magnitude of the displacement</u> and the <u>component of the force</u> parallel to the displacement. In general,



Alternatively, if the force is not constant, then



2 Work done by a gas

For a gas enclosed in a cylinder, the work done by the gas in expanding at constant pressure is given by:

$$W = p\Delta V$$

In general, the work done by a gas in changing the volume from V_1 to V_2 is given by:

$$W = \int_{V_1}^{V_2} p \, dV$$

or the area under the pressure-volume graph.

3 Common examples of energy

Chemical, nuclear, radiant, electrical, internal, kinetic, gravitational potential, elastic energy

4 Kinetic energy

Kinetic energy of an object is the energy it possess by virtue of its motion.

$$E_k = \frac{1}{2}mv^2$$

Derivation of equation:

Consider an object of mass m moving to the right under a constant force F_{net} . Then,

$$F_{net} = mo$$

From kinematics,

$$v^2 = u^2 + 2as \rightarrow s = \frac{v^2 - u^2}{2a}$$

 $W = F_{net}s = mas$

From forces,

$$W = ma\left(\frac{v^2 - u^2}{2a}\right) = \frac{1}{2}m(v^2 - u^2)$$

Therefore,

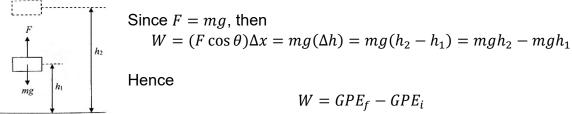
 $W = \Delta KE$

For a body initially at rest, u = 0, and where work done is fully converted into kinetic energy, the KE of an object of mass *m* moving with speed *v* is defined as the equation above.

This equation is an expression for the Work Energy theorem, which states that the net work done by forces acting on a particle is equal to the change in kinetic energy of the particle.

5 Gravitational potential energy

Consider a book being raised from a height h_1 to a height h_2 by a constant force F equal and opposite to the weight mg of the book, ignoring air resistance.



We define the gravitational potential energy of a body as the product of its weight mg and its height h above a reference level. Hence

GPE = mgh

6 Different forms of potential energy

Type of P.E.	Due to
Gravitational	2 masses
Electric	An electric current
Elastic	A deformed spring

7 Conservation of energy

The Principle of Conservation of Energy (c.o.e.) states that energy can neither be <u>created</u> nor <u>destroyed</u>, but can be <u>converted</u> from one form to another. The total energy in a <u>closed system</u> is always <u>constant</u>.

Situation	Energy converted	
	From	То
Bullet strikes block and	KE of bullet	Internal energy, KE, GPE
becomes embedded in it		of bullet and block
Shooting an arrow using	Elastic PE of the bow	KE of the arrow
the bow		
Swinging pendulum	GPE ←→KE	

In general,

8 Power (symbol *P*, units W, base units kgm^2s^{-3}) Power is defined as the <u>rate at which work is done</u> or the <u>rate at which energy is</u> transferred.

$$P = \frac{\Delta W}{\Delta t}$$

For a constant force F acting on an object with velocity v,

$$P = Fv$$

Derivation:

$$P = \frac{dW}{dt} = \frac{d(Fx)}{dt} = F\frac{dx}{dt} = Fv$$

9 Efficiency

In reality, no device can convert its total work done into useful energy. Most practical devices dissipate energy, usually due to fraction.

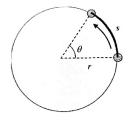
Efficiency can be defined as the percentage ratio of useful power output to power input.

$$\eta = \frac{\text{useful power output}}{\text{power input}}$$

Chapter 7: Motion in a Circle

1 Angular displacement (symbol θ , units rad)

Angular displacement is given by the formula



 $\theta = \frac{s}{r}$

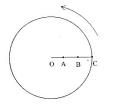
Can be thought of as the angle "travelled". The radian is the angle subtended at the centre of a circle by an arc of equal length to the radius.

2 Angular velocity (symbol ω , units rad s⁻¹) Angular velocity is the rate of change of angular displacement.

$$\omega = \frac{d\theta}{dt}$$

Period *T* is the time taken for one complete revolution. For one revolution, $\theta = 2\pi$. Hence

$$\omega = \frac{2\pi}{T} = 2\pi f$$



Note: for a body rotating in circular motion at constant speed, the angular velocity of the object is the same at every point on the body. However, the velocity at different points on the body varies.

$$v_C > v_B > v_A$$
 but $\omega_A = \omega_B = \omega_C$

3 Tangential velocity (symbol *v*)

Tangential velocity refers to the linear velocity of a point moving in a circle. It is directed tangentially to the circular path.

$$s = r\theta$$
$$\frac{ds}{dt} = r\frac{d\theta}{dt}$$
$$v = r\omega$$

4 Centripetal acceleration (symbol a)

Consider a mass *m* moving from point A to B in a circular path of radius *r* with uniform speed *v*. At A, its initial velocity is v_a and at B, its velocity is v_B . (magnitude is the same, but direction is different)

Hence $\Delta v = v_B - v_A = v_B + (-v_A)$

By small angle approximation, $\Delta v = v\theta$

$$a = v\omega = r\omega^2 = \frac{v^2}{r}$$

Centripetal acceleration is always directed towards the centre.

5 Centripetal force (symbol *F_c*)

Qualitative explanation of motion in a curved path:

Newton's first law states that an object in motion will continue in its state of uniform motion <u>in a straight line</u> unless an external force acts on it. Hence, for an object moving with constant speed to execute uniform circular motion, an external force must act on the object.

According to Newton's second law, this resultant force is proportional to its acceleration. Hence this resultant force acts in the same direction as the acceleration, i.e. perpendicular to the motion and directed towards the centre. This force is known as the centripetal force.

$$F_c = ma = m\left(\frac{v^2}{r}\right) = mr\omega^2$$

Note: Centripetal force is a <u>resultant force</u>. Hence, it should not be labelled in free body diagrams.

Note: Centripetal force is <u>never constant</u> because its direction is always changing.

Chapter 8: Gravitational Field

1 Law of Universal Gravitation

Newton proposed a law, which says:

Two particles attract each other with a <u>force proportional to the product of their</u> <u>masses</u> and <u>inversely proportional to the square of their distance apart</u>.

$$F = \frac{GMm}{r^2}$$

Note: As gravitational force is always attractive, the force is always negative, i.e.

$$F = -\frac{GMm}{r^2}$$

However, in calculations, the negative sign can be dropped. This is because gravitational force is always attractive in nature (i.e. there is only 1 type, negative). r^2 is the square of the distances between their geometrical centres.

2 Gravitational field (strength) (symbol g, units N kg⁻¹, base units ms⁻²) The gravitational field strength g at a point in a gravitational field is defined as the <u>force per unit mass</u> acting on any mass placed at that point.

$$g = \frac{F}{m}$$

A mass *M* establishes a field of strength *g*

$$g = \frac{\frac{GMm}{r^2}}{m} = \frac{GM}{r^2}$$

Due to sign convention, gravitational field is negative. However, the negative sign can be dropped as it is always attractive in nature.

Near the earth's surface, g is approximately constant at 9.81 ms^{-2} .

3 Satellite motion

For a satellite of mass m orbiting a planet of mass M, it is said that the gravitational force acting on the satellite *provides* the centripetal force.

Hence, for all satellites, gravitational force provides the centripetal force.

$$\frac{F_G = F_C}{\frac{GMm}{r^2} = \frac{mv^2}{r}}$$

In this equation, the orbital velocity can be calculated.

$$v_{orb} = \sqrt{\frac{GM}{r}}$$

4 Geostationary satellites

A geostationary satellite is a satellite that is doing <u>circular orbit</u> above the <u>equator of</u> <u>the Earth</u> and moving in a <u>direction similar to the earth's rotation</u> with a <u>period</u> <u>exactly equal to the period of the Earth</u> as it turns about its axis, which is 24 hours. The orbit is also sometimes called a 'parking orbit'.

For a geostationary satellite, gravitational force provides the centripetal force.

5 Kepler's third law

Consider a planet of mass m orbiting a star of mass M. Let the distance between them be r. Since gravitational force provides the centripetal force,

$$F_{G} = F_{C}$$

$$\frac{GMm}{r^{2}} = mr\omega^{2}$$

$$\frac{GMm}{r^{2}} = mr\left(\frac{2\pi}{T}\right)^{2}$$

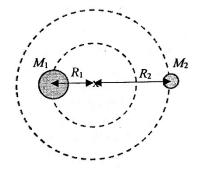
$$T^{2} = \frac{4\pi^{2}}{GM}r^{3} \text{ or } T^{2} \propto r^{3}$$

6 Gravitational potential energy (symbol *U*, units J, base units kgm^2s^{-2}) With <u>infinity</u> being taken as the zero reference, the GPE at a point is the <u>work done</u> by an <u>external agent</u> in bringing a <u>test mass</u> from infinity to the point without a change in kinetic energy.

$$U = -\frac{GMm}{r}$$

Note: The negative sign cannot be dropped in calculations as U is zero at infinity. The sign implies the attractive nature of the gravitational field.

7 Binary stars



A binary star system is one where two stars rotate about their common centre of mass.

To find the period of rotation:

$$F_G = \frac{GM_1M_2}{(R_1 + R_2)^2}$$

For mass
$$M_1$$
,
 $F_{CM_1} = M_1 R_1 \omega^2$. Since $F_G = F_{CM}$,
 $\frac{GM_1M_2}{(R_1+R_2)^2} = M_1 R_1 \omega^2$ so $\frac{GM_2}{(R_1+R_2)^2} = R_1 \omega^2$ —(1)

For mass M_2 , $F_{CM_2} = M_2 R_2 \omega^2$. Since $F_G = F_{CM}$, $\frac{GM_1M_2}{(R_1+R_2)^2} = M_2 R_2 \omega^2$ so $\frac{GM_1}{(R_1+R_2)^2} = R_2 \omega^2$...(2) (1) + (2):

$$\frac{G(M_1 + M_2)}{(R_1 + R_2)^2} = (R_1 + R_2) \left(\frac{2\pi}{T}\right)^2$$
$$T = \sqrt{\frac{4\pi^2(R_1 + R_2)^3}{G(M_1 + M_2)}}$$

Calculating GPE:

$$GPE_{1} = \frac{GM_{1}M_{2}}{R_{1} + R_{2}} \qquad GPE_{2} = \frac{GM_{1}M_{2}}{R_{1} + R_{2}}$$
$$GPE_{T} = \frac{2GM_{1}M_{2}}{R_{1} + R_{2}}$$

8 Escape speed (symbol v_e)

Escape speed is the velocity which will allow the object to go to infinity and never return. Assuming that at infinity, the speed of the object is zero,

$$\frac{1}{2}mv^{2} \geq \frac{GMm}{r}$$

$$v_{e} = \sqrt{\frac{2GM}{r}}$$

where v_e is the minimum escape speed.

9 Gravitational potential (symbol ϕ , units J kg⁻¹, base units m²s⁻²) The gravitational potential at a point is the <u>work done per unit mass</u> by an <u>external</u> agent in bringing an object from infinity to the point.

Similarities		Differences	
GPE U	${\sf GP} \phi$	GPE U	$GP\phi$
Measure of	work done	Work done for the whole object	Work done per unit mass
Sca	lars		-
Infinity as ze	ero reference		

$$\phi = \frac{U}{M} = -\frac{GM}{r}$$

10 Orbital energies

The total energy of a satellite of mass m moving with speed v around the earth in an orbit of radius r around the earth of mass M is

$$E_T = K + U$$

$$E_T = \frac{1}{2}mv^2 + \left(-\frac{GMm}{r}\right) - -(1)$$

For the circular motion of the satellite in orbit:

$$\frac{GMm}{r} - \frac{mv^2}{r}$$

$$\frac{1}{r^2} = \frac{1}{r}$$
$$mv^2 = \frac{GMm}{r} \implies K = \frac{GMm}{2r}$$
Therefore,

$$E_T = K + U = \frac{GMm}{2r} + \left(-\frac{GMm}{r}\right) = -\frac{GMm}{2r}$$

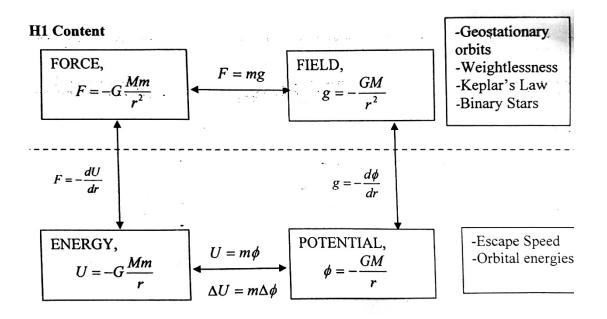
11 Relating F, U, g, ϕ

The gravitational force F can be derived from the negative of the potential energy gradient.

$$F = -\frac{dU}{dr}$$

The gravitational field strength g can be derived from the negative of the potential gradient.

$$g = -\frac{d\phi}{dr}$$



Chapter 9: Temperature and Ideal Gases

1 Temperature

A measure of the degree of hotness or coldness of an object. Rises when thermal energy is supplied to an object, unless there is a phase change. Heat flows from regions of high temperature to regions of low temperature.

Temperature of a substance is directly related to the amount of kinetic energy in the substance. The higher the temperature, the higher the kinetic energy.

Kinetic energy comes in 3 forms: Translational, Rotational and Vibrational

2 Internal Energy

Internal energy, U, is the sum of the kinetic energy and intermolecular potential energy of all the molecules of the system.

U = KE + PE of all molecules

3 Thermal Equilibrium

Heat energy is transferred from a hot body to a cooler body through 3 mechanisms: Conduction, Convection and Radiation.

2 bodies in thermal contact will exchange energy between them, but if there is no net heat flow, they are in thermal equilibrium and at the same temperature.

4 Empirical Scales

Based on the observation of thermometric properties as they change with temperature.

Usually has 2 fixed points which assumes a certain thermometric property varies linearly with temperature. The quantity of the thermometric property must have a unique value at every temperature. E.g. Must show a one-one graph For the centigrade scale, the fixed points are the steam point (100°C) and ice points

For the centigrade scale, the fixed points are the steam point (100°C) and ice point (0°C).

If the value X is the value of the thermometric property, then

$$\frac{\theta}{100} = \frac{X_{\theta} - X_i}{X_s - X_i}$$

where X_{θ} is the value of X at temperature θ

 X_i is the value of X at ice point

 X_s is the value of X at steam point

Assumption of linearity of the thermometric properties leads to inaccurate temperature readings as it is inherently a wrong assumption. Instead, the actual behaviour of the thermometric property is non-linear. Thus empirical scales are always slightly wrong, except at the fixed points.

Thus, thermometers based on different thermometric properties will disagree, except at the 2 fixed points.

5 Kelvin (absolute) Scale

The lower point is absolute 0. This is the lowest possible temperature all matter can have, where all atoms have the minimum possible kinetic energy possible.

At absolute 0, all atoms have minimal internal energy.

The upper fixed point is the triple point of water, which is assigned 273.16 K. This is the temperature where water can exist in all 3 states in equilibrium. It was chosen as it is easily reproducible and precise.

Thus the Kelvin, K, is defined as $\frac{1}{273.16}$ of the triple point of water. Melting point of water is 273.15 K.

6 Celsius scale

It uses the same value as the centigrade scale, where 0°C is the melting point of water and 100°C is the boiling point of water. Thus,

$$X/(^{\circ}C) = K/(K) - 273.15$$

7 Ideal Gas Behaviours

Ideal gases are a collection of perfectly hard spheres which collide but otherwise do not interact with each other. No real gases obey this completely, but they almost behave like ideal gases at high temperature and at low pressures. The ideal gas equation

$$pV = nRT = NkT$$

$$R = 8.31 \text{ J mol}^{-1} \text{ K}^{-1} = k \frac{N}{n} = kN_A$$
where *k* is the Boltzmann's constant, $k = \frac{R}{N_a} = 1.38 \times 10^{-23} \text{ J K}^{-1}$

Boyle's law	0	$p \propto \frac{1}{u}$
	inversely proportional to its volume if the	$p \propto \frac{1}{V}$
	temperature is constant.	M,T = constant
Charles' law	Volume of a fixed mass of gas is	
	proportional to its thermodynamic	$V \propto T$
	temperature if the pressure of the gas is	M, p = constant
	constant.	
Pressure law	Pressure of a fixed mass of gas is	m or T
(Gay-Lussac law)	proportional to the thermodynamic	$p \propto T$
	temperature if its volume is constant.	M, V = constant
Avogadro's law	The volume of a gas is proportional to the	$V \propto n$
-	number of moles of the gas if its pressure	
	and temperature are constant.	p,T = constant

Ideal gases obey the following laws.

8 Avogadro's constant and mole

A mole is the amount of substance that contains as many particles as there are atoms in 0.012 kg of carbon-12.

This number is known as Avogadro's constant, $N_a = 6.02 \times 10^{23} \text{ mol}^{-1}$

Molar mass M_m is the mass of 1 mole of a substance.

If each molecule has mass m, $M_m = m \times N_a$.

9 The kinetic theory

An attempt to explain the macroscopic properties of the gas by looking at its microscopic properties.

Assumptions of ideal gas using the acronym DAVE:

Duration of impact among molecules and with walls are negligible.

<u>A</u>ttraction forces among molecules are negligible and do not exert forces on each other unless they collide.

<u>V</u>olume of the atoms or molecules is negligible compared to the volume of the container.

Elastic collisions, so that molecules do not lose any kinetic energy from collisions.

10 Root mean square speed

The molecules in a gas do not travel at the same speed, thus root mean square speed is used like an average speed.

Let M be the total mass of the gas, and m be the mass of one molecule.

$$M = Nm$$

$$C_{rms} = \sqrt{\langle c^2 \rangle} = \sqrt{\frac{c_1^2 + c_2^2 + \dots + c_N^2}{N}}$$
$$c_{rms} = \sqrt{\frac{3RT}{M}} = \sqrt{\frac{3kT}{m}}$$

Total kinetic energy is translational energy, $\frac{1}{2}Mc_{rms}^{2}$

$$p = \frac{1}{3} \frac{M}{V} < c^{2} > \text{where } M \text{ is the mass of the gas}$$

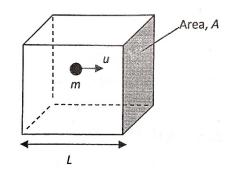
$$p = \frac{1}{3} \rho < c^{2} > \text{where } \rho \text{ is the density of the gas}$$

$$Mean \ KE = \frac{3}{2} kT = \frac{1}{2} m \langle c^{2} \rangle$$

$$Total \ KE = \frac{3}{2} NkT = \frac{3}{2} nRT$$

Derivation of $p = \frac{1}{3} \frac{M}{V} \langle c^2 \rangle$ and $p = \frac{1}{3} \rho \langle c^2 \rangle$

We start by analysing the momentum of a random molecule in the diagram shown on the right.



Initial momentum:
$$p_i = mu$$

Final momentum: $p_f = -mu$
Change in momentum: $\Delta p = p_f - p_i$
 $\Delta p = -2mu$
Time between consecutive collisions on
the same wall:
Force / Rate of change of momentum
(N2L): $\frac{\Delta p}{\Delta t} = -2mu \div \frac{2L}{u}$
 $\frac{\Delta p}{\Delta t} = -\frac{mu^2}{L}$
 $F = -\frac{mu^2}{L}$
Force exerted by one molecule on wall
(N3L): $F = \frac{mu^2}{L}$
 $F = \frac{m(u_1^2 + u_2^2 + \dots + u_n^2)}{L}$
 $F = \frac{mN(u^2)}{L}$
where $\langle u^2 \rangle = \frac{u_1^2 + u_2^2 + \dots + u_n^2}{N}$

Pressure exerted by all molecules:

$$p = \frac{F}{A}$$

$$p = \frac{mN\langle u^2 \rangle}{L(L^2)}$$

$$p = \frac{mN\langle u^2 \rangle}{V} - eqn(1)$$

Since there are 3 directions, x, y, z, each molecule has velocities u, v, w in the x, y, z, direction. Hence, we have $c^2 = u^2 + v^2 + w^2$ Averaging, we have $\langle c^2 \rangle = \langle u^2 \rangle + \langle v^2 \rangle + \langle w^2 \rangle$ Since the motion is random and the number of molecules is very large, $\langle u^2 \rangle = \langle v^2 \rangle = \langle w^2 \rangle$ $\langle u^2 \rangle = \frac{1}{3} \langle c^2 \rangle$ From eqn (1): $p = \frac{mN}{V} (\frac{1}{3} \langle c^2 \rangle)$ $p = \frac{1}{3} (\frac{Nm \langle c^2 \rangle}{V})$ Nm = total mass of particles = M $p = \frac{1}{3} (\frac{M}{V} \langle c^2 \rangle)$ $\frac{M}{V} = \text{density} = \rho$ $p = \frac{1}{3} \rho \langle c^2 \rangle - -(1)$ Derivation of the relationship between c_{rms} and T of a given gas: From Ideal Gas law,

$$p = \frac{nRT}{V} - -(2)$$

Equating (1) and (2),

$$\frac{nRT}{V} = \frac{1}{3}\rho\langle c^2 \rangle$$
$$\frac{M}{M_m} * \frac{RT}{V} = \frac{1}{3}\rho\langle c^2 \rangle$$
$$\frac{\rho RT}{M_m} = \frac{1}{3}\rho\langle c^2 \rangle$$
$$\frac{RT}{M_m} = \frac{1}{3}\langle c^2 \rangle$$
$$\langle c^2 \rangle = \frac{3RT}{M_m}$$
$$c_{rms} = \sqrt{\frac{3RT}{M_m}}$$

Chapter 10: Thermodynamics

1 Internal Energy

Heat is defined as the energy that is transferred from one body at a higher temperature to one at lower temperature. Matter DOES NOT contain heat. Once transferred to an object, heat becomes internal energy.

Internal Energy, U, of a system is the sum of the KE (movement of molecules) and PE (Forces between molecules in the system)

U = KE + PE

When a substance absorbs or gives off heat, internal energy changes, either through changing KE or PE.

2 Temperature

All matter is made of atoms and molecules which are in constant random motion. Hence they all possess KE. Temperature is actually a measure of the average KE of a substance. This is the <u>microscopic</u> view of temperature.

When there is no net heat flow between 2 substances, they are in thermal equilibrium and said to be at the same temperature. This is the <u>macroscopic</u> view of temperature.

3 Internal energy and Temperature

Internal KE is related to the absolute temperature, T. But internal PE is related to the forces between the atoms or molecules, depending on their separation. When a system undergoes state change, PE increases but KE remains constant. Thus it cannot be said that an increase of Internal Energy reflects an increase in temperature.

4 Kinetic energy and Temperature

Temperature is related to the internal KE of the gas, NOT the bulk (i.e moving the whole container with gas inside) KE of the gas.

5 Internal Energy of an Ideal Gas

As the molecules do not exert any forces on each other, the Internal Energy of Ideal Gases is solely the KE of the molecules.

6 Heat Capacity

The <u>heat capacity</u> of a body is defined as the quantity of heat energy required to raise its temperature by 1 K.

C is the heat capacity of an object and has the units J K^{-1}

 $Q = C\Delta T$ where C is the heat capacity

The <u>specific heat capacity</u> of a body is defined as the heat required to raise the temperature of 1 kg of the substance by 1 K.

c is the specific heat capacity of an object and has the units J kg^{-1} K⁻¹

 $Q = mc\Delta T$ where c is the specific heat capacity

The molar heat capacity of a body is defined as the heat required to raise the temperature of 1 mole of a substance.

 C_n is the molar heat capacity of an object and has the units J mol⁻¹ K⁻¹

7 Latent Heat

The <u>specific latent heat of fusion</u> is of a substance is defined as the amount of heat energy needed to change a unit mass of the substance from solid to liquid state without a change in temperature.

 $Q = nC_n\Delta T$

The <u>specific latent heat of vaporisation</u> is of a substance is defined as the amount of heat energy needed to change a unit mass of the substance from liquid to gaseous state without a change in temperature.

A transition from one state of matter to another is known as a phase change. During a phase change, latent heat is given off or absorbed and does not change the temperature of the object.

 $Q = m\ell$ where ℓ is the specific latent heat of fusion/vaporisation

8 Kinetic Theory of matter

Solid: atoms are arranged closely together in a regular pattern. Intermolecular forces are strong and the motion of the atoms are limited to vibrations around their mean positions. Internal energy comprises of vibrational KE and PE.

Liquid: atoms are slightly further apart than in a solid. Occur in clusters and are free to slide around each other. Intermolecular forces are weaker. Internal energy comprises of KE (translational, rotational and vibrational) and more PE than solid.

Gas: atoms are very far apart from each other. Intermolecular forces are very weak and atoms move in a random manner. Internal Energy comprises of KE and PE (largest).

9 Melting

At the melting point, the amplitude of vibration of the atoms is large enough for atoms to move past the adjacent atoms and move to new positions (resulting in them being further apart) and hence have higher PE.

The latent heat of fusion = work required at the molecular level to transform from a ordered solid phase to a disordered liquid phase but does not increase the average KE. Hence temperature does not increase during melting. Hence,

$$\Delta U = 0, \Delta Q_s = \Delta W_{by}$$

10 Boiling

Atoms in a liquid are close together and the forces between them are stronger than the widely separated gas atoms.

The latent heat of vaporisation = work required to pull apart the atoms from being close together to being far apart in a gas.

Specific Latent heat of vaporisation > fusion because atoms must move further apart for boiling than melting. Additionally, work must be done against the external atmospheric pressure.

11 Cooling by evaporation

Evaporation occurs when a liquid becomes a vapour without a heat source. Atoms near the surface gain enough energy to overcome the attractive forces of the surrounding atoms and thus escape from the liquid.

It is the faster and more energetic atoms that evaporate. Thus, average KE of the remaining atoms decreases and thus the temperature decreases and evaporation is a cooling process.

12 First Law of Thermodynamics

The first law of thermodynamics states that the <u>internal energy</u> of a system depends only on its state, and that the increase in internal energy of the system is the sum of the heat <u>supplied to the system</u> and the external work done <u>on the system</u>.

Increase in internal energy = Heat supplied + External work done **on** the system $\Delta U = \Delta Q_s + \Delta W_{on}$

Term	Positive	Negative
U	Increase in internal energy	Decrease in internal energy
Q	Heat absorbed by the system	Heat given off by the system
W	Work done on the system	Work done by the system

Note: As from Chapter 9, internal energy is dependent on temperature ONLY for ideal gases ($E_k = \frac{3}{2}nRT$, negligible PE)

13 Calculations involving work done

At constant pressure, the work done <u>by</u> the gas in displacing a piston through a small distance dx is:

$$W_{by} = F \ dx = (pA) \ dx = p \ dV = p \Delta V$$

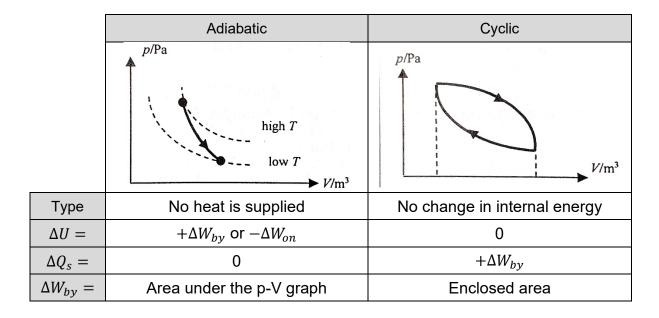
Therefore

$$W_{by} = \int_{V_i}^{V_f} p \, dV$$

Or, work done is the area under the pressure-volume graph.

14 Processes

	Isochoric	Isothermal	Isobaric
	p/Pa high T low T	p/Pa high T low T	p/Pa high T low T
Туре	Constant volume	Constant temperature	Constant pressure
$\Delta U =$	ΔQ_s	0	$\Delta Q_s + \Delta W_{by}$
$\Delta Q_s =$	Calculated using molar heat capacity $Q = mc\Delta\theta$ or $Q = mc_n\Delta\theta$	$+\Delta W_{by}$	Calculated using molar heat capacity $Q = mc\Delta\theta$ or $Q = mc_n\Delta\theta$
$\Delta W_{by} =$	0	Area under the p-V graph	$p\Delta V$



Chapter 11: Oscillations

1 Examples of free oscillations

Cylinder oscillating vertically in water

Spring-mass system oscillating horizontally on a frictionless surface

2 Definitions of some key terms

Amplitude:	maximum displacement of the oscillating body. Given by x_0 or A		
Period:	the time taken for one complete oscillation. Given by T		
Frequency:	the total number of oscillations per unit time. Given by $f = T^{-1}$ and has units of s ⁻¹ or Hz		
Angular frequency:	the product of 2π and the frequency. Given by $\omega = 2\pi f = 2\pi T^{-1}$		
Phase difference:	two particles are oscillating in phase if they are in the same state of disturbance. Otherwise, they are said to be out of phase. A phase difference of 2π means the particles are in phase, while a phase		

3 Simple harmonic motion (SHM)

Requirement 1: There must be a restoring force proportional to the displacement but directed opposite to it. For a spring-mass system,

$$F = -kx - -(1)$$

Requirement 2: The acceleration of the system must be proportional to the displacement but directed opposite to it.

difference of π means they are in antiphase.

$$F = ma - -(2)$$

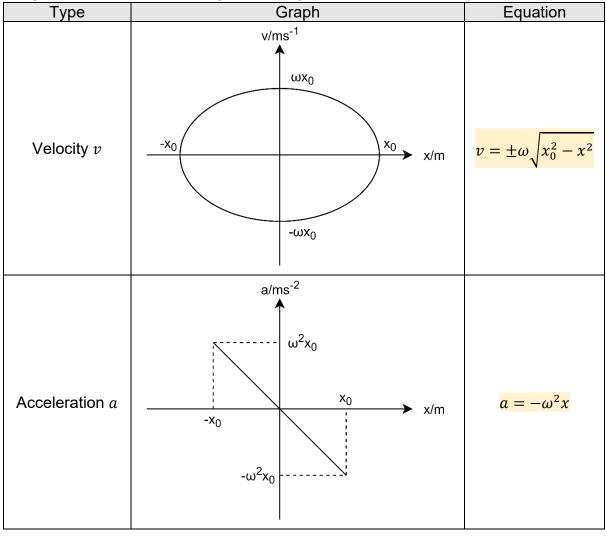
(1) = (2) $\rightarrow ma = -kx$
 $a = -\frac{k}{m}x \text{ or } a \propto -x$

If *a* is proportional to -x, the system is said to execute SHM.

4 Equations in SHM with respect to time t

Туре	Starting from equilibrium position ($x = 0 at t = 0$)	Starting from extreme position ($x = x_0 at t = 0$)
Displacement x	$x = x_0 \sin(\omega t)$	$x = x_0 \cos(\omega t)$
Velocity v	$v = \frac{dx}{dt} = \omega x_0 \cos(\omega t)$	$v = \frac{dx}{dt} = -\omega x_0 \sin(\omega t)$
Acceleration a	$a = \frac{dv}{dt} = -\omega^2 x_0 \sin(\omega t)$	$a = \frac{dv}{dt} = -\omega^2 x_0 \cos(\omega t)$

Note: Use the GC to sketch the graph to see the general shape.



5 Equations in SHM with respect to displacement *x*

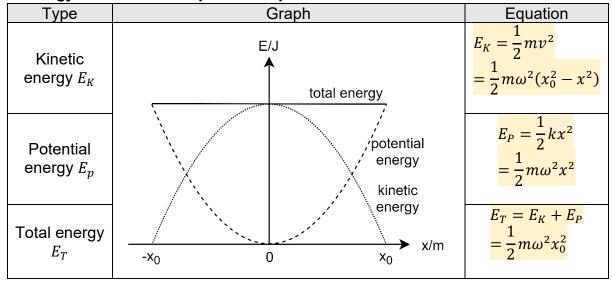
Derivation of $v = \pm \omega \sqrt{x_0^2 - x^2}$ $v = \omega x_0 \cos(\omega t) - -(1)$ $\sin^2(\omega t) + \cos^2(\omega t) = 1$ $\cos(\omega t) = \pm \sqrt{1 - \sin^2(\omega t)} - -(2)$ $x = x_0 \sin(\omega t)$ $\sin(\omega t) = \frac{x}{x_0} - -(3)$

Sub (2) in (1):

$$v = \pm \omega x_0 \sqrt{1 - \sin^2(\omega t)} - -(4)$$

Sub (3) in (4):

$$v = \pm \omega x_0 \sqrt{1 - \left(\frac{x}{x_0}\right)^2} = \pm \omega \sqrt{x_0^2 - x^2}$$



6 Energy in SHM with respect to displacement *x*

7 Energy in SHM with respect to time t (at t = 0, particle is at maximum displacement)

Туре	Graph	Equation
Kinetic energy E_K	E/J ↑	$E_{K} = \frac{1}{2}m\omega^{2}(x_{0}^{2} - x^{2})$ $= \frac{1}{2}m\omega^{2}x_{0}^{2}\sin^{2}(\omega t)$
Potential energy <i>E_P</i>	E _{total}	$E_P = \frac{1}{2}m\omega^2 x^2$ $= \frac{1}{2}m\omega^2 x_0^2 \cos^2(\omega t)$
Total energy E_T	0 Epotential T/2 t/s	$E_T = E_K + E_P$ $= \frac{1}{2}m\omega^2 x_0^2$

Derivation of $E_K = \frac{1}{2}m\omega^2 x_0^2 \sin^2(\omega t)$

$$x = x_0 \cos(\omega t)$$

$$x^2 = x_0^2 \cos^2(\omega t)$$

$$x^2 = x_0^2 (1 - \sin^2(\omega t)) = x_0^2 - x_0^2 \sin^2(\omega t) - -(1)$$

$$E_K = \frac{1}{2}m\omega^2(x_0^2 - x^2) - -(2)$$

Sub (1) in (2):

$$E_{K} = \frac{1}{2}m\omega^{2}[x_{0}^{2} - (x_{0}^{2} - x_{0}^{2}\sin^{2}(\omega t))]$$
$$E_{K} = \frac{1}{2}m\omega^{2}x_{0}^{2}\sin^{2}(\omega t)$$

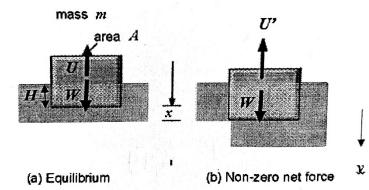
7 Period of a spring-mass system

Comparing:

$$a = -\frac{m}{m}x$$
$$a = -\omega^{2}x$$
$$\omega^{2} = \frac{k}{m}$$
$$T = 2\pi \sqrt{\frac{m}{k}}$$

k

8 Period of a rectangular block oscillating in SHM in water (floating on surface)



Let h be the height of the block.

$$W = U$$

 $mg = \rho A H g - -(1)$

Let the block move downwards by an additional distance *x*

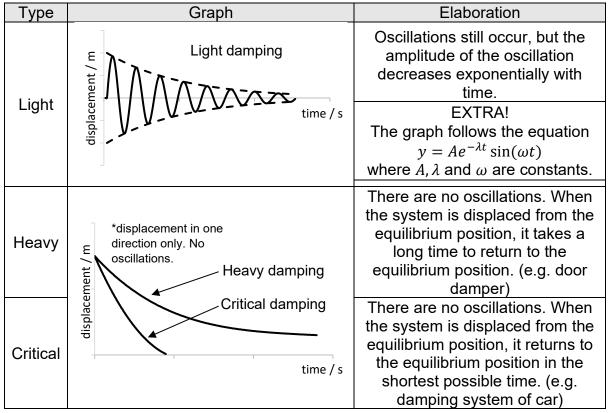
$$W - U' = ma$$
$$mg - \rho A(H + x)g = ma$$
$$mg - \rho AHg - \rho Axg = ma$$
$$-\rho Axg = ma$$
$$a = -\frac{\rho Ag}{m}x$$
$$a = -\omega^2 x$$

Comparing:

$$\omega^{2} = \frac{\rho Ag}{m}$$
$$T = 2\pi \sqrt{\frac{m}{\rho Ag}}$$

9 Damped oscillations

Oscillations gradually die away due to the effects of a damping force acting on the oscillator.



Importance of critical damping: Car suspension system

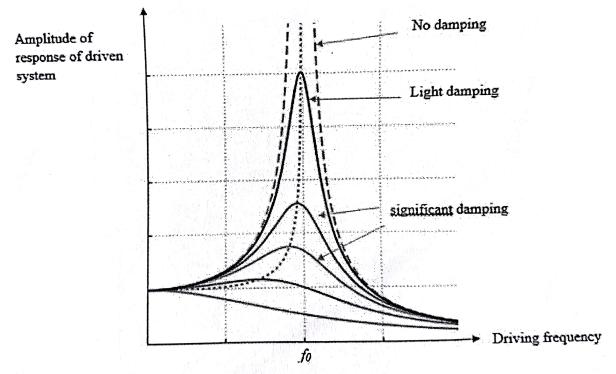
10 Forced oscillations and resonance

Forced oscillations occur when a system is subjected to an <u>external force</u>. Instead of oscillating at its natural frequency, the system will oscillate at the frequency of the <u>external force (driving frequency)</u>.

When the driving frequency is very close to the natural frequency of a system, the system will oscillate at maximum amplitude. This phenomenon is known as <u>resonance</u>, where there is maximum transfer of energy from the driving system to the driven system.

Example: A singer shattering a wine glass, Barton's pendulum

11 Frequency response



Notes:

The amplitude of response is theoretically infinite in the absence of any damping. With damping, the amplitude of the system (and thus the total energy since $E = \frac{1}{2}m\omega^2 x_0^2$) will rise till a point where the rate of work done on the system = the rate of work done against air resistance.

i.e. rate of energy input = rate of energy loss. This can be due an increase in the resistive forces at higher speeds. ($F_{air} \propto v^2$)

With light damping, the amplitude of response is high within a narrow range of driving frequencies close to the natural frequency. The response is described as <u>sharp</u>.

With significant damping, the amplitude of response is smaller, and takes place over a wider range of driving frequencies. The response is no longer sharp.

The frequency of response at resonance also shifts slightly to the left of f_0 in the case of increased damping.

Useful	Destructive	
Production of sound in many musical instruments, especially wind instruments	Collapse of buildings which resonate to seismic waves in an earthquake	
Cooking of food in a microwave oven	Collapse of bridges resonating to strong gusts of wind	
Magnetic resonance imaging (MRI)		

12 Examples of resonance

Chapter 12: Wave Motion

1 Types of waves

Transverse waves: e.g. wave on rope		The direction of vibration of the wave particles is <u>perpendicular</u> to the direction of propagation of the wave.	
Longitudinal waves: e.g. sound wave		The direction of vibration of the wave particles is <u>parallel</u> to the direction of propagation of the wave	
Progressive waves:		The wave profile moves in the direction of propagation of the wave. Progressive waves transfer energy in the direction of the wave velocity.	
2 Key terms Wavelength: The distance between 2 consecutive particles in the wave w		distance between 2 consecutive particles in the wave who in phase with one another. Given by λ	
Frequency:			
r requeitcy.		number of complete oscillations undergone by a wave icle per unit time. Given by <i>f</i>	
Period:	Period: The time taken for a wave particle to undergo one complet oscillation. Given by T		
Amplitude:	ude: The maximum displacement that a wave particle can have awa from its equilibrium position. Given by <i>A</i>		
		velocity of the advancing wave. Note: for EM waves – 10^8 ms^{-1} in vacuum, for sound waves 330 ms^{-1} in air. In by v	

3 Derivation of $v = f\lambda$

Distance travelled by wave in 1 period = λ Time taken = T (1 period) since speed = $\frac{\text{distance travelled}}{\text{time taken}}$ $v = \frac{\lambda}{T} = \frac{\lambda}{\frac{1}{f}}$ $\therefore v = f\lambda$

4 Mechanical waves

Mechanical waves are types of waves that <u>transfer energy through a medium</u>. Both transverse and longitudinal waves are mechanical waves.

E.g.: Water waves, seismic waves

5 Wave equations

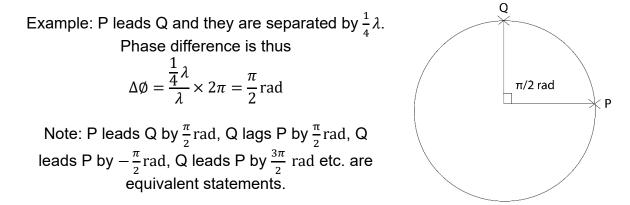
	Displacement-distance	Displacement-time
Equation	$y = A \sin\left(\frac{x}{\lambda} \times 2\pi\right)$	$y = -A\sin(\omega t)$
Graph (transverse)	displacement, y Direction of wave velocity A particle M Distance, x -A	displacement of particle M, y
Graph (longitudinal)	1 2 3 4 5 6 7 8 9 y A B B B B B B B B B B B B B B B B B B	-A

6 Phase difference

For a given displacement-distance graph:

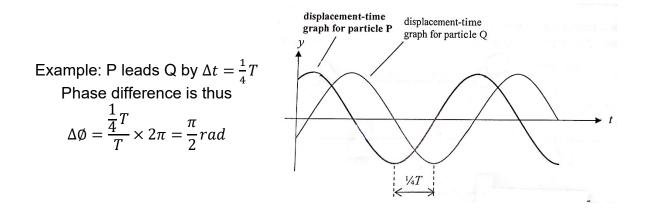
phase difference (rad) =
$$\Delta \phi = \frac{\Delta x}{\lambda} \times 2\pi$$

Note: this formula does not apply to stationary waves. See CH13 Superposition



For 2 different displacement-time graphs of 2 particles:

phase difference (rad) =
$$\Delta \phi = \frac{\Delta t}{T} \times 2\pi$$



7 Intensity (symbol *I*, units Wm^{-2}) **and Amplitude** (symbol *A*, units m) Intensity is the energy transmitted per unit time across a unit area of a surface perpendicular to the direction of the energy flow.

Intensity =	power	_	energy
mensity –	area	_	time × area

Hence, when the power of the source is kept constant, the relation between intensity and distance from the source follows the inverse square law.

Given uniform spreading of wave, then the intensity at a distance r is:

$$I = \frac{P}{A} = \frac{P}{4\pi r^2} \therefore I \propto \frac{1}{r^2}$$

The relation between intensity and amplitude is as such:

Intensity \propto (Amplitude)²

EXTRA!

$$I = \frac{E}{t \times \text{Area}} = \frac{\frac{1}{2}m\omega^2 A^2}{t \times \text{Area}} = (\frac{\frac{1}{2}m\omega^2}{t \times \text{Area}})A^2 \therefore I \propto A^2$$

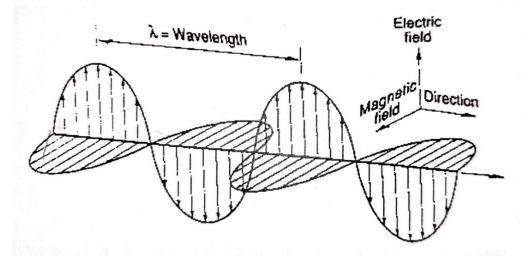
Hence, assuming <u>no energy is lost during propagation</u>, the amplitude of a wave decreases with increasing distance from the source, due to the spreading of waves (amplitude of original source remaining constant). For example, the light from a star decreases with increasing distance.

8 Polarisation

Polarisation occurs with all transverse waves. Longitudinal waves cannot be polarised.

Polarisation is usually applied to electromagnetic waves, e.g. light. Electromagnetic waves consist of an electric field and a magnetic field which vibrate in phase in directions perpendicular to each other.

We usually only keep track of the direction of the oscillation of the electric field vector only. Hence, the direction of the e-field vector is the direction of polarisation of the wave.



View of a vertically polarized EM wave with magnetic and electric components

9 Polarisation calculations

When an <u>unpolarised</u> EM wave is incident on a polariser, its intensity will be reduced by $\frac{1}{2}$. Hence,

$$I = \frac{1}{2}I_0$$

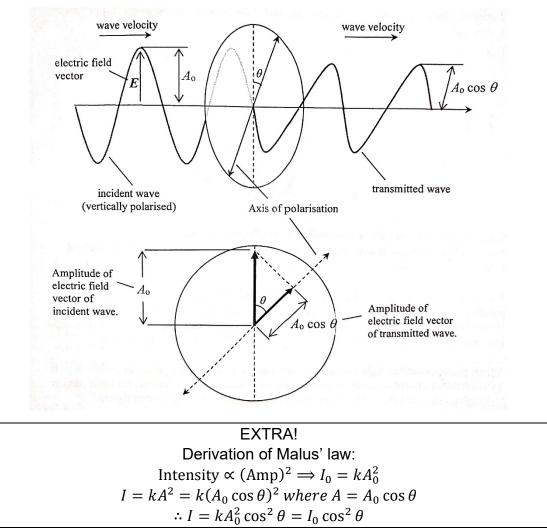
This means that 50% of the wave's energy passes through the polariser (energy \propto intensity as derived above).

When a <u>polarised</u> EM wave is incident on a polariser, its intensity can be found by the formula:

$$I = I_0 \cos^2 \theta$$

where θ is the angle between the direction of polarisation of the incident wave and the polarizing axis.

This is known as Malus' Law.



Chapter 13: Superposition

1 Principle of superposition

The principle of superposition states that the resultant wave displacement (and hence the amplitude) is given by the vectorial sum of the individual wave displacements (amplitude) at the point where the waves meet.

According to the principle of superposition, when a wave meets a <u>fixed</u> end, the reflected wave will be <u>inverted</u>. When a wave meets a <u>free</u> end, the reflected wave will <u>not be inverted</u>.

Strategy: Extend the wave past the end. Mirror the wave from the right onto the left (reflection in the y-axis). If it is a fixed end, invert the wave (reflection in the x-axis).

Note: Waves which are polarised perpendicular to one another will <u>not</u> produce observable interference patterns on a screen.

They obey the Principle of Superposition, but there will be no destructive interference, hence there will be no observable bright and dark fringes. These waves can still interfere with each other.

Common examples of end types:

Free end	Fixed end	
End of a rope	End of a rope attached to a wall	
Wall of water tank		

2 Stationary / Standing waves

A standing or stationary wave is formed when two waves of the same propagation speed, wavelength and amplitude overlap while travelling in opposite directions.

A standing or stationary wave is one which the wave profile does not travel in the direction of the wave velocity, though the wave particles still execute oscillatory motion (SHM) about their rest positions. Hence, a stationary wave does not transport any energy. Energy is trapped within the stationary wave.

Particles on the wave oscillate with different amplitudes vertically.

Nodes are a point at which the displacement of a particle is permanently zero. Antinodes are a point with maximum amplitude.

Particles in a N-A-N segment are in phase with one another. They are 180° out of phase with the adjacent N-A-N segment.

_	Progressive wave	Stationary wave
<u>F</u> requency	Same for all points	Same except at the nodes $(f = 0)$
<u>E</u> nergy	Transfers from one point to	No transmission of energy along
	another	the wave
Wavelength	Distance between two	Twice the distance between a pair
<u>w</u> avelength	consecutive points in phase	of adjacent nodes or antinodes
<u>P</u> hase	Different phases in 1 wavelength	All particles in the same segment are vibrating in phase. Particles in alternate segments are 180° out of phase.
<u>A</u> mplitude	Same for all points	Varies; largest at antinode, 0 at the node
<u>W</u> aveform	Travels in direction of wave	Does not advance

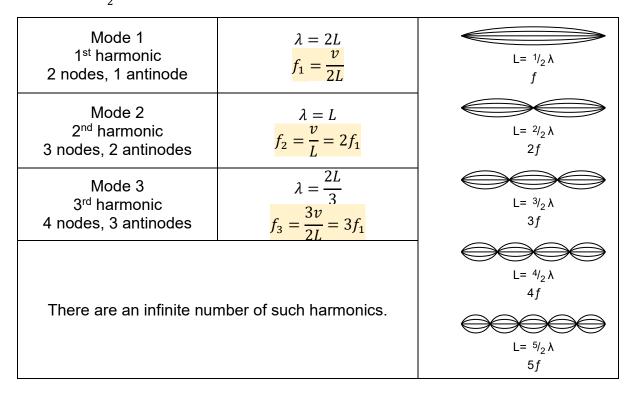
Characteristics of progressive and stationary waves:

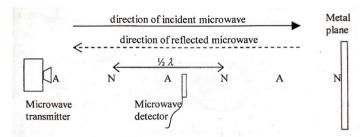
Method to remember: **FEW PAW**

3 Transverse stationary waves

Supposed a string is stretched tightly between fixed supports. It is then able to vibrate in a direction perpendicular to its length if it is plucked. Waves travel away from the position it is plucked, reach the fixed ends, get reflected and travel back along the string, forming stationary waves.

As the ends are fixed, they are nodes. The distance between adjacent nodes / antinodes is $\frac{1}{2}\lambda$.





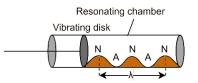
Experiment: finding wavelength of microwave using a detector

The detector will read 0 at a displacement antinode or pressure node. The distance between nodes can be calculated, and from this the wavelength through $v = f\lambda$ The same setup can be used for other EM waves.

4 Longitudinal stationary waves

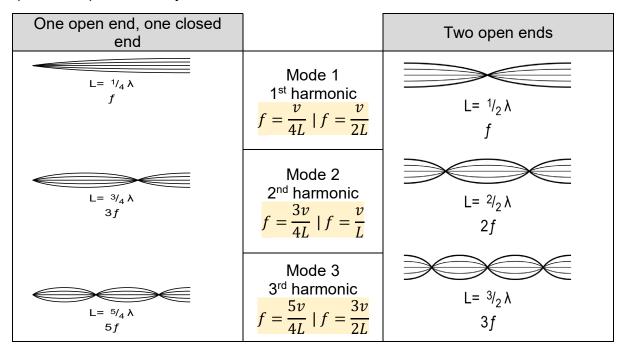
Stationary waves can be formed in either a pipe that has 2 closed ends or 1 open end.

Kundt's' tube - experiment to demonstrate longitudinal standing waves



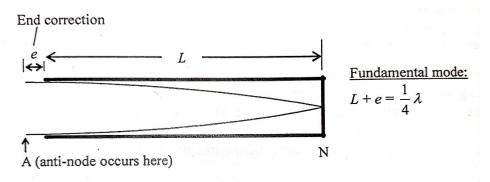
At the nodes, the air particles are not moving. Hence, they allow powder to collect at the nodes.

At the antinodes, the air particles are moving with maximum amplitude. Hence, they "push" the powder away from the antinodes to the nodes.



Strategy for calculations: find the wavelength in terms of *L*.

End correction: In reality, the antinode occurs at a short distance – called the end correction – away from the open end.

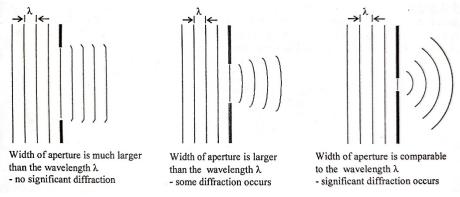


The wavelength is now $\lambda = 4(L + e)$.

5 Diffraction

Diffraction refers to the spreading of waves when they encounter an aperture or obstacle whose linear dimension (e.g. width of aperture) is comparable to the wavelength of the waves.

Generally, the smaller the width of the aperture, the greater the spreading or diffraction of the waves.



Wavelength remains unchanged before and after diffraction.

6 Two-source interference

Interference refers to the phenomenon of two or more waves of the same type meeting at a point in space to produce a resultant wave disturbance given by the superposition of individual waves at that point.

There are 3 conditions needed for steady (constant) observable interference patterns – The 2 wave sources must be coherent, must be unpolarised or polarised in the same plane, and must have roughly the same amplitude.

Condition		Notes		
Both waves must be coherent		Sources of waves are said to be coherent if they have a constant phase difference. This implies that they have the same wavelength or frequency and maintain a constant phase relation with one another.		
Both waves must be unpolarised or polarised in the same plane		Waves polarised perpendicular to one another will not produce any observable interference pattern, just a bright fringe. This is because there is no destructive interference.		
Both waves must have roughly the same amplitude		If one wave has a larger amplitude than the other wave, when destructive interference occurs ($A_{net} = A_1 - A_2$), the net amplitude will not be 0, resulting in no dark fringes.		
Constructive int		e interference	Destructive interference	
Description	Resultant amplitude is a maximum. It occurs at a point where the waves meet in phase; they reinforce each other. The amplitude of the resultant wave at a given time or position is greater than that of either individual wave.		Resultant amplitude is 0 if the 2 waves are identical. It occurs at a point where they meet exactly out of phase. The amplitude of the resultant wave at a given time or position is smaller than that of either individual wave.	
Condition	ondition 2 coherent waves meet in phase		2 coherent waves meet out of phase	
Doth		$\Lambda x - x^{2}$	$\Delta u = \begin{pmatrix} n & 1 \\ n & 1 \end{pmatrix} 1$	

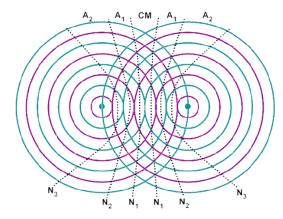
mula	Path difference	$\Delta x = n\lambda$ where $n \subseteq \mathbb{Z}^+$	$\Delta x = \left(n + \frac{1}{2}\right)\lambda$ where $n \subseteq \mathbb{Z}^+$
Fo	Phase	$\Delta \phi = 2n\pi$	$\Delta \phi = (2n+1)\pi$
	difference	where $n \subseteq \mathbb{Z}^+$	where $n \subseteq \mathbb{Z}^+$

Note: the above formula holds true when the 2 sources are in phase. If the 2 sources have a phase difference of 180° or π , then the equations will swap around. e.g. constructive interference if $\Delta x = \left(n + \frac{1}{2}\right)\lambda$, where $n \subseteq \mathbb{Z}^+$

Example: identifying the nodal lines (drawn) and antinodal lines (not drawn) in two-source interference.

Path difference at $A_1 = 0$, $A_2 = \lambda$...

Path difference at N₁ = $\frac{\lambda}{2}$, N₂ = $\frac{3\lambda}{2}$...



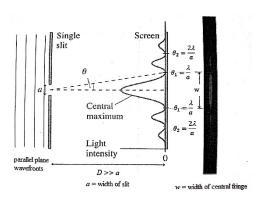
7 Single Slit diffraction

In a slit, each portion of the slit acts as a 'source' of light waves.

The intensity of each maxima decreases as the angle of diffraction of light increase.

Equation for single slit is (details not in syllabus)

$$\sin\theta = \frac{n\lambda}{a}$$



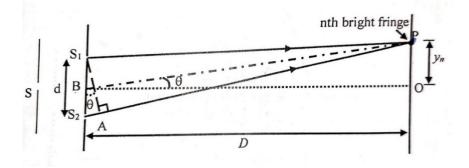
For small θ , sin $\theta \approx \theta$ for n = 1, 2

Width of slit (using prior trigonometry knowledge): $\tan \theta = \frac{w}{2D}$

Two-source interference on single slit: two sources must be at a distance D from the slit in order to be just resolved (the central maximum of one image falls on the first minimum of the other). This is known as Rayleigh's criterion.

$$\theta_{min} = \frac{\lambda}{a}$$

8 Double Slit diffraction



Note: source S needs to be exactly the same distance from S_1 and S_2 . Note: If light from S_1 and S_2 are polarised 90° to one another, they will not interfere. A bright band will be observed on the screen.

$$\Delta y = \frac{\lambda D}{d}$$

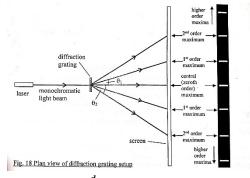
EXTRA!
Path difference =
$$S_2A = n\lambda$$
.
 $\therefore \sin \theta = \frac{n\lambda}{d}$ (basic equation for slit diffraction)
 $\tan \theta = \frac{y_n}{D}$ and $\tan \theta \approx \sin \theta$ for small θ
 $\frac{n\lambda}{d} = \frac{y_n}{D} \Rightarrow y_n = \frac{n\lambda D}{d}$
 $\therefore \Delta y = y_{n+1} - y_n = \frac{(n+1)\lambda D}{d} - \frac{n\lambda D}{d}$
 $\therefore \Delta y = \frac{\lambda D}{d}$

9 Multiple slit diffraction or diffraction grating

Diffraction gratings are useful for analysing light sources. Fringes formed by the grating is sharper than the double slits because of the enormously large number of slits in the grating.

Formula:

$$d \sin \theta = n\lambda$$



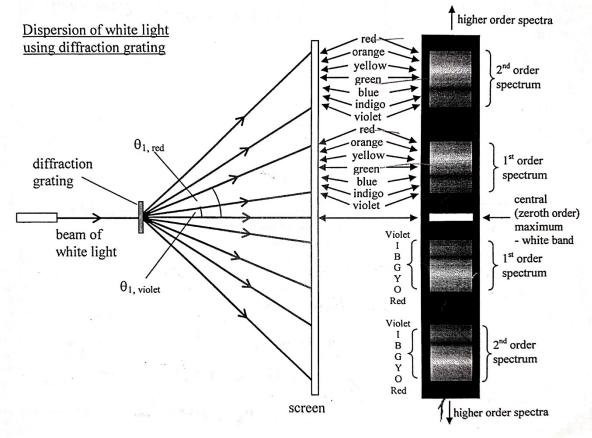
Maximum number of orders (maximums) visible is given by $n_{max} \leq \frac{d}{\lambda}$. This is because the maximum value of $\sin \theta$ is 1.

If white light is shone through the grating, then each wavelength in white light will be diffracted slightly differently. This results in a spectrum of colours to be seen on the screen, with one spectrum for each order n.

Red light will diffract more than violet light as the wavelength of red light is more than that of violet light.

It is possible for two spectra of colours to overlap (i.e. violet of 1st order and red of 2nd order may overlap).

At the straight-through position, there is no diffraction, and hence a white band is observed.



Chapter 14: Current of Electricity

1 Electric current (symbol *I*, units A or Cs^{-1})

The current *I* is the rate of flow of electric charge Q through a given cross-section of a conductor.

$$I = \frac{dQ}{dt}$$
 for varying current,
$$I = \frac{Q}{t}$$
 for steady current

Direction of conventional current is taken to be the flow of positive charges. Direction of conventional current can be thought of as opposite to the flow of negative charges.

Therefore, net charge flow in a direction is the total charge of positive charges moving in that direction plus the total charge of negative charges moving in the opposite direction.

Flow of positive charges	Flow of negative charges	Net charge flow to the right
5C to the right	5C to the right	5C - 5C = 0C
5C to the right	5C to the left	5C + 5C = 10C

2 Charge (symbol Q, units C or coulumb)

The charge Q which flows past a given cross section is the product of the steady current *I* that flows past the section and the time during which the current flows. Q = It

The coulomb *C* is the amount of electrical charge that passes through a given cross section of a circuit when a steady current of one ampere flows in one second.

An electron or proton carries an electric charge of approximate magnitude 1.6×10^{-19} C.

One coulomb is approximately 6.24×10^{18} electrons or protons.

Total charge is given by Q = Ne.

Given the charge, the number of electrons can be calculated:

$$N = \frac{Q}{e} = \frac{1C}{1.6 \times 10^{-19}}$$

The charge density *n* is the number of charge carriers per unit length, surface area, or volume in a conductor.

*We usually consider the volume charge density.

Assume charge flows through a conducting wire of length l and cross-sectional area A.

Volume of this area $\Delta V = A\Delta x$ and number of charge carriers is $N = n\Delta V$ Therefore $N = nA\Delta x$. Total charge in this section $Q = Nq = nA\Delta xq$. Hence current $I = \frac{nA\Delta xq}{\Delta t}$ Since drift velocity $v_d = \frac{\Delta x}{\Delta t}$, $I = nAv_d q$

3 Potential difference (symbol *V*, units V or JC⁻¹)

The potential difference between two points in a circuit is the amount of electrical energy converted to other forms of energy per unit charge passing from one point to another.

$$V = \frac{W}{Q}$$

The volt is the potential difference between two points in a circuit if one joule of electrical energy is converted to other forms of energy per coulomb of charge passing from one point to the other.

4 Electromotive force (symbol ε , units V)

The electromotive force (e.m.f.) of any source of electrical energy is the total energy converted into electrical energy per unit charge supplied.

$$E = \frac{W}{Q}$$

Differences between e.m.f. and p.d.			
e.m.f. p.d.			
the energy converted from non-	the energy converted from electrical to		
electrical to electrical per unit charge	non-electrical per unit charge passing		
delivered by the source of e.m.f.	from one point to another		

5 Resistance (symbol *R*, units Ω or VA⁻¹)

Resistance of a conductor is the ratio of potential difference across the conductor to the current flowing through it.

$$R = \frac{V}{I}$$

The ohm is defined as the resistance of a conductor such that a current of 1 ampere flows in it when a p.d. of 1 volt is applied across it.

Ohm's law: The current flowing through a conductor is directly proportional to the p.d. across it, providing temperature and other physical conditions are constant.

 $I \propto V$ for constant temperature and physical conditions

Devices which obey Ohm's law (i.e. have constant resistance at a certain temperature) are called ohmic conductors (e.g. pure metals). The I - V or V - I graph will give a straight line through the origin.

Resistance will increase when temperature increases. This is because of the increase in amplitude of the vibration of the atoms, causing electrons to collide more frequently with them.

Devices which do not obey Ohm's law are called non-ohmic conductors. The I - V or V - I graph does not give a straight line.

	Examples of non-ohmic conductors				
Type	Filament lamp	NTC (negative temperature coefficient) thermistor	Semiconductor diode		
Graph		I 0 V	Forward bias I Reverse bias 0 V		
Characteristics	Temperature $\approx 2000K$ Number of charge carriers stays relatively constant <i>R</i> increases as temperature increases due to increase in collisions of charge carriers	Resistance is sensitive to temperature. Number of charge carriers increases markedly with temperature. <i>R</i> decreases greatly as temperature increases.	Made up of 2 semiconductors swhich conduct well in 1 direction but not the other.		
Special note for semiconductor diode: In reality, the $I - V$ graph for the diode will look like this. There is a minimum forward bias p.d. for which a current will flow called the <u>threshold voltage</u> . If the reverse p.d. is too large, larger than the <u>breakdown voltage</u> , the diode may break down resulting in a large "avalanche" current (short circuit). Note: The expression $R = \frac{V}{I}$ is not a representation of Ohm's law, rather, it is a representation of the resistance					

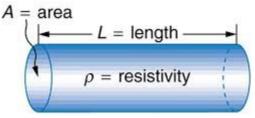
R. To find the resistance at a point from a V-I graph, take the gradient of the lone connecting that point and the origin. Take the inverse gradient for an I-V graph.

9 Resistivity (symbol ρ , units Ω m)

Resistivity is a property of the material, independent of the shape and size of the conductor.

$$R = \frac{\rho L}{A}$$

Definition: Resistivity is given by $R = \frac{\rho L}{A}$, where *R* is the resistance of the component, *A* is its cross-sectional area, and *L* is its length.

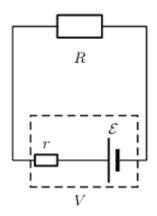


10 Internal resistance

An ideal battery should produce a voltage (termed "terminal voltage" – p.d. at the terminals) that is equal to the e.m.f.. However, in reality, all batteries have a small internal resistance r that causes the terminal voltage to be less than the stated e.m.f.

When connecting a voltmeter across the ends of the battery, we see that the voltage V is less than the e.m.f. ε . This is because the drop in voltage is used to drive the current through the internal resistance r.

Therefore drop in voltage = voltage across resistor and $r = \frac{V_r}{I} \Longrightarrow V_r = Ir$. Since $V_{terminal} = \varepsilon - V_r$, therefore $V_{terminal} = \varepsilon - Ir$



and $\varepsilon = I(R + r)$ where *R* is the total external resistance.

Note: p.d. across total external resistance is equal to the terminal p.d.. Hence

$$IR_{ext} = \varepsilon - Ir$$

11 Power (symbol P, units Js^{-1} or W)

Whenever a current passes through a resistance, power is transformed from one type to another.

From the equations of p.d. and e.m.f., the following equations can be obtained.

$$P = VI = I^2 R = \frac{V^2}{R}$$

 $P = I^2 R$ is usually used for <u>series</u> circuits. $P = \frac{V^2}{R}$ is usually used for <u>parallel</u> circuits.

Chapter 15: Electric Fields

In the case of electrons, gravity is assumed to be negligible, as $m_e \ll e$. However, for particles that are made up of at least a few molecules (e.g. oil drop), gravity cannot be assumed to be negligible.

1 Electric field

An electric field is a region of space where a <u>charged</u> particle will experience an <u>electric</u> force. (refer to gravitational field notes)

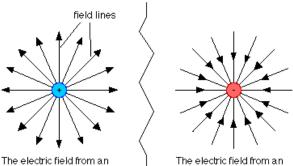
It can be represented using lines of force. starting from a positive charge and ending on a negative charge.

When drawing electric fields, take note of:

a) Electric field lines must begin from a positive charge and end on a negative charge

b) Electric field lines must not intersect each other

c) Electric field lines must be at a normal to the surface of the charge



The electric field from an isolated positive charge

The electric field from an isolated negative charge

d) The number of lines per unit area is proportional to the strength of the electric field in the given region.

On an irregularly shaped conductor, the charge concentrates at locations where the radius of curvature is the smallest – at pointed regions.

2 Electric field strength (general) (symbol E, units NC⁻¹)

The electric field strength E at a point is the electric force acting per unit positive charge placed at that point.

$$E = \frac{F}{+q}$$

E is a vector whose direction at a point is the direction of the electric force that would be exerted on a small positive test charge placed at that point.

Hence, a small positive test charged placed (u = 0) along a field line will travel along that field line.

Note: if the initial velocity is not 0, then it will not exhibit such behaviour.

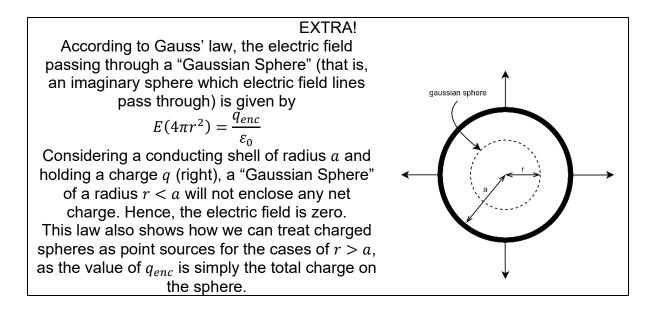
A charged particle entering a uniform electric field perpendicular to it will follow a parabolic path.

Proof: $u_x = u$ and $u_y = 0$ Hence $s_x = u_x t = ut$ and $s_y = u_y t + \frac{1}{2}at^2 = \frac{1}{2}at^2$ Since F = ma = Eq, $a = \frac{Eq}{m}$ Therefore $s_y = \frac{1}{2} \left(\frac{Eq}{m}\right) \left(\frac{1}{u_x^2}\right) s_x^2 \implies y = kx^2$ The electric field strength at a point P due to a point charge at distance r is given by

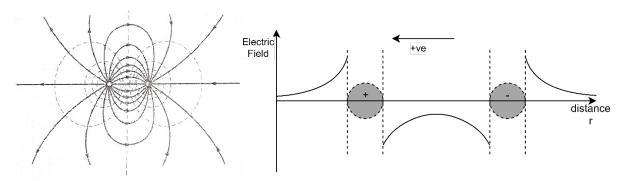
$$E = \frac{Q}{4\pi\epsilon_0 r^2}$$

This formula is also applicable to a sphere of radius R. The distance r will then be the distance of the point to the centre of the sphere.

Note that there is no electric field inside an isolated electrically charged sphere. Any excess charge on an isolated conductor resides entirely on its surface.



To draw the electric field-distance graph for 2 objects, first define the sign convention, e.g. $\rightarrow +ve$. Then, draw the strength of the e-field according to the line density.



Left: visualisation of the electric field lines. Right: Electric field vs distance graph

3 Coulomb's Law

To calculate the force acting on a charged particle due to another charged particle, we can use Coulomb's Law.

Coulomb's Law states that the electric force between two point particles carrying charges Q_1 and Q_2 separated by a distance r apart is proportional to the product of the two charges and inversely proportional to the square of their distance apart.

$$F \propto \frac{Q_1 Q_2}{r^2} \qquad \varepsilon = 8.85 \times 10^{-12} \,\mathrm{Fm}^{-1}$$
$$F = \frac{Q_1 Q_2}{4\pi\varepsilon_0 r^2}$$

Coulomb's Law also applies to charged spherical conductors. The distance r will then be measured from the centre of the sphere.

4 Electric Potential Energy (symbol U, units J)

The electric potential energy (EPE) between two point charges Q_1 and Q_2 separated by a distance r apart is given by

$$U = \frac{Q_1 Q_2}{4\pi\varepsilon_0 r}$$

EPE can be negative, if one charge is positive and the other is negative.

Note: The work done <u>on</u> a particle by an external agent is equal to the negative of the work done <u>by</u> the electric field.

5 Electric Potential (symbol *V*, units JC^{-1})

The electric potential at a point in an electric field is the work done per unit positive charge by an external agent in bringing a small test charge from infinity to that point without a change in kinetic energy.

Note: the electric potential at infinity is taken to be zero by convention (see gravitational field notes).

$$V_c = \frac{W}{q} = \frac{Q}{4\pi\epsilon_0 r}$$

Electric potential can be negative, if the charge Q is negative.

Potential energy U and potential V is related by the equation

$$J = Vq$$

Positive charges will move from places of high potential to places of lower potential. Negative charges will move from places of low potential to places of higher potential.

Change in EPE by a particle moving between 2 points whose potential difference is ΔV is given by the formula

$\Delta U = q \Delta V$

Charges moving along an equipotential line will not experience a change in potential energy.

Note: while the electric field strength inside a charged sphere of radius *R* is 0, the potential energy in a charged sphere of radius *R* is not 0, it is $\frac{Q}{4\pi\varepsilon_0 R}$

6 Relationship between electric field quantities

Generally, the electric field strength at a point is equal to the negative of the potential gradient at that point.

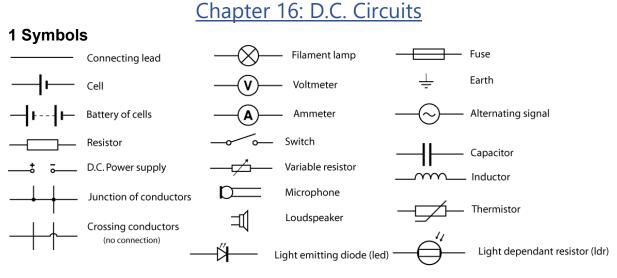
$$E = -\frac{dV}{dr}$$

Generally, the force acting on a charge is numerically equal to the negative of the potential energy gradient at that point.

$$F = -\frac{dU}{dr}$$

7 Summary

$F = \frac{Qq}{4\pi\varepsilon_0 r^2}$	$\Leftarrow E = \frac{F}{q} \Longrightarrow$	$E = \frac{Q}{4\pi\varepsilon_0 r^2}$
ſ		ſ
$F = -\frac{dU}{dt}$		$E = -\frac{dV}{dV}$
ar ↓		ar ↓
$U = \frac{Qq}{4\pi\varepsilon_0 r}$		$V = \frac{Q}{4\pi\varepsilon_0 r}$



Common devices:

Thermistor:	Resistance decreases as temperature increases
rnormstor.	

Light-dependent resistor (LDR):	Resistance decreases as light intensity increases
Diode:	Allows electrical current only in one direction

Common assumptions made:

Ammeter:	Zero resistance
Galvanometer:	Zero resistance
Voltmeter:	Infinite resistance

2 Current

Kirchhoff's Current Law states that the algebraic sum of the currents at a junction of a circuit is zero.

$$\sum_{junction} I_i = 0$$

where currents <u>entering</u> the junction are given a <u>positive</u> (+) sign and currents <u>leaving</u> the junction are given a <u>negative</u> (-) sign.

Generally

sum of currents entering a junction = sum of currents leaving the junction

This law arises from the Principle of Conservation of Charge.

3 Voltage

<u>Electromotive force</u> (symbol ε) of a cell is the amount of energy converted from nonelectrical to electrical form per unit charge that moves from one terminal of the cell to the other outside of the cell.

$$\varepsilon = \frac{\Delta W}{q}$$

<u>Potential difference</u> (symbol *V*) across a circuit component is the amount of energy converted from electrical to non-electrical forms per unit charge that flows through the component.

$$\Delta V = \frac{\Delta W}{q}$$

Note: the <u>negative</u> terminal of a battery can still have a <u>positive</u> potential. Negative terminal just means that it is at a lower potential, compared to the positive terminal.

Ohm's law: For a resistor obeying Ohm's law, this equation holds true:

i

$$\Delta V = IR$$

Kirchhoff's Voltage Law states that the algebraic sum of all electrical potential changes around any closed loop is zero.

$$\sum_{unction} \Delta V_i = 0$$

This law arises from the Law of Conservation of Energy.

Note: current always flows from higher potential to lower potential.

4 Resistors

For resistors in <u>series</u>, the total resistance is simply *the sum of the resistances*.

$$R_{eff} = R_1 + R_2 + R_3 + \dots + R_n$$

For resistors in <u>parallel</u>, the reciprocal of the total resistance is the sum of the reciprocals of the individual resistances.

$$\frac{1}{R_{eff}} = \frac{1}{R_1} + \frac{1}{R_2} + \frac{1}{R_3} + \dots + \frac{1}{R_n}$$

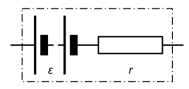
Note: effective resistance of resistors in parallel is always lower than the lowest resistance in the network.

5 Ammeters and Voltmeters

Ammeters / galvanometers should be connected in <u>series</u>. An ideal ammeter / galvanometer has 0 resistance. Voltmeters should be connected in <u>parallel</u>. An ideal voltmeter has infinite resistance.

6 Non-ideal cells

All cells inherently possess an internal resistance r. This is represented with a dotted line as such:



The internal resistance is in <u>series</u> with the battery.

This causes the <u>terminal p.d.</u> to be lower than the e.m.f.

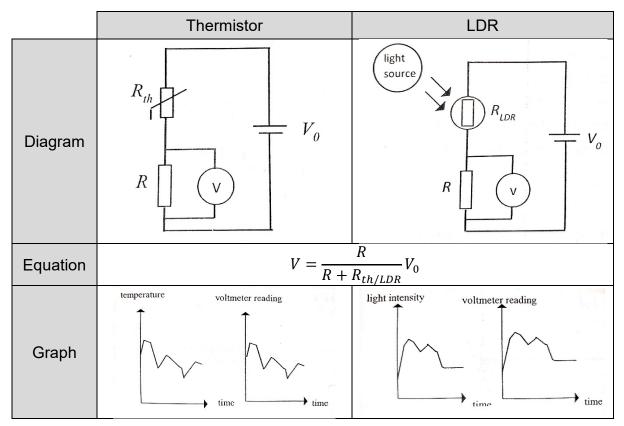
 $V = \varepsilon - Ir$

7 Potential-divider Principle

The <u>Potential-Divider</u> rule states that if a voltage exists across several resistors connected in <u>series</u>, then the voltage across each resistor is proportional to the total resistance.

$$V_1 = \frac{R_1}{R_1 + R_2} * V = \frac{R_1}{R_T} * V$$

When used in a circuit in series with a <u>thermistor</u> or <u>LDR</u>, the following graph can be obtained.



The potential-divider principle can also be used for a <u>continuous resistor</u>, e.g. a resistance wire.

If a wire has a <u>constant resistance per metre</u>, then the <u>voltage drop</u> across any section of the wire will be <u>proportional</u> to the length of the section as a fraction to the wire's total length.

$$V_{AB} = \frac{l_{AB}}{l_T} * V_0$$

EXTRA!

If a wire has non-uniform resistance per metre, integrating over the length of the wire will give the total resistance of the wire.

e.g. the resistance per metre of wire *X*, which is *L* cm long, is given by r = A + Bx. Hence total resistance is $\int_0^L A + Bx \, dx$

8 Potentiometer

The potentiometer is essentially a device used to <u>measure an unknown potential</u> <u>difference</u>. The unknown voltage is compared with a fraction of a voltage from a known source across a resistance wire.

The <u>balance length</u> is the point when the galvanometer shows <u>no current</u>.

This means that the potential difference across PJ is the same as the unknown e.m.f. i.e. $V_1 = V_2$. Hence, no current flows.

The unknown voltage E can be calculated by

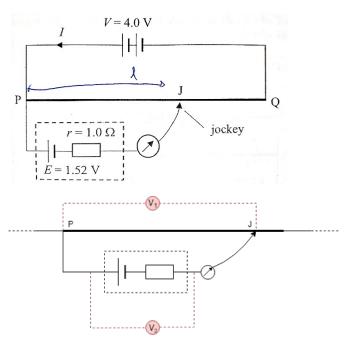
$$E = \frac{L_{PJ}}{L_{PQ}} * V$$

Note: the potentiometer is very accurate (provided that the e.m.f. source for the main circuit is known accurately)

Note: The presence of any resistors in series with the unknown e.m.f. <u>does not affect</u> <u>the balance length at all</u>.

$$V_{resistor} = IR_{resistor}$$

When I = 0 as there is no current flow, there is <u>no potential drop across the resistor</u>. Hence, the balance length is not affected by resistors.



9 Wheatstone bridge

The Wheatstone Bridge is a way to determine an unknown resistance accurately.

 R_x = unknown resistance R_1 = variable resistance R_2 and R_3 = known resistances (fixed)

Like a galvanometer, the resistance R_1 is adjusted until the <u>balance point</u> (I = 0).

This means that the potential difference between 1) A and B is 0. Hence,

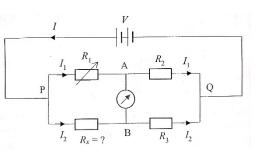
$$V_{PA} = V_{PB}$$
$$I_1 R_1 = I_2 R_x$$

 $V_{AQ} = V_{BQ}$ $I_1 R_2 = I_2 R_3$

Hence,

and

$$\frac{l_1}{l_2} = \frac{R_x}{R_1} = \frac{R_3}{R_2}$$
$$R_x = \frac{R_1}{R_2} \times R_3$$



Chapter 17: Electromagnetism

1 Magnets

Magnets produce magnetic fields. The region of space close to a magnet is called a magnetic field.

The field is represented by <u>lines of force</u>, starting from the North pole and ending at the South pole.

<u>Hard</u> magnetic materials like cobalt and nickel are difficult to magnetise but tend to retain their magnetism.

<u>Soft</u> magnetic materials like iron are easily magnetised but tend to lose their magnetism easily.

2 Magnetic field (symbol *B*, units T)

Magnetic field is a vector. The strength of the magnetic field at a point is represented by the symbol B.

Note: Magnetic field = magnetic flux density (covered later)

The <u>tangent</u> to the magnetic field line at a point in the magnetic field gives the direction of the field at that point.

The <u>number of lines per unit cross section area</u> is an indication of the <u>strength</u> of the field.

Arrow: Magnetic field acting in direction of arrow

Cross: Magnetic field acting into plane of paper

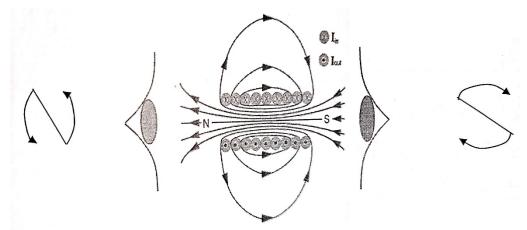
Dot with a circle: Magnetic field acting out of the plane of paper

3 Magnetic field patterns of different magnetic objects

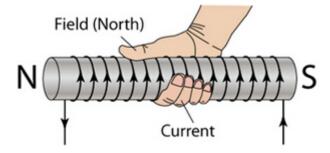
Magnetic field due to				
Current flowing in long straight wire	Current flowing in a long solenoid	Current in a circular coil		
Direction of Current Flow	N OCOCOCOCOCOCOCOCOCOCOCOCOCOCOCOCOCOCOC	Electric current loop current		
$B = \frac{\mu_0 I}{2\pi r}$	$B = \mu_0 n I$	$B = \frac{\mu_0 NI}{2r}$		

4 Solenoid

The polarity of the ends of the solenoid can be determined by the "Clock" rule. Looking at the end of a solenoid, if the current is flowing <u>counter-clockwise</u>, that end is a <u>north pole</u>. If the current is flowing <u>clockwise</u>, that end is a <u>south pole</u>.



The polarity can also be determined by grasping the core, with the fingers curled in the direction of the current. The thumb will then point to the north pole.

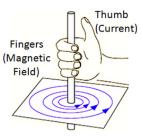


A <u>ferrous core</u> can increase the magnetic field strength of a solenoid very dramatically. This is because of the presence of "domains" in the ferromagnetic material, which are composed of tiny magnets made up of atoms.

In the unmagnetised state, these atomic magnets point randomly in all directions. When subjected to an external magnetic field, the domains start to become <u>aligned</u> in one direction, and the magnetic field becomes very strong.

5 Right hand grip rule

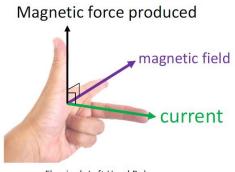
If the wire is grasped in the right hand, with the thumb in the direction of the current, the fingers will curl in the direction of the magnetic field.



6 Fleming's Left Hand rule

Fleming's Left Hand rule can be used to find the direction of the force acting on a currentcarrying conductor in a magnetic field.

The se<u>C</u>ond finger represents the <u>C</u>urrent, the <u>F</u>irst finger represents the magnetic <u>F</u>ield and the thumb represents the magnetic force.



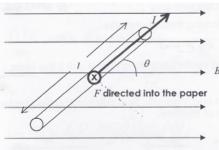
Fleming's Left Hand Rule

7 Magnetic force acting on a current-carrying conductor

The direction of the force can be calculated using Fleming's left hand rule. The <u>magnitude</u> of the force can be calculated using:

 $F = BIL \sin \theta$

where θ is the angle between the magnetic field vector and the direction of the current.



8 Magnetic Flux Density (symbol *B*, units T)

Magnetic flux density is defined as the <u>force acting per unit length</u> of a conductor which carries <u>unit current</u> and is of <u>right angles to the magnetic field</u>.

$$B = \frac{F}{IL\sin\theta}$$

The <u>Tesla</u> is the unit of magnetic flux density equivalent to a force of 1 N experienced by a straight conductor of length 1 m and carrying a current of 1 A when it is placed perpendicular to the magnetic field.

9 Forces acting on a pair of current-carrying wires

If the currents are flowing in the <u>same direction</u> (A), then the wires will <u>attract</u> each other.

If the currents are flowing in the <u>opposite direction</u> (B), then the wires will <u>repel</u> each other.

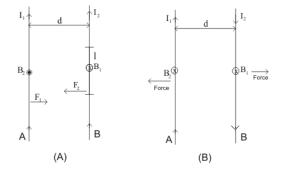
The directions of the forces can be explained using (1) the <u>right hand grip rule</u>, to determine the direction of the magnetic field, followed by <u>Fleming's left hand rule</u>, to determine the direction of the force.

The force of interaction is given by

$$F \propto \frac{I_1 I_2}{d}$$

Derivation: See later:

$$F_1 = BI_1 L = \frac{\mu_0 I_2}{2\pi r} I_1 L$$



$B \xrightarrow{P} Q$ $Current \\ P \\ (into " paper)$ Current "out of" paper F

10 Forces (torque) acting on a current-carrying rectangular loop

The force acting on the vertical arm is F = BIb. Since the forces are a couple, the torque τ experienced is:

$$\tau = Fa = (BIb)a = BIA$$

For a loop of N turns, the torque experienced is:

$$\tau = BIAN$$

A <u>current balance</u> can be used to figure out the strength of a magnetic field. It does this by aiming to balance the (1) force acting on the current-carrying rectangular loop (via electromagnetism) and a (2) known force (from a <u>nonconducting rod</u>) in a see-saw configuration.

The torque acting on side CD (clockwise) is $\tau_{clockwise} = F_B x = (BIL)x$, where *L* is the length of the wire CD The torque acting on side AF (anticlockwise) is

 $\tau_{anticlockwise} = F_{rod}y = mgy$ To achieve rotational equilibrium,

 $\tau_{anticlockwise} = \tau_{clockwise}$

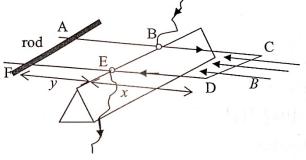
$$BILx = mgy$$

$$B = \frac{mgy}{ILx}$$

11 Force acting on a charged particle moving in a magnetic field

To find out the direction of force acting on a charged particle moving in a magnetic field, first we must resolve the B field into a component parallel and perpendicular to the direction of velocity v.

Note: take the direction of <u>conventional current</u> (i.e. positive charge flow), not electron flow.



Then, using <u>Fleming's Left Hand Rule</u>, the direction of force can be obtained. The magnitude of the force is given by:

$$F = B_{\perp}qv = B\sin\theta \, qv$$

Since the direction of the force is always perpendicular to the direction of motion v, a particle in a large enough region of magnetic field will experience <u>circular motion</u>.

Thus, the magnetic force <u>provides</u> the centripetal acceleration.

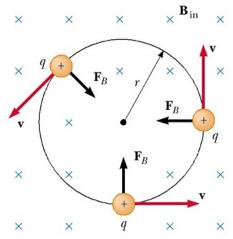
$$F_{c} = F_{B}$$

$$mr\omega^{2} = B_{in}qv$$

$$mr\omega(2\pi f) = B_{in}q(r\omega)$$

$$2\pi fm = B_{in}q$$

$$f = \frac{B_{in}q}{2\pi m}$$



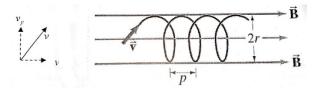
 $\boldsymbol{F} = q(\boldsymbol{v} \times \boldsymbol{B})$

This frequency is called the cyclotron frequency.

The cyclotron frequency and hence the period is independent of the speed of the particle.

If the particle was not moving at right angles to the magnetic field, the particle will still exihibit circular motion, but with an added velocity perpendicular to the circular motion.

As a result, the path of the particle will follow the shape of a <u>spring's coils</u>, and it will move in a helical path.



 v_y causes circular motion about the field lines, while v_x causes the particle to move parallel to the field lines.

$$Bqv_{y} = \frac{mv_{y}^{2}}{r}$$
$$Bq = \frac{mv_{y}}{r}$$
$$Bqr = mv\sin\theta$$

From <u>Electric Fields</u>, we know that an electron is attracted to positively charged plates and repelled from negatively charged plates.

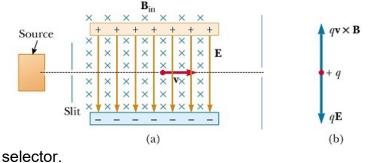
We can make use of this behaviour, along with the behaviour of charged particles in <u>magnetic fields</u>, to produce a <u>velocity selector</u>.

12 Velocity selector

The velocity selector is a device which makes use of <u>crossed</u> magnetic and electric fields to <u>select a beam of charged particles of a fixed velocity</u> originally travelling together with a bunch of other particles with different velocities.

A charged particle moving in a velocity selector (a) will experience both a <u>magnetic</u> and an <u>electric</u> force.

The selector is set up so these forces act in different directions. Only particles where the downward force equals the upwards force will pass through the slit on the other side of the selector.



Other particles are slightly deflected, and as such cannot pass through the slit.

The selected velocity can be determined by:

$$B_{in}qv_0 = qE$$
$$B_{in}v_0 = E$$
$$v_0 = \frac{E}{B_{in}}$$

As the electric force is constant but the magnetic force is proportional to velocity v, particles with velocity $v > v_0$ will deflect upwards, and

particles with velocity $v < v_0$ will deflect downwards.

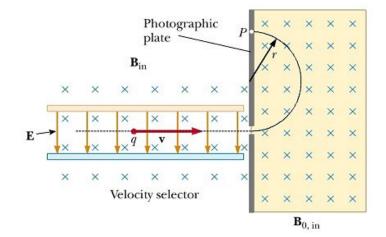
These particles will not pass through the second slit.

With a stream of particles of constant velocity, they can be separated by mass using a <u>mass spectrometer</u>.

13 Mass spectrometer

A mass spectrometer is a device used to measure the mass of charged particles. It consists of a deflection chamber (a region of magnetic field), as well as a photographic plate.

The mass spectrometer is placed after the velocity selector.



For particles in a magnetic field, the magnetic force provides the centripetal force:

$$F_B = F_c$$

$$B_0 q v_0 = \frac{m v_0^2}{r} \Longrightarrow B_0 q = \frac{m v_0}{r}$$

$$r = \frac{m v_0}{B_0 q}$$

$$r = \frac{m \left(\frac{E}{B_{in}}\right)}{B_0 q}$$

$$r = \frac{m e}{B_{in} B_0 q}$$

If $B_{in} = B_0$ (same magnetic field in velocity selector and mass spectrometer),

$$r = \frac{me}{qB^2}$$

If *B*, *q* and *E* are constant, then $r \propto m$.

14 Magnetic field due to current flowing in a long, straight wire

The magnetic field in the vicinity of a long, straight wire takes the form of <u>concentric</u> <u>circles</u> with the wire passing through the common centre.

The direction of the magnetic field can be found through Fleming's Left Hand rule

The strength of the field at a point varies proportionately with the current I and the distance r.

$$B \propto \frac{I}{r}$$

$$B = \frac{\mu_0 I}{2\pi r}$$
*given in formula sheet

Where $\mu_0 = 4\pi \times 10^{-7} Hm^{-1} =$ permeability of free space

15 Magnetic field due to current flowing in a solenoid

The direction of the magnetic field (and hence the polarity) can be determined by the "clock" rule, or, by grasping the solenoid core in the direction of current flow. The thumb will point towards the north pole. This is known as the <u>Right Hand grip rule</u>. The magnitude of the magnetic field is proportional to the current *I* in the solenoid and the number of turns per unit length *n* of the solenoid.

$$\frac{B \propto nI}{^*\text{given in formula sheet}}$$

where $n = \frac{N}{L}$

16 Magnetic field due to current in a circular coil

The magnetic field takes the form of <u>closed loops</u> about any curved portion of the wire in the circular coil.

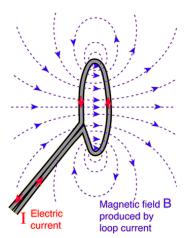
Theoretically, the magnetic field at the centre of the coil is directed through its centre. This field line appears to come from infinity and end at infinity.

One side of the coil corresponds to the <u>North</u> pole and the other side corresponds to the <u>South</u> pole and the polarity can be determined using the <u>right-hand grip rule</u>.

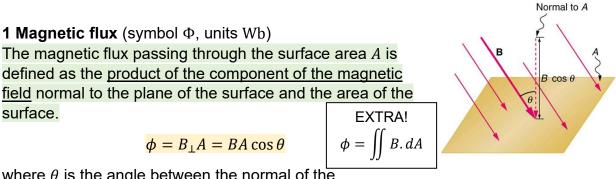
The magnitude of the magnetic field at the centre is proportional to the current I and the number of coils N, and inversely proportional to the radius r.

$$B \propto \frac{NI}{r}$$
$$B = \frac{\mu_0 NI}{2r} \quad * e^{\frac{1}{2}}$$

given in formula sheet



Chapter 18: Electromagnetic Induction



where θ is the angle between the normal of the plane and the magnetic field.

The unit of magnetic flux is the weber (Wb).

The weber is the flux of a uniform magnetic field *B* of flux density 1 *T*, through a plane surface of area *A* of 1 m^2 , placed normally to the *B* field.

The total magnetic flux linkage passing through a <u>coil of N turns</u> can be calculated using:

$$\Phi = N\phi = NBA\cos\theta$$

where *N* is the number of turns in the coil.

2 Faraday's law

Faraday's law states that when the magnetic flux linkage with a circuit is changed, an <u>induced e.m.f.</u> is set up whose magnitude is proportional to the rate of change of flux linkage.

$$\varepsilon = \left| \frac{d\Phi}{dt} \right|$$

3 Lenz's law

Lenz's law states that the <u>direction</u> of the induced e.m.f. (and hence current flow in a closed circuit) is always such as to <u>opposite the change in flux causing it</u>.

$$\varepsilon = -k\frac{d\phi}{dt} = -\frac{d\phi}{dt}$$

where k = 1, and ε , ϕ and t are in their respective S.I. units.

4 Applications of Lenz's law

Pushing a magnet into a solenoid	Pulling a magnet out of a solenoid
An <u>induced north pole</u> is set up at the right of the solenoid to <u>oppose the</u> <u>increasing flux linkage</u> from the approaching north pole. A current is induced and flows as shown. In layman terms, the induced north pole is trying to repel the incoming north pole.	An <u>induced south pole</u> is set up at the right of the solenoid to <u>oppose the</u> <u>decreasing flux linkage</u> from the retreating north pole. A current is induced and flows as shown. In layman terms, the induced south pole is trying to prevent the north pole from leaving.
At the instant switch S is closed	At the instant switch S is opened
Solenoid	
The current in the solenoid increases, increasing the strength of the magnetic field. Hence, <u>the magnetic flux linkage with</u> <u>the coil increases</u> . Induced current in the coil will flow to generate a south pole facing the electromagnet to oppose the flux growth. In layman terms, the coil is trying to prevent the south pole from becoming stronger.	The current in the solenoid decreases, decreasing the strength of the magnetic field. Hence, the <u>magnetic flux linkage with</u> <u>the coil decreases</u> . Induced current in the coil will flow to generate a north pole facing the electromagnet to oppose the flux decay. In layman terms, the coil is trying to prevent the south pole from becoming weaker.
No e.m.f. / current induced when the current is at steady state.	

A rectangular coil entering a magnetic field	A rectangular coil leaving a magnetic field
Induced B field (out of plane) 	Induced B field (into plane) x + x + x + x + x + x + x + x + x + x +
Area of coil linking with the B-field	Area of coil linking with the B-field
increases.	decreases.
Induced current will flow to generate a	Induced current will flow to generate a
B-field that opposes the flux increase.	B-field that opposes the flux decrease.
No e.m.f. / current induced when the entire coil is in the B-field.	

5 Eddy currents

A useful application of Lenz's law is in electromagnetic braking through the formation of eddy currents.

Eddy currents are induced currents flowing in loops.

A simple experimental set-up can be designed to show the effects of eddy currents.

A metal disc is set to swing on a pivot, freely. As the disc passes the magnet, it experiences a <u>changing magnetic flux</u>. Hence, <u>eddy currents</u> are induced.



These eddy currents set up a magnetic field which opposes the field of the magnet, which retards the motion of the disc.

Eddy currents are used in applications such as stopping rollercoasters, galvanometers, voltmeters and ammeters.

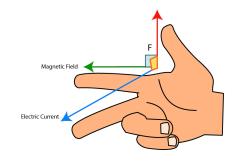
To minimize eddy currents, slots can be cut in the disc. This hinders the formation of large induced currents, which reduces the braking effect.

This is important when eddy currents should be <u>reduced</u> to avoid the loss of energy in applications such as a transformer, as induced currents do work and raise the temperature of the iron core and cause energy loss.

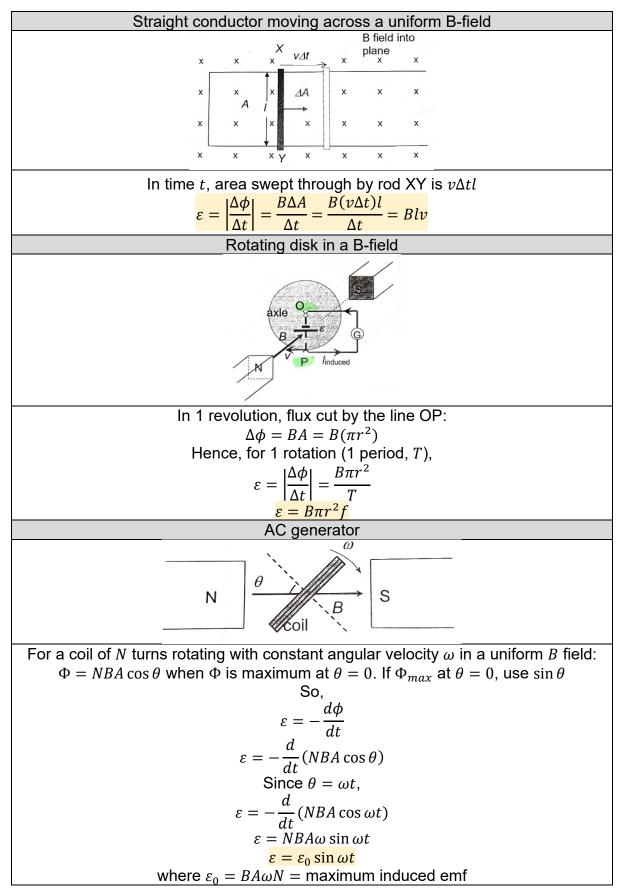
6 Fleming's Right Hand rule

Fleming's right hand rule allows us to identify the direction of the induced e.m.f. and hence current in a circuit.

The se<u>C</u> ond finger represents the <u>C</u>urrent, the <u>F</u> irst finger represents the magnetic <u>F</u> ield and the thumb represents the force (or the direction of motion).



7 Calculations of induced e.m.f.



8 Factors affecting magnitude of induced e.m.f.

From $\varepsilon = BA\omega N$, the factors affecting the magnitude of e.m.f. produced are: magnetic flux density *B*, Plane area of the coil *A*, Angular frequency of rotation ω , Number of turns of the coil *N*

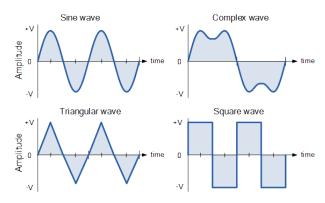
Chapter 19: Alternating Current

1 Introduction

Direct current (D.C.) is one where the current only flows in one direction.

The magnitude of the current may vary, but the direction of the current must never change.

Alternating current (A.C.) is one where the direction of the current flow changes periodically with time (with a constant period).

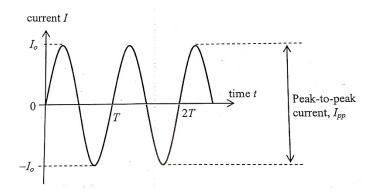


2 Key terms

Period:	The time taken for the current to undergo one complete cycle. Given by T
Frequency:	The number of complete cycles undergone by the current per unit time. Given by f , and $f = \frac{1}{r}$
Angular	I
frequency:	The frequency in terms of radians per unit time rather than cycles per unit time. Given by ω
	$\omega = \frac{2\pi}{T} = 2\pi f$
Peak current:	The maximum magnitude of the current attained in each cycle. Given by I_0
Peak-to-peak	
current:	The difference between the maximum and minimum values of the current within one cycle. Given by I_{pp} For a sinusoidal wave, $I_{pp} = 2I_0$

3 Alternating current (sinusoidal) properties

In the H2 syllabus, the most commonly encountered form of alternating current is that with a sinusoidal waveform.



This sinusoidal alternating current is represented mathematically by $I = I_0 \sin \omega t$

where $\omega = \frac{2\pi}{T} = 2\pi f$

NOTE: if the current starts at the peak value instead, use $I = I_0 \cos \omega t$.

4 Instantaneous Power

To calculate instantaneous power, we simply have to take the value of I or the value of V at the desired time, and apply the formula:

$$P = I^2 R = \frac{V^2}{R}$$

However, to calculate average power, we need to use the <u>mean-square</u> value of the current.

5 Mean-square and root-mean-square

The mean-square value is the <u>mean (average value) of the square of the current</u> over the time interval being considered.

$$\langle I^2 \rangle = \frac{I_1^2 + I_2^2 + \dots + I_n^2}{n}$$

EXTRA!

For currents that vary periodically with time, the mean-square current between time t_1 and time t_2 can be calculated as such:

$$\langle I^2 \rangle = \frac{1}{T} \int_{t_1}^{t_2} I^2 \, dt$$

NOTE: $\langle I^2 \rangle \neq \langle I \rangle^2$

The root-mean-square current of an alternating current (a.c.) is that value of the direct current (d.c.) that would produce thermal energy at the same rate in a resistor.

The root-mean-square value is the <u>square root of the mean of the square of the</u> <u>current</u> over the time interval being considered.

It is simply the square root of the mean-square current. Hence,

 $I_{rms} = \sqrt{\langle I^2 \rangle}$

6 Average power

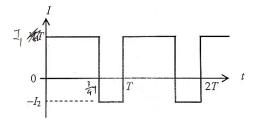
Instead of using the instantaneous current / voltage, we will now replace it with the mean-square / root-mean-square currents and voltages respectively. Hence,

$$\langle P \rangle = \langle I^2 R \rangle = I_{rms}^2 R$$
$$\langle P \rangle = \langle \frac{V^2}{R} \rangle = \frac{V_{rms}^2}{R}$$

$$\langle P \rangle = V_{rms} I_{rms}$$

For square waves, it is necessary to calculate the rootand mean-square values manually.

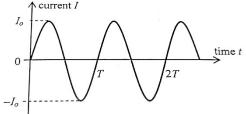
To do this, first <u>square each straight section of the</u> <u>square wave</u> and plot a new graph with the values. Then, calculate the mean of these values, which is the area under the graph divided by the total time.



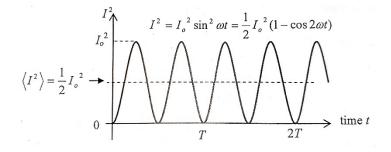
$$\langle I^2 \rangle = \frac{\frac{3}{4}T \times I_1^2 + \frac{1}{4}T \times I_2^2}{T} = \frac{3}{4}I_1^2 + \frac{1}{4}I_2^2$$

For sinusoidal waves, there are some special equations to help us calculate rootmean-square values.

$$I = I_0 \sin \omega t$$
$$I^2 = I_0^2 \sin^2 \omega t$$
$$I^2 = \frac{1}{2} I_0^2 (1 - \cos 2\omega t)$$



Plotting this graph, we obtain the following:



Hence, we can see that the mean-square value for I_0 is $\frac{1}{2}I_0^2$. Hence, the root-mean-square value is

$$I_{rms} = \sqrt{\frac{1}{2}I_0^2} = \frac{I_0}{\sqrt{2}}$$

Similarly,

$$V_{rms} = \frac{V_0}{\sqrt{2}}$$

Then, for average power,

$$\langle P \rangle = \langle I^2 \rangle R = \left(\frac{I_0}{\sqrt{2}}\right)^2 R = \frac{I_0^2 R}{2} = \frac{P_0}{2}$$

Hence, we can see that

$$\langle P \rangle = \frac{P_0}{2}$$

These equations hold true only for a sinusoidal waveform.

7 Transformers

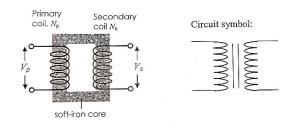
The reason why we use transformers in long-distance power transmission is due to the issue of <u>Joule heating</u>, given by the formula:

 $P_{dissipated} = I^2 R_{transmisson\,wire}$

In order to reduce the current, for the same power, we can increase the voltage (P = VI). This can be done using a <u>transformer</u>. Since D.C. currents are hard to step down or up using a transformer, A.C. current is used for long distance power transmission

for long-distance power transmission.

The high voltage reduces the energy loss as heat generation in the cables.



An alternating voltage V_p applied to the primary coil causes an alternating current I_p to flow inside the coil.

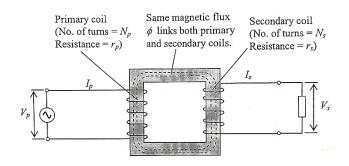
This generates an alternating magnetic flux ϕ (through $\phi = BA = (\mu_0 nI)A$) in the soft iron core.

Assuming no flux leakage (transformer is ideal), the same flux passes through the primary and secondary coils.

This alternating magnetic flux passing through the secondary coil <u>induces an</u> <u>alternating voltage in the secondary coil</u>.

Completing the circuit (i.e. by connecting a load) will allow an induced current to low through the load.

Hence, the transformer transfers electrical power from an input voltage of V_p to an output voltage of V_s .



For an ideal transformer, the relationship between the primary (input) voltage V_p and the secondary (output) voltage V_s is: $\frac{V_s}{V_p} = \frac{N_s}{N_p} = turns \ ratio$

If $N_s > N_p$, then $V_s > V_p$ and the transformer is a <u>step-up</u> transformer. If $N_s < N_p$, then $V_s < V_p$ and the transformer is a <u>step-down</u> transformer. Assuming no power loss (ideal transformer),

$$V_p I_p = V_s I_s$$

Hence,

$$\frac{V_s}{V_p} = \frac{I_p}{I_s}$$

8 Non-ideal transformers

While an ideal transformer transfers 100% of power supplied to the primary coil to the secondary coil, in reality, all transformers lose some amount of energy within the transform.

The <u>efficiency</u> of a transformer is:

efficiency =
$$\frac{\text{power output from secondary coil}}{\text{power input to primary coil}}$$

Below are the reasons for this inefficiency.

Reason	Explanation
Eddy currents	The alternating magnetic flux in the core will also induce circular currents in the soft iron core, called eddy currents.
	To reduce the effect of eddy currents, transformer cores are usually <u>laminated</u> , which means they are made from thin sheets of soft iron separated by layers of insulation. This prevents the formation of large eddy currents and the associated energy loss.
Joule heating	A small amount of energy will be lost as heat in the primary and secondary coils due to Joule heating. $P = I^2 r$
	where r is the resistance of the coil.

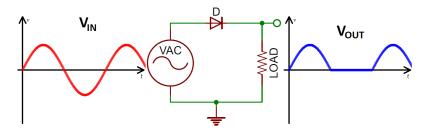
Heating within soft core due to Hysteresis	The soft iron core will <u>retain some of the magnetism</u> caused by the magnetic field. Hence, when the magnetic field is reversed, energy is needed to force the magnetic domains to reverse. This energy is lost as heat generated in the core.
	Hard magnetic materials cannot be used for the core because they strongly retain their magnetisation. Hence, it takes a lot of energy to reverse the magnetisation.
Magnetic flux leakage	In reality, the efficiency of flux coupling between the 2 coils cannot be perfect and there will be some flux leakage.
	This limits the power that can be transferred from the primary to the secondary coil (as the flux linkage is reduced).

9 Half-wave rectification of an alternating current

A diode is a circuit component that only allows current to flow one way through it. If a potential difference applied in the direction that allows it to conduct current, the diode is said to be <u>forward-biased</u>.

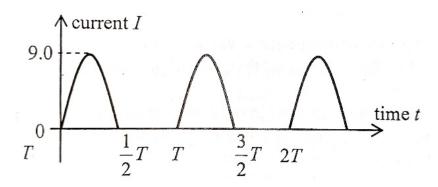
If a potential difference applied in the direction that causes it to block current, the diode is said to be <u>reverse-biased</u>.

Hence, if a diode is connected in <u>series</u> with a <u>sinusoidal alternating current</u>, then the diode will block the current during the half-cycle in which it is reverse-biased.



Hence, the diode is able to convert an A.C. current into a D.C. current. This process is called <u>rectification</u>, and the diode is said to <u>rectify</u> the A.C.

Similar to full sinusoidal waveforms, we can calculate the root-mean-square values for a rectified A.C.



To find the root-mean-square values, we first need to get the graph of I^2 against time.

$$\langle I^2 \rangle = \frac{1}{2}(81) = 40.5$$

 0
 $\frac{1}{2}T$
 T
 $\frac{3}{2}T$
 $2T$
 $\frac{5}{2}T$
 T
 $\frac{1}{2}T$
 $\frac{3}{2}T$
 $\frac{1}{2}T$
 $\frac{1}{2}T$
 $\frac{3}{2}T$
 $\frac{1}{2}T$
 $\frac{1}{2}T$

From t = 0 to $t = \frac{1}{2}T$, the graph obeys the equation $I^2 = 9^2 \sin^2 \omega t$. Hence, the average value $\langle I^2 \rangle$ is $\frac{1}{2}I_0^2 = \frac{1}{2}(81) = 40.5$.

From $t = \frac{1}{2}T$ to t = T, the average current is simply 0. Hence, the mean value of the current $\langle I^2 \rangle$ for <u>one entire cycle</u> is:

$$\langle I^2 \rangle = \frac{\frac{1}{2}I_0^2 \times \frac{1}{2}T + 0 \times \frac{1}{2}T}{T} = \frac{1}{4}I_0^2 = 20.25 \text{ A}^2$$

Hence, the root-mean-square current I_{rms} is:

$$I_{rms} = \sqrt{\frac{1}{4}I_0} = \frac{I_0}{2} = 4.5 \text{ A}$$

Chapter 20: Quantum Physics 1

1 Introduction

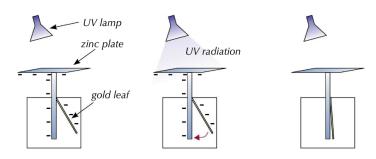
We know that electromagnetic radiation (e.g. light) can behave as waves. However, some phenomena cannot be explained by classical wave theory. For example, the <u>photoelectric effect</u> is a phenomenon that cannot be explained classically.

2 Photoelectric effect

The photoelectric effect is the phenomenon of electrons being ejected from a given meal surface when light about a certain minimum frequency falls upon its surface.

The photoelectric effect can be demonstrated with an <u>electroscope</u>.

The gold leaf is initially repelled due to electron repulsion. When UV light shines on the zinc plate, the gold plate falls. This means that electrons have been emitted.



These electrons are called photoelectrons.

3 Observations that result from the Photoelectric effect

The photoelectric effect results in a disagreement between predictions based on classical theory and the actual observations. The table below lists three wrong predictions and one correct prediction.

Prediction based on classical theory	Observation
Since intensity is a measure of energy, if we use	Wrong prediction.
radiation of enough intensity, electrons would	There <u>is</u> a threshold frequency, below which no
absorb enough energy to escape.	electrons are emitted, no matter the frequency.
Frequency of the light should be irrelevant.	
Electrons can absorb the light energy over time.	Wrong prediction.
Hence, a lower intensity light will take a longer	Photoelectrons are emitted immediately.
time to emit photoelectrons and vice versa.	
For a light of a frequency, a greater intensity	Wrong prediction.
(and hence larger energy) of light should result	Max KE of electrons is independent of intensity
in the emitted photoelectrons having a greater	and dependent on frequency.
kinetic energy.	
For light of any frequency, a greater intensity	Correct prediction.
should result in more emitted photoelectrons per	In fact,
unit time.	rate of emission of photoelectrons \propto intensity

Hence, Einstein came up with his Quantum Theory to explain the discrepancies, which postulated that the energy carried by electromagnetic radiation might exist as discrete packets called quanta.

Hence, incident light was not wave-like, but behaved like projectiles, subsequently called <u>photons</u>, striking the surface of the metal.

4 Photons

A photon is an <u>indivisible</u> quantum (or packet) of electromagnetic energy which is emitted, transmitted and absorbed as a whole.

$$E = hf = \frac{hc}{\lambda}$$

where $h = 6.63 \times 10^{-34}$ Js = planck constant and *E* is the energy of <u>one</u> photon.

The energy of photons is often expressed in <u>electron-volts</u>, as it is easier to read and write.

The <u>electron-volt</u> is the amount of kinetic energy that an electron gains when it is accelerated through a potential difference of one volt.

$$1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$$

With the formula E = hf, we can calculate values such as the rate of emission / arrival of photons in light.

Suppose a lamp has power *P* and emits light of wavelength λ .

The energy of one photon in the light is:

$$E = \frac{hc}{\lambda}$$

If the lamp produces *P* J of energy in 1 second, then the number of photons per second can be deduced:

$$P = \frac{\text{Energy}}{\text{time}}$$
$$P = \frac{Nhf}{t}$$
$$\frac{N}{t} = \frac{P}{hf}$$

Hence,

number per second
$$= \frac{N}{t} = \frac{P}{E}$$

We can also prove the relationship of intensity, frequency and photons. Since intensity is defined as

Intensity =
$$\frac{Power}{Area} = \frac{P}{Area} = \frac{NE}{t} \times Area$$

Substituting E = hf,

Intensity =
$$\frac{\text{Number of photons} \times hf}{t} \times \text{Area}$$

Hence

Intensity =
$$(\frac{n}{t}) \times \frac{hf}{\text{Area}}$$

assuming that the area of incident light is constant.

Assuming that <u>frequency is constant</u>, <u>doubling the intensity</u> will <u>double the rate of</u> <u>photons</u>.

Assuming that <u>intensity is constant</u>, <u>doubling the frequency</u> of light will <u>halve the rate</u> <u>of photons</u>.

5 Work function

When a photon hits an free electron in the metal, the photon must be able to transfer sufficient energy to the electron such that it is able to overcome the electrostatic forces of attraction keeping the electron in the metal.

The minimum amount of energy a photon must have is given by the <u>work function of</u> the metal.

The <u>work function</u> (ϕ) of a metal is the minimum amount of energy required for an electron to escape from the surface.

To find the work function, we can form this equation:

$$E_{photon} = E_{metal} + E_{electron}$$

Since the work function is the minimum energy an electron must possess, from the equation, this is only possible if $E_{electron}$ is zero, i.e. $KE_{electron} = 0$. Hence,

$$E_{photon} = \text{work function}$$

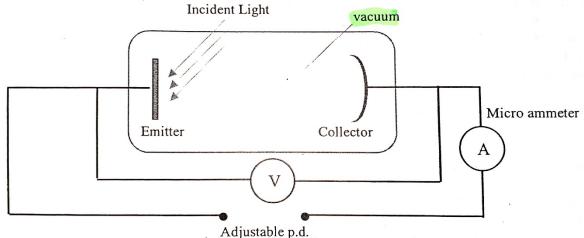
 $\phi = hf_0$

where f_0 is the threshold frequency.

This is the minimum frequency where surface electrons are ejected with <u>zero KE</u>. Below this frequency, no electrons will be emitted.

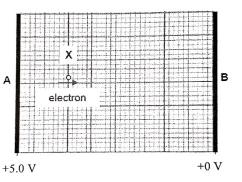
6 Energy of photoelectrons

In order to measure the energy of photoelectrons, we can make an experimental setup.



Imagine that the adjustable p.d. is set such that the collector is of a lower potential than that of the emitter, where A is the emitter and B is the collector.

In order to reach the collector (B), since the plate is of negative potential, the electron must have sufficient kinetic energy to convert into electric potential energy such that it has $KE \ge 0$ at plate B.



From electric fields, we know that $E = q\Delta V$. In the case of an electron,

 $E = e\Delta V$

Since we know that the electron must have enough kinetic energy to convert into electric potential energy,

$$\begin{split} E_{electron} - E_{EPE} &\geq 0 \\ \frac{1}{2} m_e v_{max}^2 - e \Delta V &\geq 0 \end{split}$$

Hence, the potential that stops the most energetic electron is when

$$e\Delta V_s = \frac{1}{2}m_e v_{max}^2$$

Higher potentials cause the value of $E_{electron} - E_{EPE}$ to be less than 0. Hence, the electron will not reach the collector, resulting in 0 current.

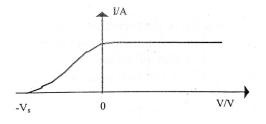
This potential that stops the most energetic photoelectron is called the stopping potential V_s .

The stopping potential is the minimum value of the retarding voltage which will just stop even the most energetic photoelectrons from reaching the collector, indicated by current just dropped to zero as measured in a micro-ammeter.

On the other hand, if <u>the collector is at a higher potential than that of the emitter</u>, this means that electrons experience an attractive force towards the emitter. Hence, all photoelectrons emitted will reach the collector.

This also means that the current will be a constant value regardless of the potentials of the emitter and the collector.

Hence, we can obtain this graph:



where V_s is the stopping potential.

Earlier, we showed how the rate of photons is dependent on intensity on light. Here, we will prove how the rate of <u>photoelectrons</u> is also dependent on the intensity of light.

The <u>photoelectric current</u> I_{PE} is a measure of the number of photoelectrons emitted per second.

$$I_{PE} = \frac{Q}{t} = \frac{N_e \times e}{t} = \left(\frac{N_e}{t}\right)e$$

Since intensity is proportional to the <u>rate of arrival of photons</u>, this should increase the <u>rate of emission of photoelectrons</u> as more photons are striking the metal surface per unit area per unit time and thus more photoelectrons will be emitted. One photoelectron is emitted when approximately 10⁶ photons have struck the surface of the metal.

7 Einstein's photoelectric equation

The equation $E_{photon} = E_{metal} + E_{electron}$ is for a general situation of the photelectric effect. However, Einstein's equation is the study of one particular energy situation, where minimum energy is used to overcome the forces of attraction of the metal and the KE of the photoelectron is the maximum. Hence,

 $E_{photon} = Work function of metal + Maximum KE of electron$

$$hf = \phi + \frac{1}{2}m_e v_{max}^2$$

where $m_e = 9.11 \times 10^{-31} \text{ kg}$

8 Various graphs

Since $eV_s = \frac{1}{2}m_e v_{max}^2$, we can rewrite the equation as

$$hf = eVs + \phi$$

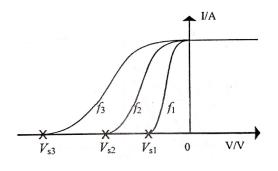
This allows us to see that by <u>varying the</u> <u>frequency of light</u> shone on one particular metal, <u>different stopping potentials V_s is</u> <u>needed</u> to reduce the measured current to zero.

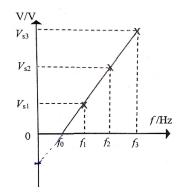
We can also rewrite the equation as

$$V_s = \frac{h}{e}f - \frac{\phi}{e}$$

This allows us to plot a graph of V_s against f.

Connecting the points, we get a straight line where the gradient is $\frac{h}{\rho}$ and the y-intercept is $-\frac{\phi}{\rho}$





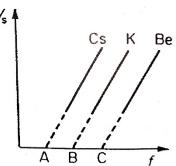
When the light is at the threshold frequency, electrons emitted have 0 KE. Hence, 0 EPE is needed to stop these electrons. By $EPE = e\Delta V$, the stopping potential will be zero.

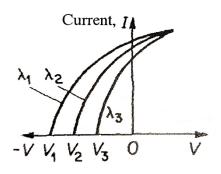
The gradient of every $V_s - f$ graph will be the same, as $\frac{h}{e}$ is a universal constant. However, different metals have different work functions. Since the y-intercept is $-\frac{\phi}{e}$, different metals will have different y-intercepts.

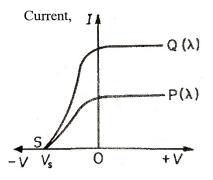
We can also consider the I-V graph for light of various frequencies, while keeping the intensity constant. This means that the <u>rate of arrival of photons</u> will be constant for each frequency, meaning that they will all have the same <u>saturation current</u>. Additionally, by E = hf, different frequencies of light will have a different stopping potential, since $\phi = hf_0$.

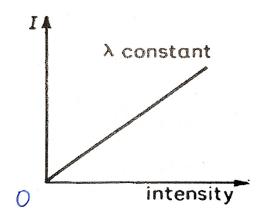
If we consider the I-V graph for light of varied intensities instead of frequency, keeping frequency constant, we can see that since the <u>rate of arrival of photons</u> is different, the <u>rate of emission of photoelectrons</u> will also be different, and hence the saturation current will be different. This is because $Intensity = {N \choose t} \times {hf \over Area}$, and a change in intensity without a change in frequency will change ${N \choose t}$.

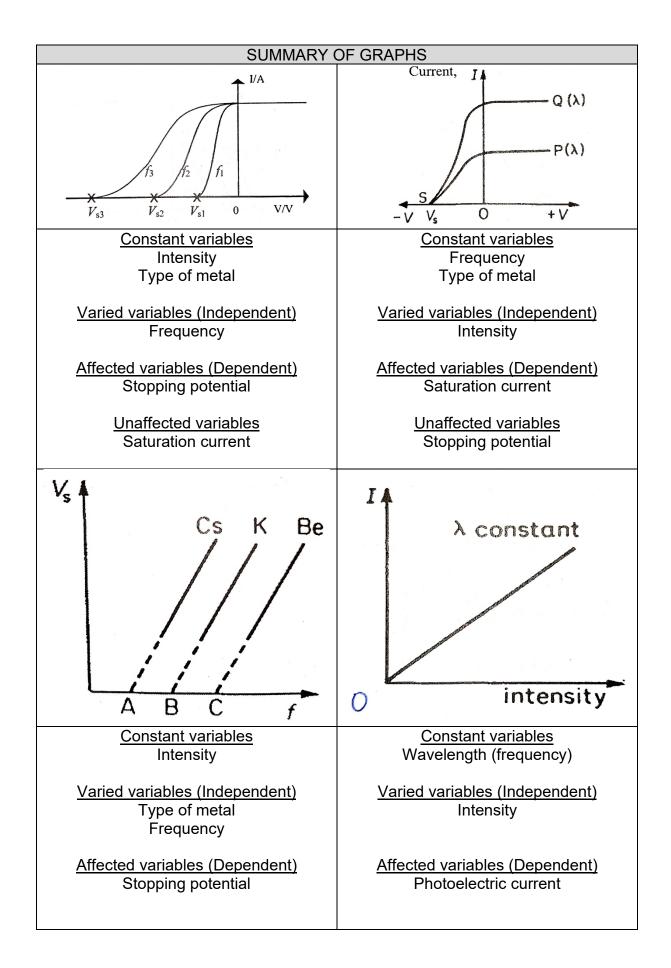
We can also consider the <u>current-intensity</u> graph for light of constant frequency. Earlier, we proved that intensity is related to the rate of emission of photoelectrons. Hence, a larger intensity of light will have a higher saturation current and vice versa.

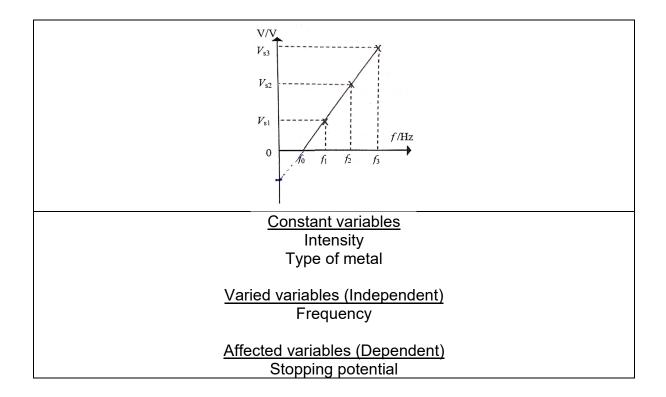












Chapter 21: Quantum Physics 2

1 Wave-particle duality

Light can behave as both a <u>particle</u> (photon) or a <u>wave</u>. However, it will never be both at the same time.

Wave characteristics: Interference, diffraction Particle characteristics: Photoelectric effect

In actual fact, <u>matter</u> can also behave like waves. This is summed up by de Broglie's equation:

$$\lambda = \frac{h}{p} = \frac{h}{mv}$$

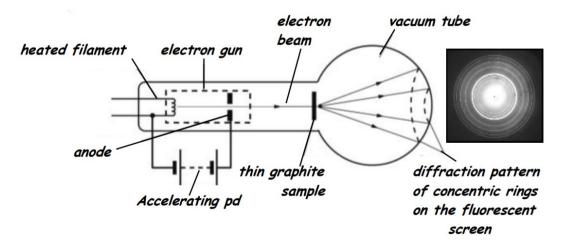
where *h* is the Planck constant $h = 6.63 \times 10^{-34}$ Js.

As a result, even though light is massless, it can exert a force, by Newton's Second Law:

$$F \propto \frac{dp}{dt} \propto \frac{\Delta p}{\Delta t}$$
 where $p = \frac{h}{\lambda}$

2 Electron diffraction

In an experiment, electrons were observed to diffract. This effectively confirmed de Broglie's equation.



The filament (cathode, -) is heated and emits electrons via <u>thermionic emission</u> (not photoelectric).

Electrons are then accelerated to the anode through a potential difference V. This leaves them with an energy $E_k = eV$.

The electrons are then diffracted by the thin graphite sample using $d \sin \theta = n\lambda$. The wavelength of the accelerated electrons can be calculated by using the alternate formula for kinetic energy:

$$E_k = \frac{1}{2}mv^2 = \frac{(mv)^2}{2m} = \frac{p^2}{2m}$$

By de Broglie's equation,

$$E_k = \frac{h^2}{\lambda^2 (2m)} = eV$$
$$V = \frac{h^2}{\lambda^2 (2m)e}$$

3 Energy levels

Each atom possesses discrete energy levels.

Electrons revolve around the nucleus with uniform motion due to the coulomb force ($F = \frac{Qq}{4\pi\epsilon_0 r^2}$).

Electrons can jump from an orbit of lower energy to one of higher energy (E_1 to E_2). This process is called <u>excitation</u> and requires energy.

Electrons can also jump from an orbit of higher energy to one of lower energy (E_2 to E_1) This process is called <u>de-excitation</u> and releases energy in the form of a photon.

Both processes are able to show evidence that discrete energy levels exist in an atom.

The ground state of an atom refers to the most <u>stable</u>, and lowest energy state available.

The energy of a given level is usually expressed in electronvolts. $(1eV = 1.6 \times 10^{-19}J)$

By convention, the energy of an electron at rest outside the atom is taken to be 0. This means that all other energy levels have negative values.

The energy level of a <u>hydrogen atom</u> is given by $E_n = -\frac{13.6}{n^2}$.

 E_{∞} is also known as the ionisation energy, if the ground state is taken to be at 0 energy.

Any state above the ground state are known as excited states.

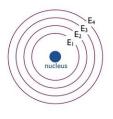
The <u>excitation energy</u> is the energy required to raise the atom from its ground state to an excited state.

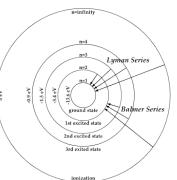
4 De-excitation

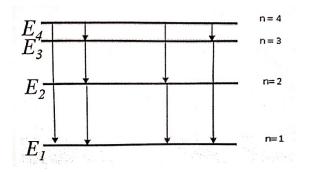
De-excitation occurs when an electron jumps from an orbit of <u>higher energy</u> to one of <u>lower energy</u>. Energy does not need to be supplied to the atom.

Electrons in higher orbits are not stable and the atom will want to de-excite by allowing the electron to jump back to the ground state.

Electrons can jump to any state that is lower energy than the current one. That is, it can go from n=4 to n=3 or n=4 to n=2 or directly from n=4 to n=1.







The number of ways an electron can jump from the n^{th} state to the ground state is given by:

number of ways = ${}^{n}C_{2}$

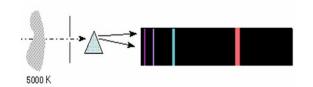
When an electron de-excites, it emits a photon equal to the energy difference between the states:

$$E_2 - E_1 = hf$$

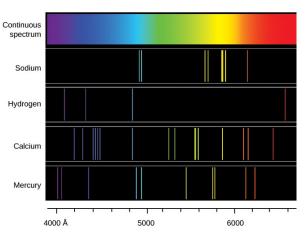
This phenomena is used to produce unique spectral fingerprints, which are evidence for the discrete energy levels.

The spectral fingerprints are produced by passing the photons emitted (light) from de-excitation through a prism, which will produce an <u>emission line spectrum</u>.

The emission line spectrum consists of <u>bright lines</u> of definite wavelengths on a <u>dark</u> <u>background</u>.



The hot gas contains electrons in excited states. These electrons then deexcite, emitting photons of <u>definite</u> <u>wavelengths</u>, which are then diffracted by the prism.



5 Excitation

Excitation occurs when an electron jumps from an orbit of <u>lower energy</u> to one of <u>higher energy</u>. As a result, energy needs to be supplied to the atom. This can take the form of bombarding the atom with 1) electrons or 2) photons.

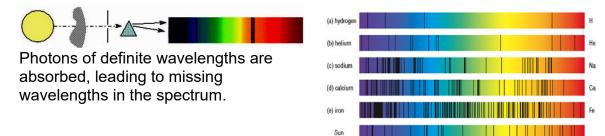
Particles of suitable energy will cause an electron to jump from a lower energy level to one that is higher.

<u>Electrons of any energy that is higher than the minimum excitation energy</u> will cause an excitation. This is because incoming electrons can collide with electrons in the atom and leave with any excess energy.

However, only photons whose <u>energy equals to that of the excitation energy</u> will cause an excitation. This is because photons are <u>indivisible</u> and are <u>absorbed as a</u> <u>whole</u> (Q Phy 1). This means that photons cannot leave with excess energy and all energy must be transferred at once. This phenomenon can demonstrate the existence of discrete energy levels through the production of <u>unique spectral fingerprints</u>.

Instead of passing photons emitted from de-excitation through a prism, white light is shone at a gas. Atoms in these gases then absorb photons whose wavelengths exactly match the energy level difference. The light is then passed through a prism, producing an <u>absorption line spectrum</u>.

The absorption line spectrum consists of a continuous spectrum crossed by dark lines due to some missing wavelengths.



Note: While the electrons excited by the light de-excite immediately, emitting the same wavelength of photons they absorbed, these photons are emitted in <u>all</u> <u>directions</u>. Hence, the intensity of that particular wavelength emitted in one direction is very small relative to the rest of the wavelengths. Hence, dark lines are seen.

6 Examples

An electron de-excites from n=2 (-10 eV) to n=1 (-6 eV). The photon emitted is of wavelength $\lambda = \frac{hc}{\Lambda E}$:

$$\lambda = \frac{6.63 \times 10^{-34} \times 3 \times 10^8}{[-6 - (-10)] \times 1.6 \times 10^{-19}} = 311 \text{ nm}$$

Electrons accelerated by a potential difference bombard an atom which has energy levels n=1 (-12 eV), n=2 (-10 eV) and n=3 (-6 eV).

a) What is the maximum wavelength the electron must have to cause an excitation? Given $m_e = 9.11 \times 10^{-31} \,\mathrm{ms^{-1}}$

Minimum energy to cause excitation $= [-10 - (-12)] \times 1.6 \times 10^{-19}$ $= 3.2 \times 10^{-19} \text{ J}$ $\frac{p^2}{2m} = 3.2 \times 10^{-19}$ $p = \sqrt{2m \times 3.2 \times 10^{-19}}$ $= 7.6357 \times 10^{-25} \text{ kgms}^{-1}$ $\lambda_{max} = \frac{h}{p}$ $= 8.68 \times 10^{-10} \text{ m}$

b) Instead of electrons, photons now bombard the atom. What is the maximum wavelength of the photon to cause excitation?

Minimum energy to cause excitation
$$= 3.2 \times 10^{-19} \text{J}$$

$$E = \frac{hc}{\lambda}$$

$$\lambda_{max} = \frac{hc}{E}$$

$$= \frac{6.63 \times 10^{-34} \times 3 \times 10^{8}}{3.2 \times 10^{-19}}$$

$$= 6.21 \times 10^{-7} \text{ m}$$

c) Why is there a difference in the maximum wavelengths?

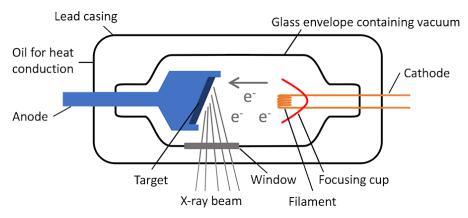
This is because we can assume that photons travel at the speed of light, while electrons cannot be assumed to do so.

7 X-rays

X-rays are <u>ionising radiation</u> with short wavelengths of about 10^{-11} to 10^{-8} m. They can be <u>diffracted</u> by passage through crystals but cannot be focused by lenses.

X-rays can be used for medical therapy, research in crystal structures or used to diagnose flaws in large welds and castings.

X-rays can be produced through a method similar to one demonstrating electron diffraction.



The cathode (-) and anode (+) are connected to a large e.m.f. source.

The cathode is heated by the filament, producing electrons by <u>thermionic emission</u>. These electrons are accelerated at high speed towards the anode through a potential difference of about 10 kV to 1 MV.

The apparatus is <u>evacuated</u> of air so that the path of the electrons is unobstructed.

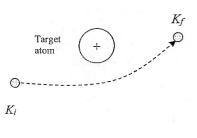
The target contains of atoms with a <u>high atomic number</u> such as tungsten which has a <u>high melting point</u>.

The intensity of X-rays produced increases with the atomic number, as heavier elements have a higher <u>effective nuclear charge</u>, resulting in stronger interactions with electrons, making them lose more energy through braking.

8 Bremsstrahlung radiation (braking)

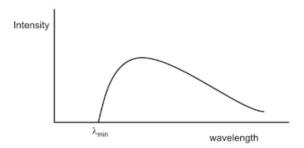
As electrons approach a nucleus, the electron slows down and loses some kinetic energy, which is released as X-ray photon energy.

$$K_i - K_f = \frac{hc}{\lambda}$$
$$\lambda = \frac{hc}{K_i - K_f}$$



As a variable amount of energy can be lost, a continuous range of wavelengths are obtained during braking. This gives rise to the continuous spectrum.

9 Continuous spectrum



The continuous spectrum is formed by a continuous range of wavelengths obtained from billions of electrons losing different amounts of kinetic energy during braking.

When the electrical energy given to an accelerated electron is <u>totally converted into</u> <u>X-ray photon energy</u> as the electron strikes the target in <u>one single interaction</u>, the wavelength of the photon produced is the minimum wavelength as shown in the graph.

$$Ve = \frac{hc}{\lambda_{min}}$$
$$\lambda_{min} = \frac{hc}{Ve}$$

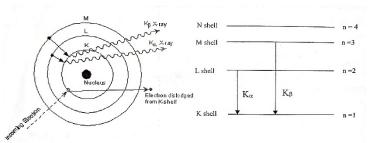
This process is sometimes called the inverse photoelectric effect.

This equation is known as the Duane-Hunt Law. This minimum wavelength depends only on the accelerating p.d. and is the same for all target materials.

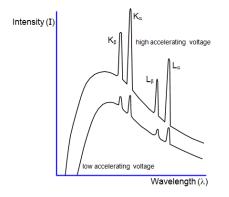
10 Characteristic spectrum

Characteristic X-ray spectra are due to electron transitions deep within atoms of high atomic numbers.

When incoming electrons dislodges electrons in a shell, that shell will have a <u>vacancy</u>. Electrons in higher orbits can then jump to that shell, releasing <u>photons</u> with a definite wavelength. There will then be a spike in intensity for photons of that wavelength.



The characteristic X-ray spectra are unique to each element.



While the spectra is the same for each element, the minimum wavelength depends on the accelerating voltage as seen in the Duane-Hunt law.

10 Heisenberg's Uncertainty Principle

Heisenberg's Uncertainty Principle states that the product of the uncertainty Δx in the position of a body at any instant and the uncertainty Δp in its momentum at the same instant is greater than or approximately equal to the Planck constant.

$\Delta x \Delta p \gtrsim h$

The more precisely one variable is known, the less precisely the other is known. Hence, it is impossible to know the exact speed (mass can be exact) or exact position of an object, as the uncertainty in the other would approach infinity.

One way to think of it as it is impossible to tell the speed of a car from a picture (knowing its position exactly), likewise it is impossible to tell the exact position of a car from a video (knowing its speed exactly).

Chapter 22: Nuclear Physics 1

1 Nucleus

In Nuclear Physics, the atom is characterised as having an <u>extremely small</u> central nucleus in which the positive charge and mass are concentrated.

The nucleus is composed of <u>protons</u> and <u>neutrons</u> (except the hydrogen nucleus). It can be represented by the notation:

 A_{Z}

where *A* is the atomic number (proton + neutron) *Z* is the proton number (proton)

The number of neutrons can be calculated by A - Z.

Atoms of the same element can have a different atomic number. These elements are known as <u>isotopes</u>.

Isotopes are atoms of the same element with the same number of protons but a different number of neutrons.

The mass of a proton is marginally different from that of a neutron.

 $m_{proton} = 1.6726 \times 10^{-27} \,\mathrm{kg}$ $m_{neutron} = 1.6749 \times 10^{-27} \,\mathrm{kg}$

Since the masses are only marginally different, we can define a separate unit which can be thought of as the average weight of a nucleon (particle in the nucleus):

 $1u = 1.66 \times 10^{-27} \text{ kg}$

where *u* is the <u>unified atomic mass</u>.

The unified atomic mass is equivalent to one-twelfth the mass of a carbon-12 atom. The mass of an atom is often expressed in terms of the unified atomic mass. For example, the mass of a carbon-12 atom is 12 u.

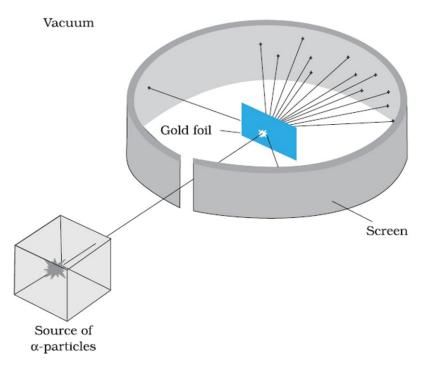
The mass of an atom can also be expressed by the <u>relative atomic mass</u>. The relative atomic mass is defined as the ratio of the mass of the atom to the mass of one-twelfth of the mass of the neutral carbon-12 isotope.

Relative atomic mass, $A_r = \frac{\text{mass of atom}}{\frac{1}{12} \text{ the mass of } \frac{12}{6} \text{C atom}}$

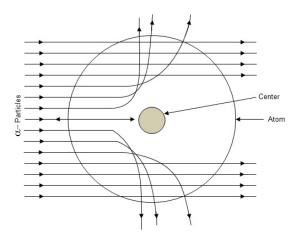
2 Rutherford's scattering experiment

This experiment, conducted in 1909, helped Rutherford to publish his model of the atom In 1911, which is now accepted as the correct model.

Rutherford and his team fired a beam of alpha particles at a gold foil a few atoms (3-4) thick in a special apparatus, shown below.

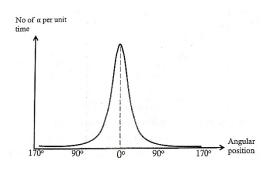


Alpha particles which are emitted with high energy bombard the thin gold foil and are scattered due to Coulomb repulsion (electrostatic repulsion between nuclei).



Most alpha particles pass straight through. Some are deflected forwards (<u>forwards</u> <u>scatter</u>), some are deflected backwards (<u>back scatter</u>) while some are deflected directly backwards (<u>back-tracking</u>)

Scattered alpha particles are then detected using the screen (fluorescent screen). When the particles strike the screen, tiny flashes of light are observed. This gives rise to the following graph:



Rutherford and his team were able to use these observations and make deductions about the model of the atom.

Observation	Deduction
Most alpha particles passed straight	The atom consists of mostly empty
through the atom.	space.
Some alpha particles were deflected	The nucleus is positively charged.
forwards and backwards.	
Very few alpha particles were	The nucleus is very small and has a
backtracked.	very large mass relative to the atom.

3 Nuclear reactions

Nuclear reactions can be represented in the following ways:

An equation where the total nucleon number and proton number are the same.	${}^{9}_{4}\text{Be} + {}^{4}_{2}\text{He} \rightarrow {}^{12}_{6}\text{C} + {}^{1}_{0}\text{n}$
A symbolic expression of the form: initial nuclide (incoming, outgoing) final nuclide	9_{A} Be $(\alpha, n) \frac{12}{6}$ C
There are different types of particles involved in pyels or reactions	

There are different types of particles involved in nuclear reactions.

Particle	Representation
α -particle	⁴ ₂ He
β -particle	$0 \\ -1^{e}$
Neutron	$\frac{1}{0}$ n

4 Conservation

Just like in Mechanics, nuclear reactions obey certain conservation laws.

Conserved quantity	
Nucleon number	The sum of the mass numbers on both sides of the equation are the same.
Proton number	The sum of the atomic numbers on both sides of the equation is the same.
Momentum	The linear momentum of all the particles in the system is conserved.
Mass-energy	Since mass and energy are interchangeable, the total mass-energy of the system remains constant.
Ferentially, in helencing nuclear reactions, the total number of nucleans and total	

Essentially, in balancing nuclear reactions, the <u>total number of nucleons</u> and <u>total</u> <u>proton number</u> must remain constant.

5 Einstein's equation

Einstein's famous equation was a result of special relativity. It stated that <u>mass</u> and <u>energy</u> are equivalent:

 $\Delta E = \Delta m c^2$

where E is the energy of the object m is the mass of the object c is the speed of light.

This equation thus relates mass and energy. As a result, any mass has an equivalent amount of energy.

This equation means that a loss of energy (e.g. in chemical reactions) should equate to a loss in mass of an object.

In <u>nuclear reactions</u>, any object that has energy (e.g. KE) will have a larger mass than at rest and any object that loses energy (e.g. nuclear fission) will have a smaller product mass.

The mass of an object that is at rest is called its rest mass.

6 Consequences of $E = mc^2$

In order to break up the nucleus by separating it into its individual components (neutrons and protons), energy needs to be <u>supplied</u> to the nucleus.

EXTRA! Energy needs to be supplied to overcome the strong nuclear force holding the nucleons together. This force is only significant at very small distances ($\approx 1 \text{ fm}$).

Since $\Delta E = \Delta mc^2$, the additional energy supplied will cause an increase in mass of the neutrons and protons. Hence, the total mass of the <u>nucleons</u> will always be greater than that of the <u>nucleus</u>.

1.5 Range (fm)

$\sum m_{neutrons} + \sum m_{protons} > m_{nucleus}$

The difference (Δm) between the mass of the <u>nucleons</u> and that of the <u>nucleus</u> is known as the <u>mass defect</u>.

The amount of energy needed to break up the nucleus into its constituent protons and neutrons is known as the <u>binding energy</u>, and can be calculated by:

binding energy = (mass defect) $c^2 = \Delta m c^2$

The <u>binding energy</u> of the nucleus is defined as the energy <u>released</u> when the nucleus is first formed from its separate protons and neutrons.

The <u>binding energy</u> of the nucleus is also defined as the <u>work done</u> on the nucleus to separate the nucleus into individual protons and neutrons.

This energy appears in the form of a γ photon or kinetic energy.

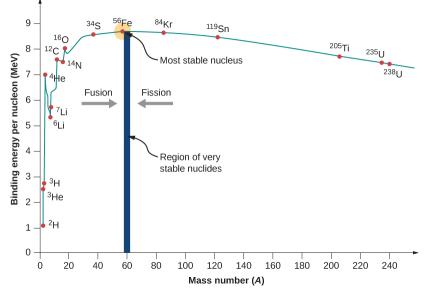
The binding energy of a proton or neutron is zero.

The <u>binding energy per nucleon</u> is simply the binding energy of the nucleus divided by the nucleon number. It is a measure of the <u>stability</u> of the nucleus.

7 Stability of nucleus

As mentioned earlier, the binding energy is a measure of the <u>stability</u> of the nucleus. The higher the binding energy per nucleon, the more stable the nucleus and vice versa.

The graph of binding energy per nucleon against the nucleon number can be plotted.



Note: it is not needed to memorise the peaks in the graph. A smooth curve can be plotted instead.

From the graph, several observations can be made.

The average binding energy per nucleon is about <u>8 MeV</u>, except for lighter nuclei.

Iron-56 is the most stable nuclide as it has the maximum binding energy per nucleon.

Nuclei with low mass numbers (< 56) may undergo fusion such that the final product will have a higher binding energy per nucleon.

Nuclei with high mass numbers (> 56) may undergo fission such that the daughter nuclei will have a higher binding energy per nucleon.

EXTRA!

Why do we compare binding energy per nucleon and not total binding energy? Stability is relative. If a daughter particle exists that has lesser total binding energy (e.g. $\frac{2}{3}BE_{parent}$) and half the number of nucleons ($\frac{A_{parent}}{2}$), even though the binding energy is less, the daughter particle is actually more stable as the binding energy per nucleon is higher.

The peak of the graph at iron means that reactions that involve 1) fusion of low mass number particles and 2) fission of high mass number particles will release energy.

In general,

Energy released = $BE_{final} - BE_{initial}$

Energy is taken in to break up the reactants $(-BE_{initial})$. Energy is then released when the final product is formed (BE_{final}) .

Einstein's equation can also be used to calculate the energy released.

8 Nuclear fission

Nuclear fission is the disintegration of a <u>heavy nucleus into 2 or more lighter nuclei</u> when bombarded by a particle, usually a neutron, and in the process releasing more neutrons.

Energy is released in the process as the total binding energy of the products is higher than that of the reactants.

Using $\Delta E = \Delta m c^2$, the mass defect Δm is positive. Hence, energy is released.

This energy is released in the form of kinetic energy of the particles or γ -radiation.

Example of a nuclear fission reaction:

$${}^{235}_{92}\text{U} + {}^{1}_{0}\text{n} \rightarrow {}^{144}_{56}\text{Ba} + {}^{90}_{36}\text{Kr} + {}^{3}_{0}\text{n} + \text{energy}$$

9 Nuclear fusion

Nuclear fusion occurs when <u>2 light nuclei</u> combine to form a <u>single more massive</u> <u>nucleus</u> leading to a release of an enormous amount of energy.

This is because the increase in binding energy per nucleon for fusion particles is greater (left side of graph is steeper), a greater amount of energy per nucleon is released.

Fusion is <u>difficult</u> to achieve because of the <u>strong electrical repulsion</u> (proton and proton) between the nuclei. Hence, the reaction needs to occur at extremely high temperatures (about 100 million K). Even then, the probability of the reaction occurring is small.

Example of a nuclear fusion reaction:

$${}^2_1\mathrm{H} + {}^3_1\mathrm{H} \rightarrow {}^4_2\mathrm{He} + {}^1_0\mathrm{n}$$

Chapter 23: Nuclear Physics 2

1 Radioactivity

Radioactivity or radioactive decay is the <u>spontaneous</u> and <u>random</u> disintegration of heavy unstable nuclei into more stable products through the emission of radiation such as <u>alpha particles</u>, <u>beta particles</u> and <u>gamma rays</u>.

Radioactivity is spontaneous because it occurs on its own accord and is not affected by external factors.

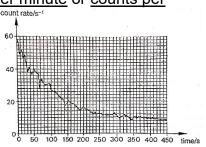
This means that <u>physical factors</u> (e.g. temperature, pressure) and <u>chemical factors</u> (e.g. state; aqueous, liquid) do not influence the process of decaying.

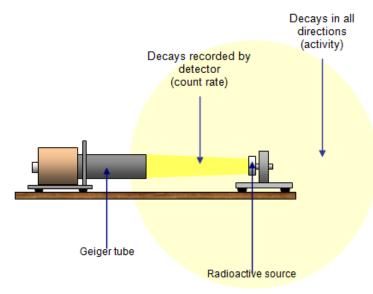
Radioactivity is random because it is not possible to predict which nucleus will decay or when it will decay.

Radioactivity can be measured using a <u>Geiger-Muller tube connected to a ratemeter</u>. The Geiger-Muller tube displays the radioactivity in <u>counts per minute</u> or <u>counts per</u> <u>second</u>.

Using the Geiger-Muller tube, the random nature of radioactivity can be shown from the uneven and random peaks and troughs in the count rate.

The Geiger-Muller tube does not detect the true amount of radiation that the source gives off. Instead, it only detects what enters the tube.





The count rate follows an inverse square law, that is:

count rate = $\frac{\text{Area}}{4\pi r^2} \times \text{Activity}$ count rate $\propto \frac{1}{r^2}$

2 Background radiation

Even when there are no explicit radioactive sources present, there will still be background radiation, for example from cosmic rays or the decay of radon.

The background radiation ranges from <u>20 to 50 counts per minute</u>, or <u>0.333 to 0.833</u> <u>counts per second</u>.

In experiments where the emission rate from radioactive sources is <u>large</u>, the background radiation is <u>negligible</u> and hence can be ignored.

In experiments where the emission rate from radioactive sources is <u>small</u>, the background radiation is <u>not negligible</u>. Hence, the actual emission rate can be calculated by subtracting the background radiation.

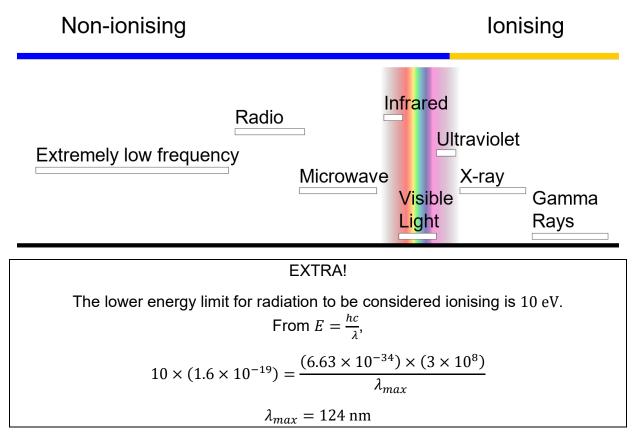
3 lonising radiation

lonising radiation refers to radiation that has enough energy to strip electrons from an atom, thus producing free electrons and ions.

Ionising radiation can be electromagnetic or nuclear (alpha, beta, gamma) in nature.

When an atom is ionised, energy is absorbed. Hence, the radiation loses energy. This is why alpha particles are stopped by a few cm of air (see below).

Gamma rays, X-rays, and radiation in the higher part of the ultraviolet spectrum are considered ionising.



When exposed to ionising radiation, there will be short-term and long-term effects.

Short-term effects include blistering of the skin, nausea, vomiting, radiation sickness, loss of hair, emaciation, and even death.

Long-term effects include anaemia, leukaemia, cancer, sterility and genetic mutation.

Safety precautions can be taken, both in the lab and in the real world to prevent this.

In the real world, workers should take precautions such as:

Storing radioactive materials in lead containers to stop radiation Handling samples by remote control or at a safe distance Using shielding from equipment producing or scattering radiation Wearing radiation badges which can inform the user of radiation levels

In the lab, students should observe the following procedures:

Handle radioactive substances with tongs rather than fingers to prevent accidental ingestion

Hold radioactive substances at a distance, pointing away from the body and from other people

Wearing rubber gloves

4 Alpha particles

Alpha particles are identical to the helium nuclei. That is,

alpha particle
$$=$$
 $\frac{4}{2}\alpha = \frac{4}{2}He$

Alpha particles have a <u>+2 positive charge</u>.

Alpha particles have a <u>rest mass of 4 u</u>.

Alpha particles have a speed of about 0.1 c

Alpha particles have <u>high ionising power</u> because it moves relatively slowly and easily ionises atoms and molecules in their way.

Alpha particles have <u>low penetrating power</u> because it loses its energy quickly due to ionisation of other particles.

They can be stopped by 0.1 mm of paper or 0.01 mm of aluminium foil and travel only about 2 cm in air.

Alpha particles are <u>deflected slightly</u> by electric and magnetic fields because of their high mass and charged state.

Alpha particles are always emitted with a <u>definite kinetic energy</u> as governed by the Principle of Conservation of Momentum (and energy).

5 Beta particles

Beta particles are electrons. That is,

beta particle =
$${0 \atop -1} \beta = {0 \atop -1} e$$

When a nucleus has too many neutrons, it disintegrates by converting a <u>neutron</u> into a <u>proton</u> and <u>electron</u>.

Beta particles have a <u>-1 negative charge</u>.

Beta particles have a speed of about 0.9 c.

Beta particles have <u>lower ionising power</u> compared to alpha particles as they are small and fast moving.

Beta particles have <u>high penetrating power</u>. They can be stopped by 5 mm of aluminium foil and travel about 5 m in air.

Beta particles are <u>easily deflected</u> by electric and magnetic fields as they are 7300 times lighter than alpha particles.

Beta particles are emitted with a <u>range of kinetic energies</u> due to the production of neutrinos. Neutrinos are needed to fulfil the conservation laws.

6 Gamma rays

When a nucleus decays, it is often left in an excited energy state. It can then decay to its ground state by emitting a photon, also called a gamma (γ) ray.

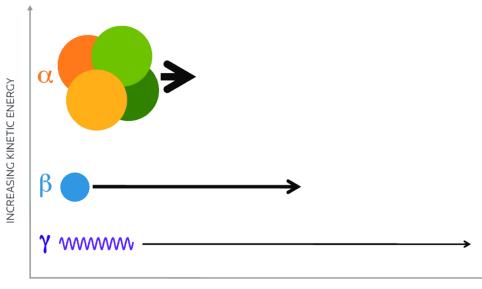
Gamma rays are <u>electromagnetic radiation</u>.

Gamma rays have no charge and no mass.

Gamma rays have a <u>speed of 1 *c*</u>.

Gamma rays cause <u>relatively little ionisation</u> as they have no charge.

Gamma rays have the <u>greatest penetrating power</u> as they travel very far. They can only be stopped with about 5 cm of thick lead or about 10 m of thick concrete and travel about 500 m in air.



INCREASING PENETRATION

7 Half-life (symbol $t_{1/2}$)

Half-life is defined as the time taken for half the number of nuclei of a radioactive element to decay.

8 Activity (symbol *A*, units becquerel = Bq)

The activity of a radioactive material is the number of disintegrations of its atoms per unit time.

9 Decay Constant (symbol λ , units s⁻¹)

The decay constant is the probability that a radioactive nuclide of an isotope in a sample would decay in one second.

10 Decay Law

The rate of radioactive decay $A = -\frac{dN}{dt}$ is directly proportional to the number N of the radioactive nuclei present.

$$-\frac{dN}{dt} \propto N$$
$$-\frac{dN}{dt} = \lambda N$$

10 Equations

$$x = \left(\frac{1}{2}\right)^n x_0$$

where x can be activity Anumber of undecayed particles Ncount rate Cmass of undecayed material m

From
$$-\frac{dN}{dt} = \lambda N$$
,
 $\frac{1}{N} dN = -\lambda dt$
 $\int_{N_0}^N \frac{1}{N} dN = \int_0^t -\lambda dt$
 $[\ln N]_{N_0}^N = [-\lambda t]_0^t$
 $\ln \frac{N}{N_0} = -\lambda t$
 $N = N_0 e^{-\lambda t}$

Generally,

 $x = x_0 e^{-\lambda t}$

where x can be activity Anumber of undecayed particles Ncount rate Cmass of undecayed material m

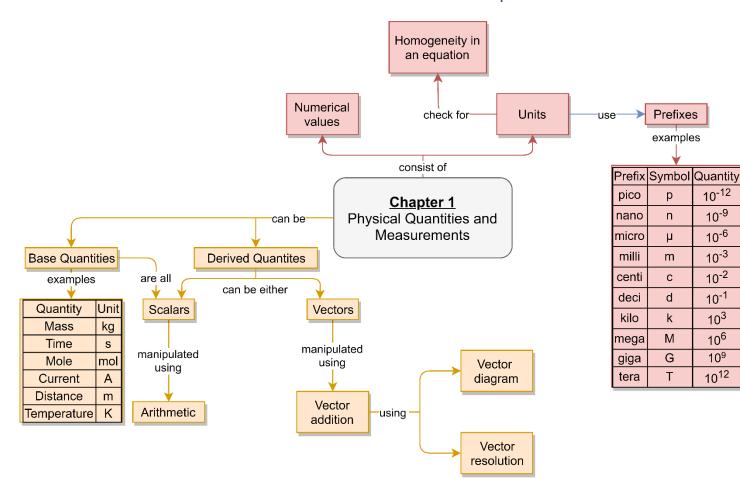
Relating half-life and the decay constant:

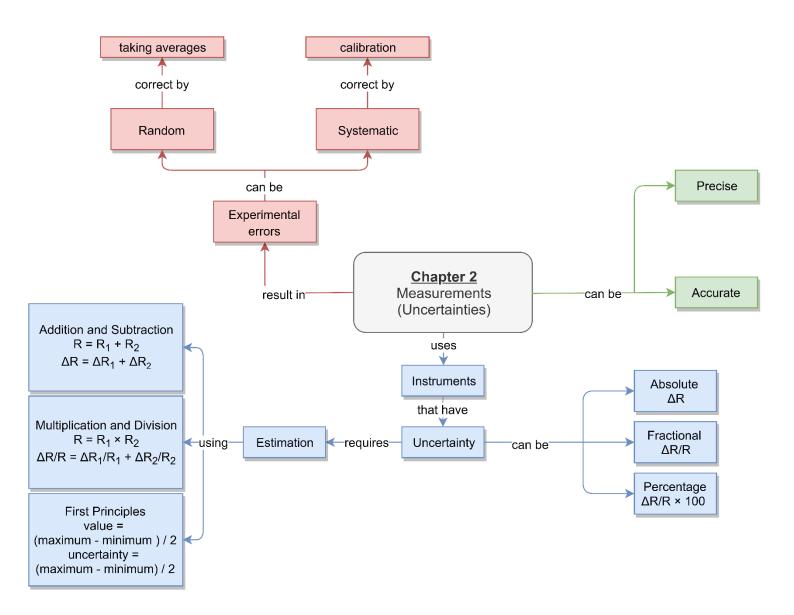
$$t_{1/2} = \frac{\ln 2}{\lambda}$$

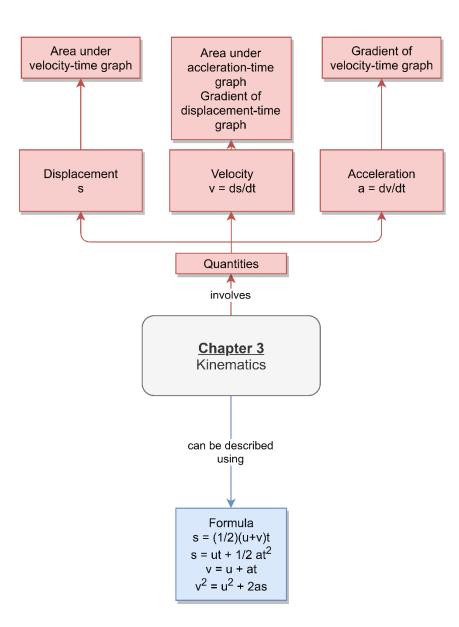
EXTRA!	
Deriving $t_{1/2} = \frac{\ln 2}{\lambda}$:	
When $t = t_{1/2}$	
$x = \frac{1}{2}x_0$	
$ \begin{array}{rcl} x &= \frac{1}{2}x_{0} \\ \frac{x_{0}}{2} &= x_{0}e^{-\lambda t_{1/2}} \end{array} $	
$\lambda t_{1/2} = -\ln\frac{1}{2}$	
$t_{1/2} = \frac{\ln 2}{\lambda}$	

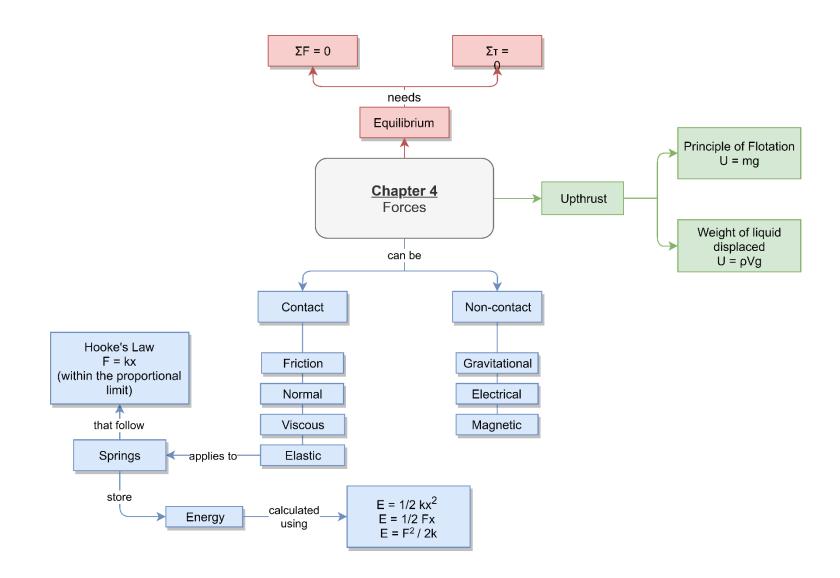
NOTES

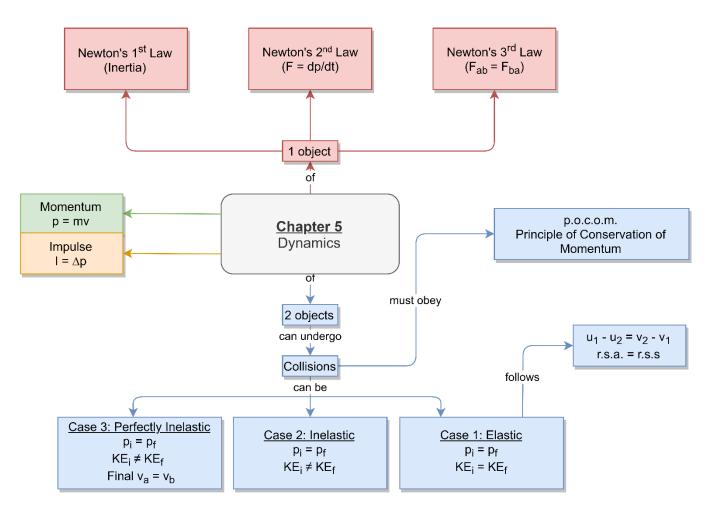
Annex 1: Mind maps

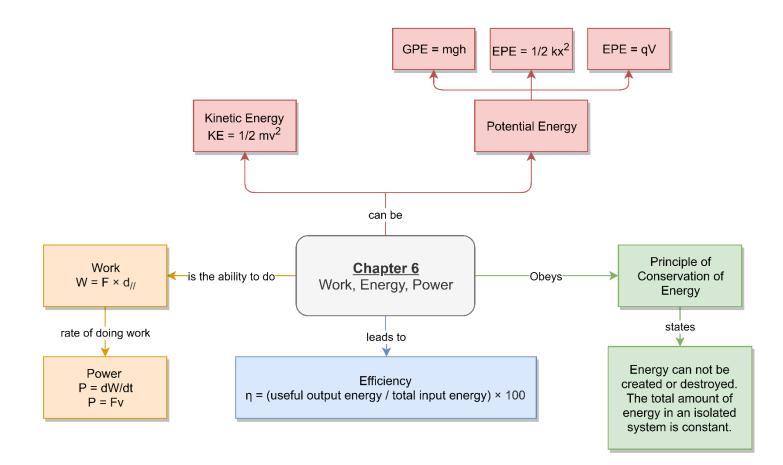


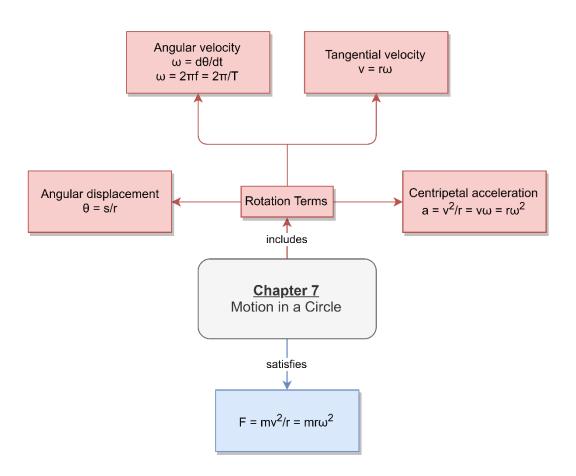


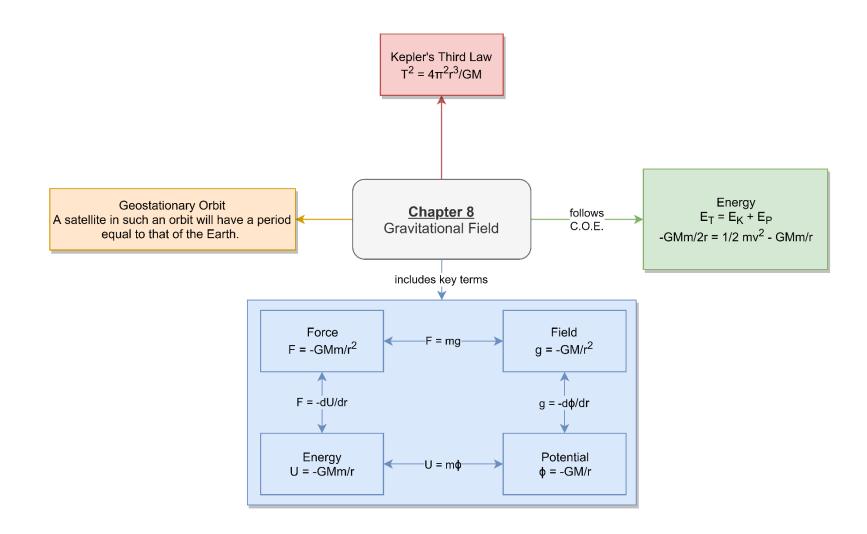


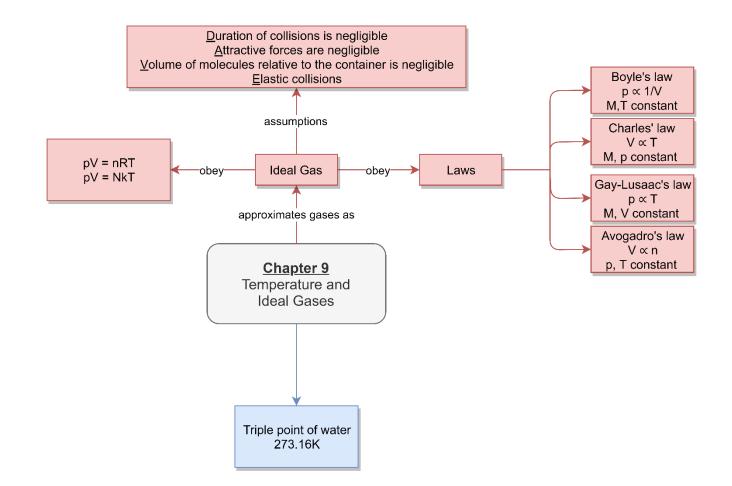


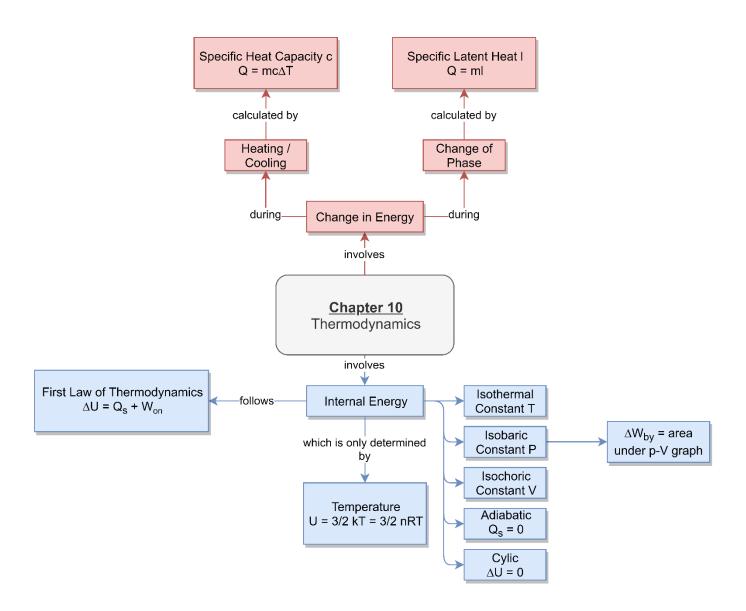


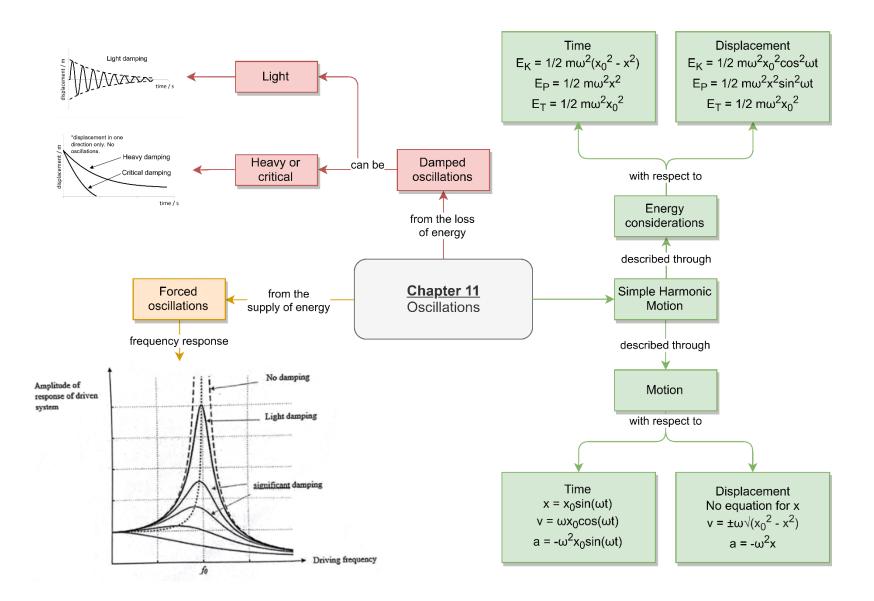


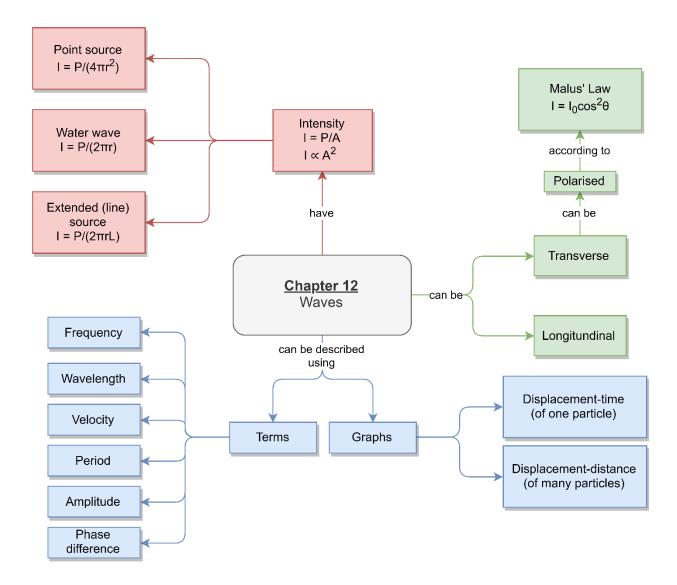


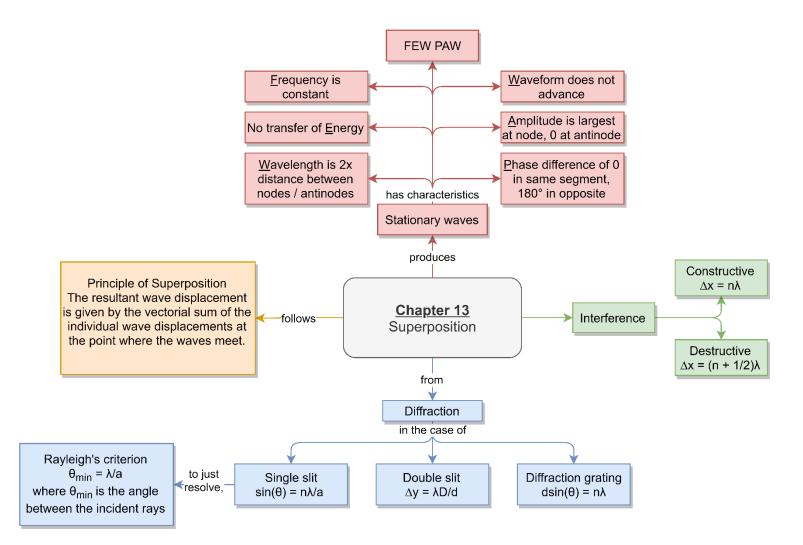


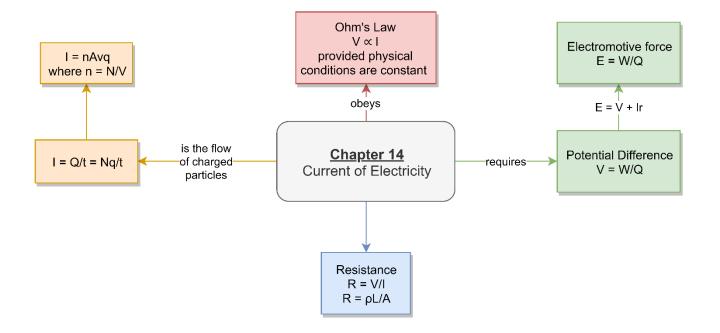


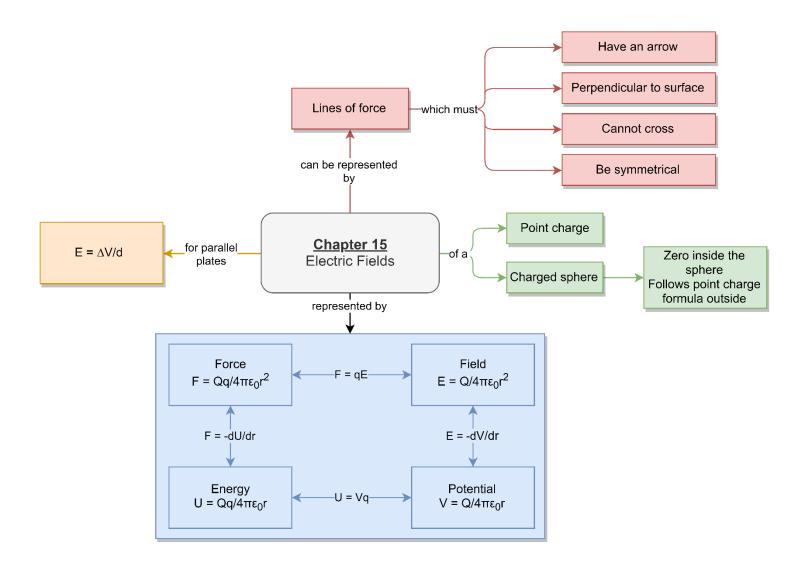


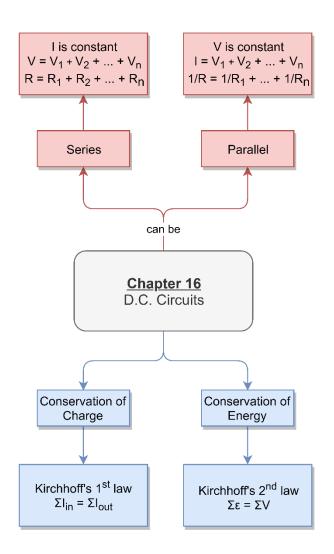


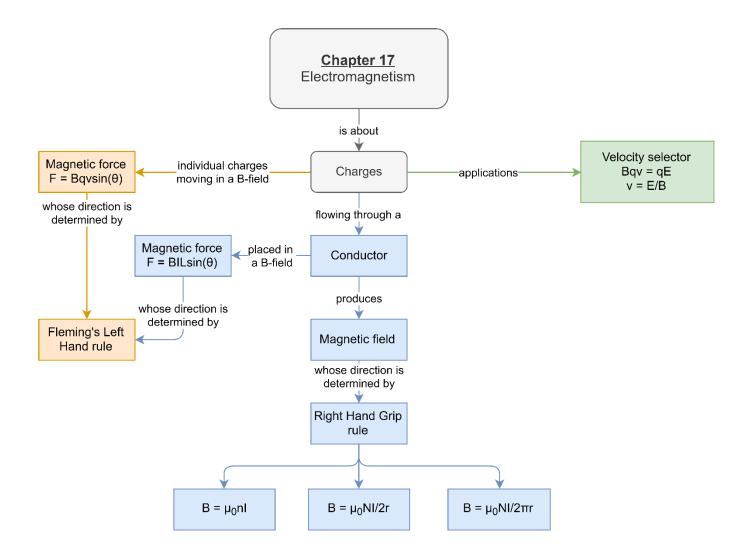


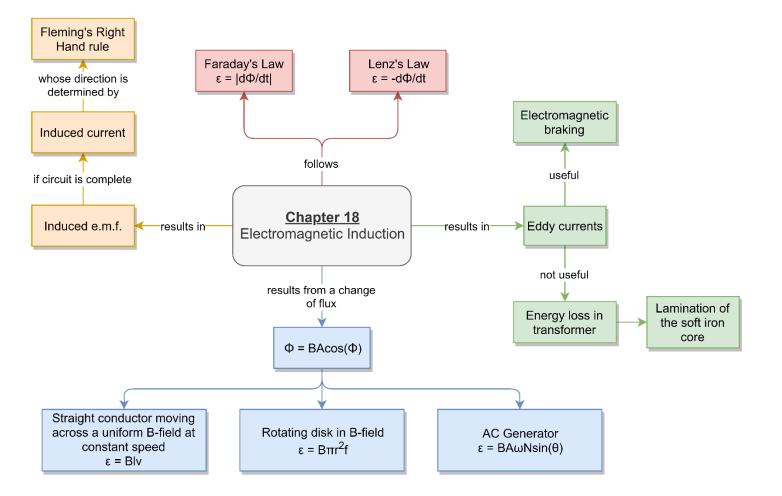


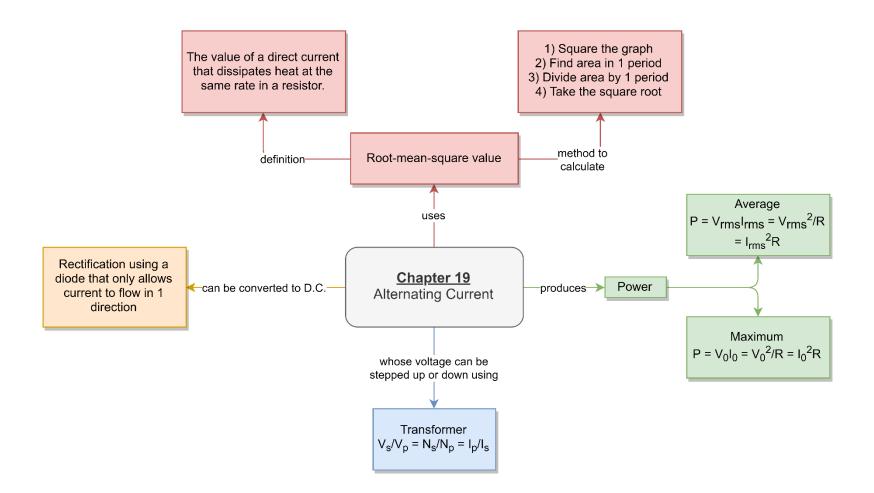


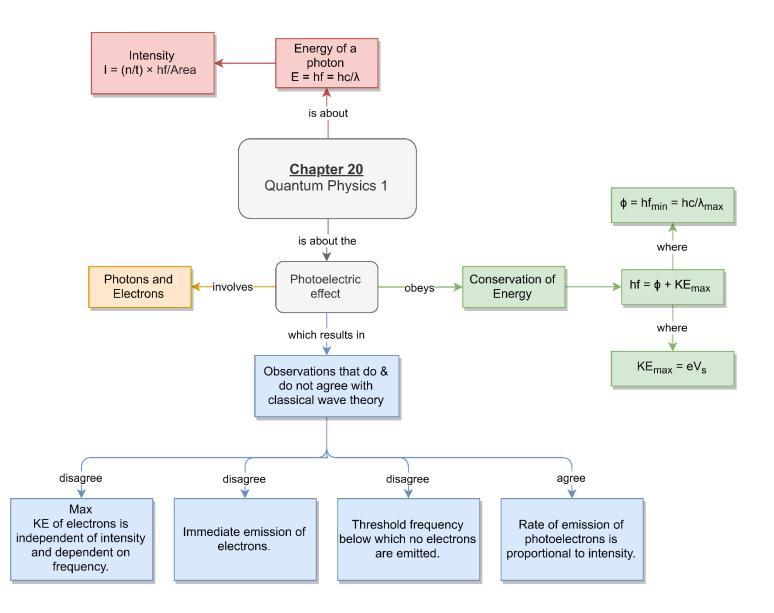


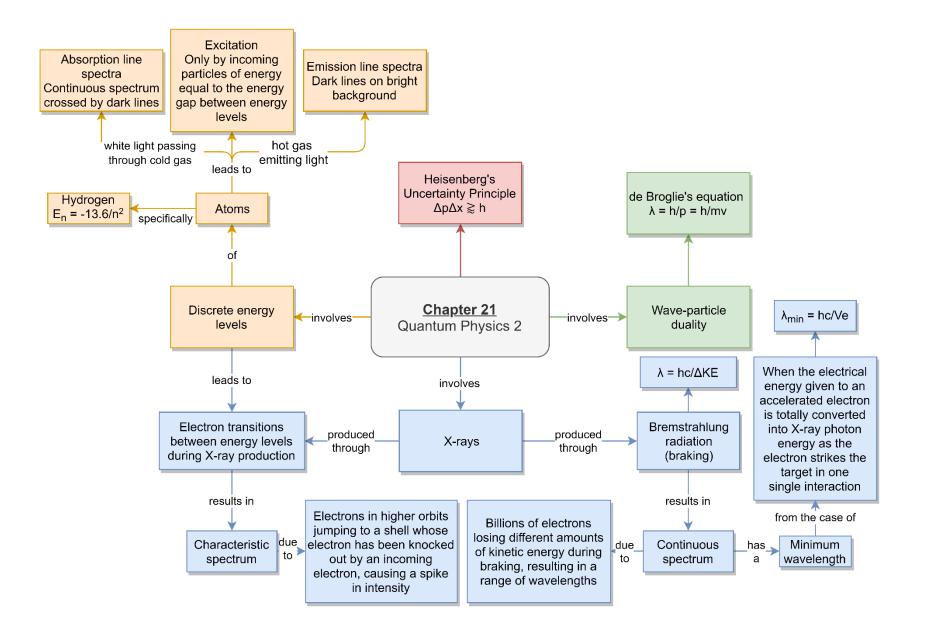


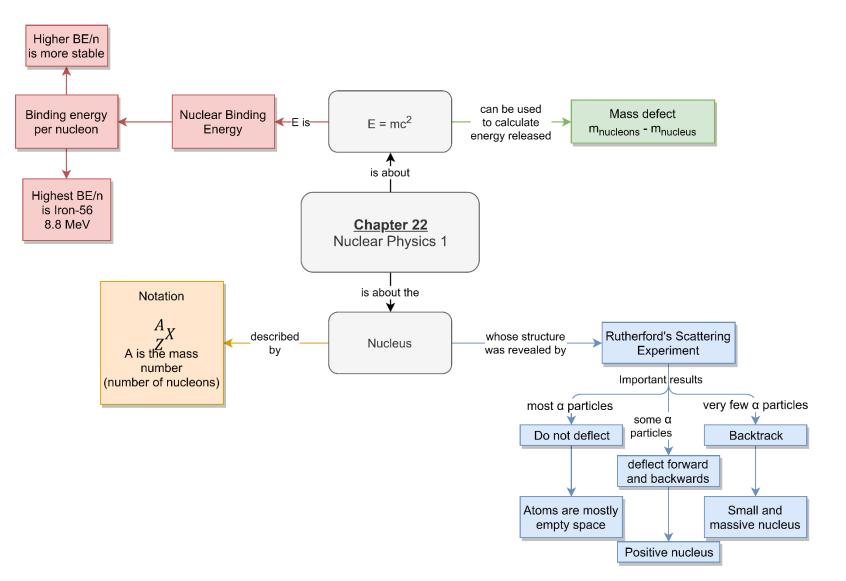


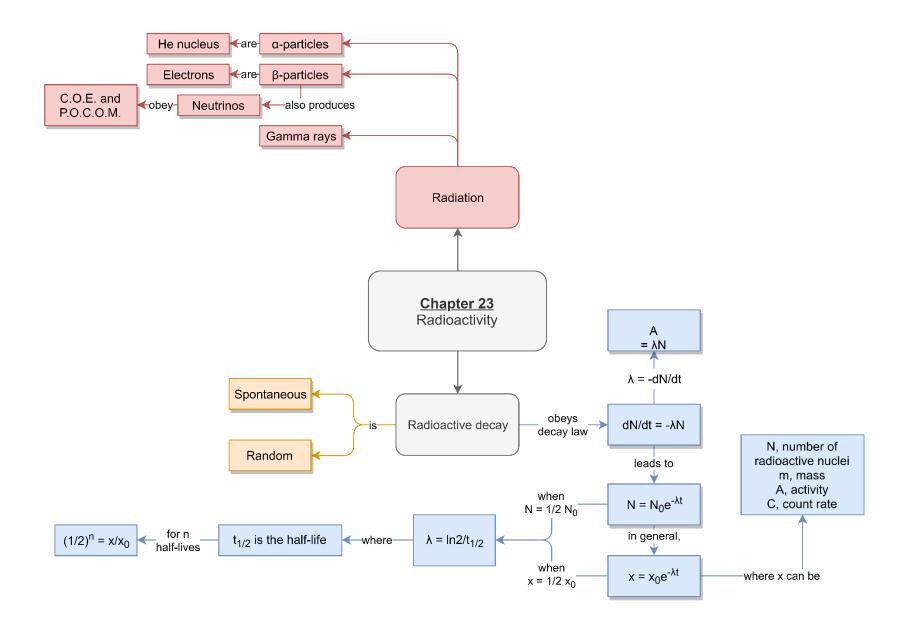












Annex 2: Formula list

Physical Quantities and Units (1)	Kinematics (3)	Dynamics (4)	Forces (5)
$R = \sqrt{A^2 + B^2 + 2AB\cos\theta}$	v = u + at	p = mv	F = kx
	$s = \frac{1}{2} (u+v)t$	$F \propto \frac{dp}{dt}$	$\frac{1}{k_{eff}} = \frac{1}{k_1} + \frac{1}{k_2} + \dots + \frac{1}{k_N}$
	$s = ut + \frac{1}{2}a t^2$	$Impulse = \int_{t_1}^{t_2} F dt$	$k_{eff} = k_1 + k_2 + \dots + k_N$
	$v^2 = u^2 + 2as$	$F_{avg} = \frac{\Delta p}{\Delta t}$	$EPE = \frac{1}{2}Fx = \frac{F^2}{2k} = \frac{1}{2}kx^2$
		$m_1 u_1 + m_2 u_2 = m_1 v_1 + m_2 v_2$	$W = \int_{x_1}^{x_2} F dx$
		$u_1 - u_2 = v_2 - v_1$	$\tau = F \times d_{//}$
		$\frac{1}{2}m_1u_1^2 + \frac{1}{2}m_2u_2^2 = \frac{1}{2}m_1v_1^2 + \frac{1}{2}m_2v_2^2$	p = ho gh
			$Upthrust = \rho LgA$ $(\rho h_1 g)A - (\rho h_2 g)A$

Work, energy, power (6)	Motion in a circle (7)	Gravitational field (8)	
$W = F\cos\theta \left(\Delta x\right)$	$\theta = \frac{s}{r}$	$F = -\frac{GMm}{r^2}$	$\phi = -\frac{GM}{r}$
$W = p \Delta V$	$\omega = \frac{d\theta}{dt}$	$g = \frac{F}{m} = -\frac{GM}{r^2}$	$E_T = -\frac{GMm}{2r}$
$E_k = \frac{1}{2}mv^2$	$\omega = \frac{2\pi}{T} = 2\pi f$	$v_{orb} = \sqrt{\frac{GM}{r}}$	$F = -\frac{dU}{dr}$
GPE = mgh	$v = r\omega$	$r_{geosat} = \sqrt[3]{\frac{GMT^2}{4\pi^2}}$	$g = -rac{d\phi}{dr}$
$P = \frac{\Delta W}{\Delta t}$	$a = v\omega = r\omega^2 = \frac{v^2}{r}$	$T^2 = \frac{4\pi^2}{GM} r^3 \text{ or } T^2 \propto r^3$	
P = Fv	$F_C = \frac{mv^2}{r} = mr\omega^2$	$T_{binary} = \sqrt{\frac{4\pi^2 (R_1 + R_2)^3}{G(M_1 + M_2)}}$	
$\eta = \frac{\text{useful power output}}{\text{power input}}$		$U = -\frac{GMm}{r}$	
		$v_{escape} = \sqrt{\frac{2GM}{r}}$	

Temperature and ideal gases (9)	Thermodynamics (10)	Oscillations (11)	
U = KE + PE		$a = -\frac{k}{m}x$	$T = 2\pi \sqrt{\frac{m}{k}}$
$\frac{\theta}{100} = \frac{X_{\theta} - X_i}{X_s - X_i}$	$Q = C\Delta T = mc\Delta T$	$T = 2\pi \sqrt{\frac{m}{\rho Ag}}$	
$X^{\circ}C = K + 273.15$	C = mc	$x = x_0 \sin(\omega t)$	
pV = nRT = NKT	$Q = m\ell$	$v = \frac{dx}{dt} = \omega x_0 \cos(\omega t)$	$v = \pm \omega \sqrt{x_0^2 - x^2}$
$k = \frac{R}{N_a}$	$\Delta U = \Delta Q_s + \Delta W_{on}$	$a = \frac{dv}{dt} = -\omega^2 x_0 \sin(\omega t)$	$a = -\omega^2 x$
$C_{rms} = \sqrt{\langle c^2 \rangle} = \sqrt{\frac{c_1^2 + c_2^2 + \dots + c_N^2}{N}}$	$W_{by} = \int_{V_i}^{V_f} p dV = p \Delta V$	$E_{K} = \frac{1}{2}mv^{2} = \frac{1}{2}m\omega^{2}(x_{0}^{2} - x^{2})$	$E_K = \frac{1}{2}m\omega^2(x_0^2 - x^2)$ $= \frac{1}{2}m\omega^2 x_0^2 \sin^2(\omega t)$
$p = \frac{1}{3}\rho < c^2 >$		$E_P = \frac{1}{2}kx^2 = \frac{1}{2}m\omega^2 x^2$	$E_P = \frac{1}{2}m\omega^2 x^2$ $= \frac{1}{2}m\omega^2 x_0^2 \cos^2(\omega t)$
$\langle KE \rangle = \frac{3}{2}kT$		$E_T = E_K + E_P = \frac{1}{2}m\omega^2 x_0^2$	

Wave Motion (12)	Superposition (13)	Current of Electricity (14)	Electric Fields (15)
v =	fλ	$I = \frac{Q}{t}$	$E = \frac{F}{+q}$
$y = A\sin\left(\frac{x}{\lambda} \times 2\pi\right)$	constructive interference: $\Delta x = n\lambda$ where $n \subseteq \mathbb{Z}^+$	$I = nAv_d q$	$E = \frac{Q}{4\pi\epsilon_0 r^2}$
$y = -A\sin(\omega t)$	$\Delta \emptyset = 2n\pi$ where $n \subseteq \mathbb{Z}^+$	$V = \frac{W}{Q}$	$F = \frac{Q_1 Q_2}{4\pi\varepsilon_0 r^2}$
$\Delta \phi = \frac{\Delta x}{\lambda} \times 2\pi$	destructive interference: $\Delta x = \left(n + \frac{1}{2}\right)\lambda$ where $n \subseteq \mathbb{Z}^+$	$E = \frac{W}{Q}$	$U = \frac{Q_1 Q_2}{4\pi\varepsilon_0 r}$
$\Delta \phi = \frac{\Delta t}{T} \times 2\pi$	$\Delta \emptyset = (2n + 1)\pi$ where $n \subseteq \mathbb{Z}^+$	$R = \frac{V}{I}$	$V_c = \frac{W}{q}$
Intensity \propto (Amplitude) ²	$\sin\theta = \frac{n\lambda}{a}$	$R = \frac{\rho L}{A}$	$V_c = \frac{Q}{4\pi\epsilon_0 r}$
$I = \frac{1}{2}I_0$	$\tan\theta = \frac{w}{2D}$	$V_{terminal} = \varepsilon - Ir$	U = Vq
$I = I_0 \cos^2 \theta$	$ heta_{min} = rac{\lambda}{a}$	$IR_{ext} = \varepsilon - Ir$	$\Delta U = q \Delta V$
	$\Delta y = \frac{\lambda D}{d}$	$P = VI = I^2 R = \frac{V^2}{R}$	$E = -\frac{dV}{dr}$
	$d\sin\theta = n\lambda$		$F = -\frac{dU}{dr}$

D.C. Circuits (16)	Electromagnetism (17)	Electromagnetic Induction (18)	Alternating Current (19)
$\sum_{junction} I_i = 0$	$F = BIL \sin \theta$ $B = \frac{F}{IL \sin \theta}$	$\phi = B_{\perp}A = BA\cos\theta$	$I = I_0 \sin \omega t$
$\varepsilon = \frac{\Delta W}{q}$	<u>A pair of current-carrying wires</u> $F \propto \frac{I_1 I_2}{d}$	$\Phi = N\phi = NBA\cos\theta$	$\langle I^2 \rangle = \frac{I_1^2 + I_2^2 + \dots + I_n^2}{n}$
$\Delta V = \frac{\Delta W}{q}$	<u>A current-carrying rectangular</u> loop $\tau = BIAN$	$\varepsilon = \left \frac{d\Phi}{dt}\right $	$I_{rms} = \sqrt{\langle I^2 \rangle}$
$\sum_{junction} \Delta V_i = 0$	$F = B_{\perp}qv = B\sin\theta qv$	$arepsilon = -krac{d\phi}{dt} = rac{d\phi}{dt}$	$\langle P \rangle = \frac{V_{rms}^2}{R} = I_{rms}^2 R = V_{rms} I_{rms}$
$\frac{\text{Series}}{R_{eff} = R_1 + R_2 + \dots + R_n}$	$f = \frac{B_{in}q}{2\pi m}$	$\frac{\text{Straight conductor moving}}{\text{across uniform B-field}}$ $\varepsilon = Blv$	$I_{rms} = \frac{I_0}{\sqrt{2}}$
$\frac{1}{R_{eff}} = \frac{\frac{\text{Parallel}}{1}}{R_1} + \frac{1}{R_2} + \dots + \frac{1}{R_n}$	Current in a long, straight wire		$V_{rms} = \frac{V_0}{\sqrt{2}}$
$V = \varepsilon - Ir$	$B \propto \frac{l}{r}$ $B = \frac{\mu_0 l}{2\pi r}$	Rotating disk in B-field	$\langle P \rangle = \frac{P_0}{2}$
$V_{1} = \frac{R_{1}}{R_{1} + R_{2}} * V = \frac{R_{1}}{R_{T}} * V$ $V_{AB} = \frac{l_{AB}}{l_{T}} * V_{0}$	Current in a solenoid	$\varepsilon = B\pi r^2 f$	$\frac{V_s}{V_p} = \frac{N_s}{N_p} = turns \ ratio$ $\frac{V_s}{V_p} = \frac{I_p}{I_s}$
$V_{AB}=rac{l_{AB}}{l_T}*V_0$	$B \propto nI \qquad \qquad B = \mu_0 nI$	$\frac{\text{AC generator}}{\varepsilon = NBA\omega\sin\omega t}$	$\frac{V_s}{V_p} = \frac{I_p}{I_s}$
	Current in a circular coil		Half-wave rectification:
	$B \propto \frac{NI}{r} \qquad \qquad B = \frac{\mu_0 NI}{2r}$		$I_{rms} = \frac{I_0}{2}$

Quantum Physics 1 (20)	Quantum Physics 2 (21)	Nuclear Physics 1 (22)	Nuclear Physics 2 (23)
$E = hf = \frac{hc}{\lambda}$	$\lambda = \frac{h}{p} = \frac{h}{mv}$	$A_r = \frac{mass \ of \ atom}{\frac{1}{12} \ the \ mass \ of \ \frac{12}{6}C \ atom}$	count rate $\propto \frac{1}{r^2}$
$1 \text{ eV} = 1.6 \times 10^{-19} J$	$E_k = \frac{1}{2}mv^2 = \frac{(mv)^2}{2m} = \frac{p^2}{2m}$	$\sum m_n + \sum m_p > m_{nucleus}$	$-\frac{dN}{dt} = \lambda N$
$Intensity = (\frac{n}{t}) \times \frac{hf}{Area}$	$\Delta x \Delta p \gtrsim h$	$BE = \Delta mc^2$	$x = \left(\frac{1}{2}\right)^n x_0$
$E_{photon} = E_{metal} + E_{electron}$		$\Delta E = BE_{final} - BE_{initial}$	$x = x_0 e^{-\lambda t}$
$\phi = h f_0$			$t_{1/2} = \frac{\ln 2}{\lambda}$
$hf = \phi + \frac{1}{2}m_e v_{max}^2$ $hf = eVs + \phi$ $V_s = \frac{h}{e}f - \frac{\phi}{e}$			

Annex 3: Definitions

	Word	Definition
CH1	Homogeneity	A physical equation is homogeneous if every term in the equation has the same units.
Errors	Systematic error	Systematic errors are errors of measurements that occur according to some 'fixed rule or pattern' such that they yield a consistent overestimation or underestimation of the true value.
CH2:	Random error	Random errors are errors with different magnitudes and signs in repeated measurements.
atics	Displacement	the shortest distance from the initial to final position of a body along a specific direction.
CH3: Kinematics	Velocity	the rate of change of displacement with respect to time
	Acceleration	the rate of change of velocity with respect to time
	Inertia	the reluctance of a body to change its state of motion or rest.
mics	Mass	Mass is the amount of matter in a body.
Chapter 4: Dynamics	Newton's First law	(resultant) force acts on it to change its state.
	Newton's Second law	
	Newton's Third law	If body A exerts a force (action) on body B, then body B will exert an equal and opposite force (reaction) on body A.
	Impulse	the integral of the force over the time interval in which the force acts.

	Impulso	impulse is equal to the change in momentum
	Impulse-	impulse is equal to the change in momentum.
	momentum	
	theorem	
	Principle of	the total momentum of a system is conserved provided no external unbalanced force acts on the
	Conservation of	system.
	Momentum	
	(p.o.c.o.m)	
es	Hooke's law	Within the proportional limit, the extension e, produced in a material is directly proportional to the force F applied.
5: Forces	Moment	the product of the force and the perpendicular distance from the line of action of the force to the pivot.
Chapter	Couple	2 equal and opposite forces whose lines of action do not coincide.
0	Principle of	For an object in rotational equilibrium, the sum of the anticlockwise moments about the pivot is equal to
	moments	the sum of clockwise moments about the same pivot.
Work	Work done	The work done by a constant force on an object is defined as the product of the magnitude of the displacement and the component of the force parallel to the displacement.
<pre></pre>	Principle of	energy can neither be created nor destroyed, but can be converted from one form to another. The total
r 6:	Conservation of	energy in a closed system is always constant.
Chapter	Energy (c.o.e.)	
Jap	Efficiency	the percentage ratio of useful power output to power input.
Ö	Enciency	

Circular motion	Radian	the angle subtended at the centre of a circle by an arc of equal length to the radius.
CH7: Circ	Angular velocity	the rate of change of angular displacement.
nal field	Gravitational field strength	the force per unit mass acting on any mass placed at that point.
Gravitational field	Gravitational potential energy	With infinity being taken as the zero reference, the GPE at a point is the work done by an external agent in bringing a test mass from infinity to the point without a change in kinetic energy.
CH8: G	Gravitational potential	the work done per unit mass by an external agent in bringing an object from infinity to the point.
	Temperature	A measure of the degree of hotness or coldness of an object.
eal gas	Boyle's law	Pressure of a fixed mass of gas is inversely proportional to its volume if the temperature is constant.
CH9: Ideal	Charles' law	Volume of a fixed mass of gas is proportional to its thermodynamic temperature if the pressure of the gas is constant.
	Gay-Lusaac's law	Pressure of a fixed mass of gas is proportional to the thermodynamic temperature if its volume is constant.

	Heat	the energy that is transferred from one body at a higher temperature to one at lower temperature.
Thermodynamics	Heat capacity	The heat capacity of a body is defined as the quantity of heat energy required to raise its temperature by 1 K.
nodyn	Specific heat capacity	The specific heat capacity of a body is defined as the heat required to raise the temperature of 1 kg of the substance by 1 K.
	Specific latent heat of fusion	
Chapter 10:	Specific latent heat of vaporisation	The specific latent heat of vaporisation is of a substance is defined as the amount of heat energy needed to change a unit mass of the substance from liquid to gaseous state without a change in temperature.
Che	First law of thermodynamics	The first law of thermodynamics states that the internal energy of a system depends only on its state, and that the increase in internal energy of the system is the sum of the heat supplied to the system and the external work done on the system.
Chapter 11: Oscillations	Amplitude	maximum displacement of the oscillating body
	Period	the time taken for one complete oscillation
	Frequency	the total number of oscillations per second
	Angular frequency	the product of 2π and the frequency
	Wavelength	the distance between 2 consecutive particles in the wave who are in phase with one another
	Simple harmonic motion	an oscillating system where (1) there must be a restoring force proportional to the displacement but directed opposite to it and (2) the acceleration of the system must be proportional to the displacement but directed opposite to it
	Light damping	oscillations still occur, but the amplitude of the oscillation decreases with time

	Heavy damping	there are no oscillations. When the system is displaced from the equilibrium position, it takes a long
	neavy damping	time to return to the equilibrium position
	Critical	there are no oscillations. When the system is displaced from the equilibrium position, it returns to the
		equilibrium position in the shortest possible time
	Resonance	when the driving frequency is very close to the natural frequency of a system, the system will oscillate
		at maximum amplitude
lotion	Transverse waves	the direction of vibration of the wave particles is <u>perpendicular</u> to the direction of propagation of the wave
′ave ∿	Longitudinal waves	the direction of vibration of the wave particles is <u>parallel</u> to the direction of propagation of the wave
Chapter 12: Wave Motion	Progressive waves	the wave profile moves in the direction of propagation of the wave. Progressive waves transfer energy in the direction of the wave velocity
Chapte	Intensity	intensity is the energy transmitted per unit time across a unit area of a surface perpendicular to the direction of the energy flow
Chapter 13: Superposition	Principle of Superposition	the principle of superposition states that the resultant wave displacement (and hence the amplitude) is given by the vectorial sum of the individual wave displacements (amplitude) at the point where the waves meet
	Standing wave	a standing or stationary wave is one which the wave profile does not travel in the direction of the wave velocity, though the wave particles still execute oscillatory motion (SHM) about their rest positions
	Diffraction	diffraction refers to the spreading of waves when they encounter an aperture or obstacle whose linear dimension (e.g. width of aperture) is comparable to the wavelength of the waves
	Interference	interference refers to the phenomenon of two or more waves of the same type meeting at a point in space to produce a resultant wave disturbance given by the superposition of individual waves at that point.

	1	
	Constructive	the amplitude of the resultant wave at a given time or position is greater than that of either individual
	interference	wave
	Destructive	the amplitude of the resultant wave at a given time or position is smaller than that of either individual
	interference	wave
	Electric current	the current <i>I</i> is the rate of flow of electric charge Q through a given cross-section of a conductor
		the burrent his the rate of new of electric charge & through a given cross section of a conductor
	Electric charge	the charge Q which flows past a given cross section is the product of the steady current / that flows
		past the section and the time during which the current flows.
	Coulomb	the coulomb C is the amount of electrical charge that peaces through a given cross section of a sireuit
	Coulonib	5 1 5 5
≳		when a steady current of one ampere flows in one second
Electricity		
sct	Charge density	
Ш		conductor
of		
Current o	Potential	
line	difference	other forms of energy per unit charge passing from one point to another
4	Volt	the volt is the potential difference between two points in a circuit if one joule of electrical energy is
, U		converted to other forms of energy per coulomb of charge passing from one point to the other
apte		
Chapter 14:	Electromotive	the electromotive force (e.m.f.) of any source of electrical energy is the total energy converted into
	force	
		571 5 11
	Resistance	Resistance of a conductor is the ratio of potential difference across the conductor to the current flowing
	rtoolotarioo	through it.
	Ohm's law	the current flowing through a conductor is directly proportional to the p.d. across it, providing
	Uniti 3 law	temperature and other physical conditions are constant

	Resistivity	Resistivity is a property of the material, independent of the shape and size of the conductor.
	Electric field	an electric field is a region of space where a charged particle will experience an electric force
Electric fields	Electric field strength	the electric field strength E at a point is the electric force acting per unit positive charge placed at that point
15:	Coulomb's law	Coulomb's Law states that the electric force between two point particles carrying charges Q_1 and Q_2 separated by a distance r apart is proportional to the product of the two charges and inversely proportional to the square of their distance apart
Chapter	Electric potential	the electric potential at a point in an electric field is the work done per unit positive charge by an external agent in bringing a small test charge from infinity to that point without a change in kinetic energy
Chapter 16: D.C. Circuits	Kirchhoff's Current law	Kirchhoff's Current Law states that the algebraic sum of the currents at a junction of a circuit is zero.
	Electromotive force	Electromotive force (symbol ε) of a cell is the amount of energy converted from non-electrical to electrical form per unit charge that moves from one terminal of the cell to the other outside of the cell.
	Potential difference	Potential difference (symbol <i>V</i>) across a circuit component is the amount of energy converted from electrical to non-electrical forms per unit charge that flows through the component.
	Kirchhoff's Voltage law	Kirchhoff's Voltage Law states that the algebraic sum of all electrical potential changes around any closed loop is zero.
	Potential-divider rule	The Potential-Divider rule states that if a voltage exists across several resistors connected in series, then the voltage across each resistor is proportional to the total resistance.

er 17: EM	Magnetic flux density / magnetic field strength Tesla	Magnetic flux density is defined as the force acting per unit length of a conductor which carries unit current and is of right angles to the magnetic field. The Tesla is the unit of magnetic flux density equivalent to a force of 1 N experienced by a straight
Chapter		conductor of length 1 m and carrying a current of 1 A when it is placed perpendicular to the magnetic field.
	Magnetic flux	The magnetic flux passing through the surface area A is defined as the product of the component of the magnetic field normal to the plane of the surface and the area of the surface.
18: EMI	Weber	The weber is the flux of a uniform magnetic field B of flux density 1 T, through a plane surface of area A of $1m^2$, placed normally to the B field.
Chapter	Faraday's law	Faraday's law states that when the magnetic flux linkage with a circuit is changed, an induced e.m.f. is set up whose magnitude is proportional to the rate of change of flux linkage.
	Lenz's law	Lenz's law states that the direction of the induced e.m.f. (and hence current flow in a closed circuit) is always such as to opposite the change in flux causing it.
CH 19: A.C.	Root-mean-square	The root-mean-square value of an alternating current (a.c.) is that value of the direct current (d.c.) that would produce thermal energy at the same rate in a resistor.

	Photoelectric effect	The photoelectric effect is the phenomenon of electrons being ejected from a given meal surface when light about a certain minimum frequency falls upon its surface.
Physics 1	Photon	A photon is an indivisible quantum (or packet) of electromagnetic energy which is emitted, transmitted and absorbed as a whole.
Quantum Physics	Electron-volt	The electron-volt is the amount of kinetic energy that an electron gains when it is accelerated through a potential difference of one volt.
Chapter 20:	Work function	The work function (ϕ) of a metal is the minimum mount of energy required for an electron to escape from the surface.
Ch	Stopping potential	The stopping potential is the minimum value of the retarding voltage which will just stop even the most energetic photoelectrons from reaching the collector, indicated by current just dropped to zero as measured in a micro-ammeter.
Phy 2	Emission line spectrum	The emission line spectrum consists of bright lines of definite wavelengths on a dark background.
Chapter 21: Q PI	Absorption line spectrum	The absorption line spectrum consists of a continuous spectrum crossed by dark lines due to some missing wavelengths.
	Heisenberg's Uncertainty Principle	Heisenberg's Uncertainty Principle states that the product of the uncertainty Δx in the position of a body at any instant and the uncertainty Δp in its momentum at the same instant is greater than or approximately equal to the Planck constant.
CH 22: N Phy 1	Isotope	Isotopes are atoms of the same element with the same number of protons but a different number of neutrons.
	Unified atomic mass	The unified atomic mass is equivalent to one-twelfth the mass of a carbon-12 atom.
	Relative atomic mass	The relative atomic mass is defined as the ratio of the mass of the atom to the mass of one-twelfth of the mass of the neutral carbon-12 isotope.
	Mass defect	The difference (Δm) between the mass of the nucleons and that of the nucleus is known as the mass defect.

Binding energy	The binding energy of the nucleus is defined as the energy released when the nucleus is first formed
	from its separate protons and neutrons.
	The binding energy of the nucleus is also defined as the work done on the nucleus to separate the
	nucleus into individual protons and neutrons.
Radioactivity	Radioactivity or radioactive decay is the spontaneous and random disintegration of heavy unstable
-	nuclei into more stable products through the emission of radiation such as alpha particles, beta
	particles and gamma rays.
lonising radiation	Ionising radiation refers to radiation that has enough energy to strip electrons from an atom, thus

cs 2	Radioactivity	Radioactivity or radioactive decay is the spontaneous and random disintegration of heavy unstable nuclei into more stable products through the emission of radiation such as alpha particles, beta particles and gamma rays.
Physics	lonising radiation	lonising radiation refers to radiation that has enough energy to strip electrons from an atom, thus producing free electrons and ions.
Nuclear	Half-life	Half-life is defined as the time taken for half the number of nuclei of a radioactive element to decay.
Chatper 23: Nu	Activity	The activity of a radioactive material is the number of disintegrations of its atoms per unit time.
	Decay constant	The decay constant is the probability that a radioactive nuclide of an isotope in a sample would decay in one second.
	Decay Law	The rate of radioactive decay $A = -\frac{dN}{dt}$ is directly proportional to the number N of the radioactive nuclei present.

ANNEX 4: SUMMARY OF KEY QUANTITIES, SYMBOLS AND UNITS

The following list illustrates the symbols and units that will be used in question papers.

Quantity	Usual symbols	Usual unit
Base Quantities		
mass	т	kg
length	1	m
time	t	S
electric current	Ι	А
thermodynamic temperature	Т	K
amount of substance	n	mol
Other Quantities		
distance	d	m
displacement	s, x	m
area	A	m ² ₃
volume	V, v	m^3
density	ρ	$kg m^{-3}$
speed	U, V, W, C	$m s^{-1}$
velocity	U, V, W, C	m s ⁻¹ m s ⁻²
acceleration acceleration of free fall	a	m s ⁻²
force	g F	N
weight	F W	N
momentum	p	Ns
work	w, W	J
energy	E,U,W	J
potential energy	<i>Ε</i> , <i>σ</i> , <i>π</i>	J
kinetic energy	E _k	J
heating	Q	J
change of internal energy	ΔU	J
power	Р	W
pressure	p	Pa
torque	T	Nm
gravitational constant	G	$N kg^{-2} m^2$
gravitational field strength	g	N kg ⁻¹ J kg ⁻¹
gravitational potential	ϕ	
angle	heta	°, rad
angular displacement	heta	°, rad
angular speed	Ø	rad s ⁻¹
angular velocity	Ø	rad s ⁻¹
period	T	S
frequency	f	Hz
angular frequency	ω	rad s ⁻¹
wavelength	λ	m _1
speed of electromagnetic waves	С	m s ^{−1}
electric charge	Q	С
elementary charge	e	C
electric potential	V	V V
electric potential difference electromotive force	V	V V
resistance	E R	
resistivity		Ω
•	ρ Ε	Ωm NC ⁻¹ , Vm ⁻¹
electric field strength permittivity of free space		Fm^{-1}
magnetic flux	E0 A	Wb
magnetic flux density	Φ B	T
permeability of free space		и Н m ⁻¹
force constant	μ_0 k	$N m^{-1}$
	Λ	INTI

Quantity	Usual symbols	Usual unit
Celsius temperature	θ	°C
specific heat capacity	С	$J K^{-1} kg^{-1}$
molar gas constant	R	$J K^{-1} mol^{-1}$
Boltzmann constant	k	J K ^{−1}
Avogadro constant	N _A	mol ^{−1}
number	N, n, m	_
number density (number per unit volume)	n	m^{-3}
Planck constant	h	Js
work function energy	Φ	J
activity of radioactive source	A	Bq s ^{−1}
decay constant	λ	s ⁻¹
half-life	t _{1/2}	S
relative atomic mass	A _r	
relative molecular mass	<i>M</i> _r	
atomic mass	m _a	kg, u
electron mass	m _e	kg, u
neutron mass	m _n	kg, u
proton mass	mp	kg, u
molar mass	Μ	kg
proton number	Z	
nucleon number	A	
neutron number	Ν	

ANNEX 5: DATA AND FORMULAE

Data			
speed of light in free space	с	=	$3.00 \times 10^8 \mathrm{m \ s^{-1}}$
permeability of free space	μ_0	=	$4\pi \times 10^{-7} \text{ H m}^{-1}$
permittivity of free space	E0	=	$8.85 \times 10^{-12} F m^{-1}$
			$(1/(36\pi)) imes 10^{-9} { m F} { m m}^{-1}$
elementary charge	е	=	$1.60 \times 10^{-19} \mathrm{C}$
the Planck constant	h	=	$6.63 \times 10^{-34} \text{Js}$
unified atomic mass constant	и	=	$1.66 \times 10^{-27} \text{kg}$
rest mass of electron	m _e	=	$9.11 \times 10^{-31} \text{kg}$
rest mass of proton	$m_{ m p}$	=	$1.67 \times 10^{-27} \text{kg}$
molar gas constant	R	=	8.31 J K ⁻¹ mol ⁻¹
the Avogadro constant	N _A	=	$6.02 \times 10^{23} mol^{-1}$
the Boltzmann constant	k	=	$1.38 \times 10^{-23} \text{J K}^{-1}$
gravitational constant	G	=	$6.67\times 10^{-11}Nm^2kg^{-2}$
acceleration of free fall	g	=	$9.81\mathrm{ms}^{-2}$
Formulae			
uniformly accelerated motion	s	=	$ut + \frac{1}{2}at^{2}$
			$u^{2} + 2as$
work done on/by a gas	W	=	pΔV
hydrostatic pressure	p	=	hogh
gravitational potential	ϕ	=	-Gm/r
temperature	T/K	=	T/°C + 273.15
pressure of an ideal gas	p	=	$\frac{1}{3}\frac{Nm}{V} < c^2 >$
mean translational kinetic energy of an ideal gas molecule	E	=	$\frac{3}{2}kT$
displacement of particle in s.h.m.	x	=	z x₀sin∞t
velocity of particle in s.h.m.	V	=	v₀cos∞t
		=	$\pm \omega \sqrt{\left(\mathbf{x}_{o}^{2}-\mathbf{x}^{2}\right)}$
electric current	Ι	=	Anvq
resistors in series	R	=	$R_1 + R_2 + \dots$
resistors in parallel	1/ <i>R</i>	=	$1/R_1 + 1/R_2 + \dots$
electric potential	V	_	Q
	V	=	$\frac{Q}{4\pi\varepsilon_0 r}$
alternating current/voltage	X	=	x _o sin∞t
magnetic flux density due to a long straight wire	В	=	$rac{\mu_0 I}{2\pi d}$
magnetic flux density due to a flat circular coil	В	=	$\frac{\mu_0 NI}{2r}$

magnetic flux density due to a long solenoid radioactive decay

decay constant

$$B = \mu_0 nI$$

$$x = x_0 \exp(-\lambda t)$$

$$\lambda = \frac{\ln 2}{t_{\frac{1}{2}}}$$